



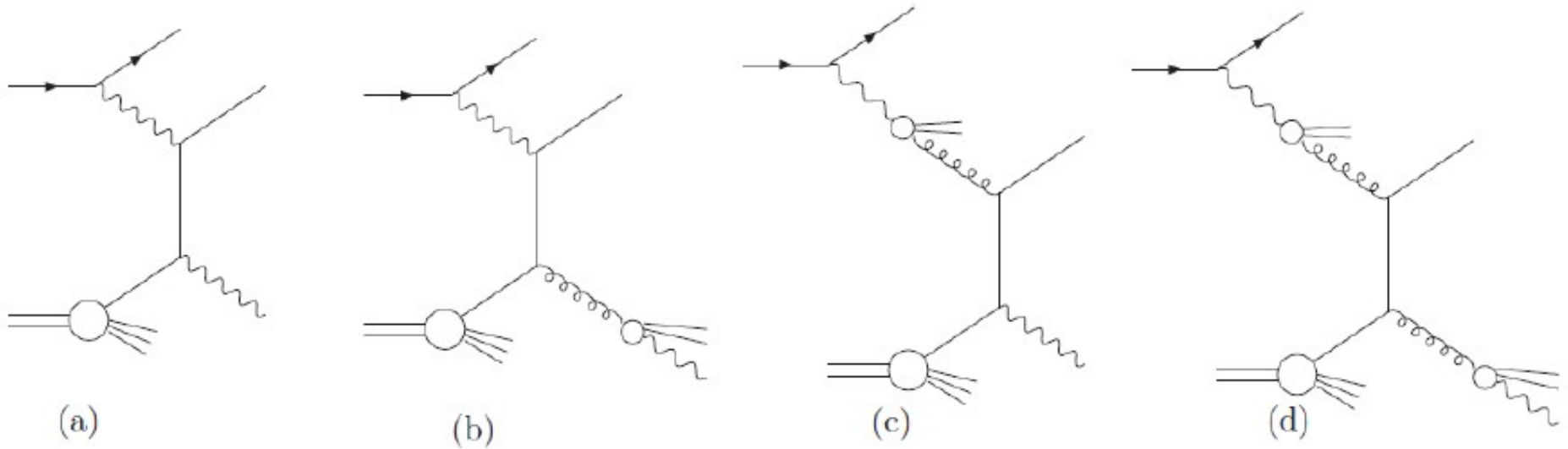
# Prompt Photon in PHP Analysis

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# Introduction



A prompt photon is one that emerges directly from a perturbative QCD process. LO diagrams are illustrated above:

- (a) direct, in which the entire incoming photon interacts,
- (c) resolved, in which a parton from the photon interacts.

Higher order pQCD processes occur and also “fragmentation” processes (b, d).

## **Motivation**

- Prompt (isolated, high pT) photons are a useful tool to study and test QCD.
- Their measurements are more precise than hadronic jets.
- Prompt photons can be used to measure and constrain the pdfs of proton and photon.

## **Differences from previous experiments**

- Luminosity  $370 \text{ pb}^{-1}$ . First ZEUS prompt photons used  $6.4 \text{ pb}^{-1}$ , latest published –  $77 \text{ pb}^{-1}$ .
- For jet building and photon finding ZUFO – ZEUS Unidentified Flying Object – are used, signal identification based on shower shapes. Previous analyses used calorimeter cells, BPRE signal.
- Phase space. Different sets of  $E_T$  and cuts  $\eta$  are used.
- For jet finding KTCLUS and Fastjet are used.
- More complex Monte-Carlo background sample (previously single-particle MC was used).

## Theory

**FGH** – LO and NLO and the box diagram term are calculated explicitly. Fragmentation processes are calculated in terms of a fragmentation function in which quark or gluon jet gives rise to a photon.

**LMZ** – The  $k_T$  factorisation method makes use of unintegrated proton parton densities at NLO + box following the approach of KMR (Kimber-Martin-Ryskin). The procedure gives a quark-radiated contribution that is enhanced relative to the LO collinear approximations. The most recent values with larger errors than before are used. LMZ has no photon structure function in their model for resolved photons.

# Analysis procedure

- Apply cleaning photoproduction cuts.
- Use Zufos as photon candidates.
- Find accompanying hadronic jet.
- Subtract the background using  $\langle\delta Z\rangle$  quantity – it is broader for background than for signal. Perform a statistical subtraction. It is done for every bin of every measured variable.
- Calculate acceptance corrections, cross-sections.

# Data Samples

•**Data:** HERA II 04p, 04/05e, 06e, 06p, 07p (Common Ntuples v06d) 370 pb<sup>-1</sup>

•**MC Signal:** 04p, 05e, 06e, 06p, 07p (CN v06b PYTHIA v6.416, HERWIG v 6.510 – used for systematics)  
Direct, Resolved

•**MC Background:** 04p, 04/05e, 06e, 06p, 07p (CN v06b PYTHIA - Heavy Flavour Group, Jet – Sebastian's + Filtered, HERWIG – Filtered – used for systematics) Direct, Resolved

## Cuts

### •Event Selection

- Trigger HPP16 on
- $0.2 < y_{JB} < 0.7$
- $|Z_{vtx}| < 40$  cm
- $|BCAL\ time| < 10$  ns
- Cal  $p_T < 10$  GeV
- No SINISTRA electron with Prob > 0.9 and Yel < 0.7

### •Prompt Photon Selection

- Tuf0[0] = 31
- $-0.7 < \eta^{z\text{ufo}} < 0.9$
- $6 < E_T^{z\text{ufo}} < 15$  GeV
- $E^{z\text{ufo}}/E^{\text{jet}} > 0.9$
- ZufoEmc/ZufoEcal > 0.9
- track isolation in cone 0.2

### ● Jet Selection

- $-1.5 < \eta^{\text{jet}} < 1.8$
- $4 < E_T^{\text{jet}} < 35$  GeV

### ● Truth level selection

- $Q^2 < 1$  GeV<sup>2</sup>
- $0.2 < y_{JB} < 0.7$
- Particle type 29
- $-0.7 < \eta^{\text{particle}} < 0.9$
- $6 < E_T^{\text{particle}} < 15$  GeV
- $E^{\text{particle}}/E^{\text{jet}} > 0.9$

# Phase space of previous measurements

## H1 (DESY 09-135)

- $Q^2 < 1 \text{ GeV}^2$
- $0.1 < y_{\text{JB}} < 0.7$
- $-1.0 < \eta^{\text{zifo}} < 2.4$
- $6 < E_{\text{T}}^{\text{zifo}} < 15 \text{ GeV}$
- $-1.3 < \eta^{\text{jet}} < 2.3$
- $E_{\text{T}}^{\text{jet}} > 4 \text{ GeV}$

## ZEUS (DESY 06-125)

- $Q^2 < 1 \text{ GeV}^2$
- $0.2 < y_{\text{JB}} < 0.8$
- $-0.74 < \eta^{\text{zifo}} < 1.1$
- $5 < E_{\text{T}}^{\text{zifo}} < 16 \text{ GeV}$
- $-1.6 < \eta^{\text{jet}} < 2.4$
- $6 < E_{\text{T}}^{\text{jet}} < 17 \text{ GeV}$

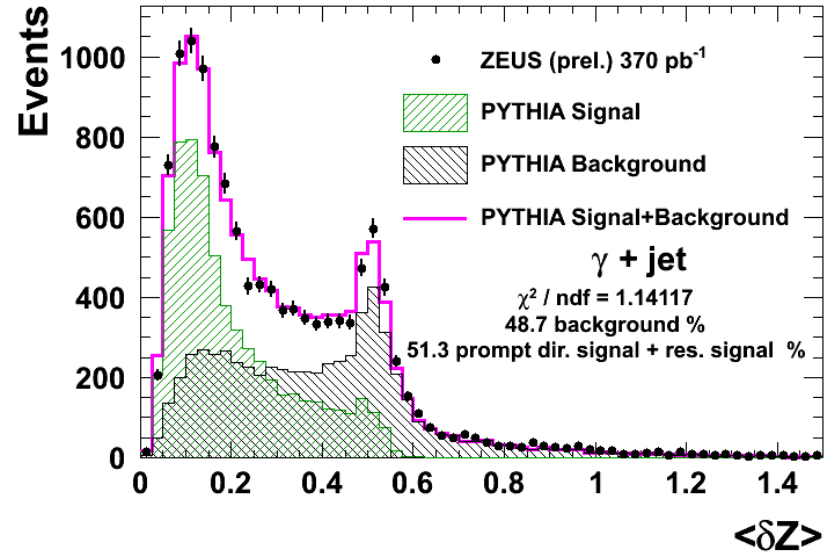
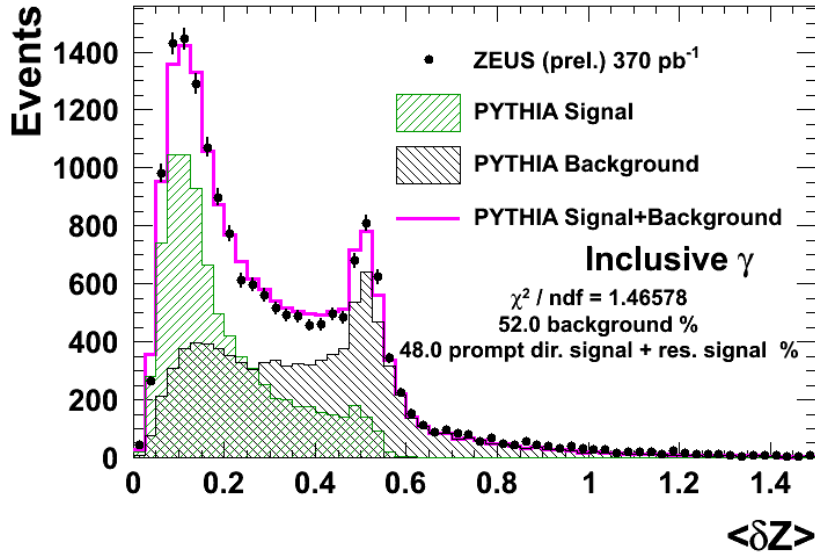
## ZEUS (DESY 99-161)

- $0.2 < y_{\text{JB}} < 0.9$
- $-0.7 < \eta^{\text{zifo}} < 0.9$
- $5 < E_{\text{T}}^{\text{zifo}} < 10 \text{ GeV}$

# Signal extraction

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$\langle \delta Z \rangle$  energy weighted mean width of the electromagnetic cluster in Z direction:

$$\langle \delta Z \rangle = \frac{\sum_i E_i |Z_i - Z_{\text{cluster}}|}{(w_{\text{cell}} \sum_i E_i)}$$

Distributions of  $\langle \delta Z \rangle$  for (a) inclusive photon events, (b) events with a jet, showing the fitted signal and background components.

Chi\*\*2 per degree of freedom in bins of examined variables is typically 1.1.

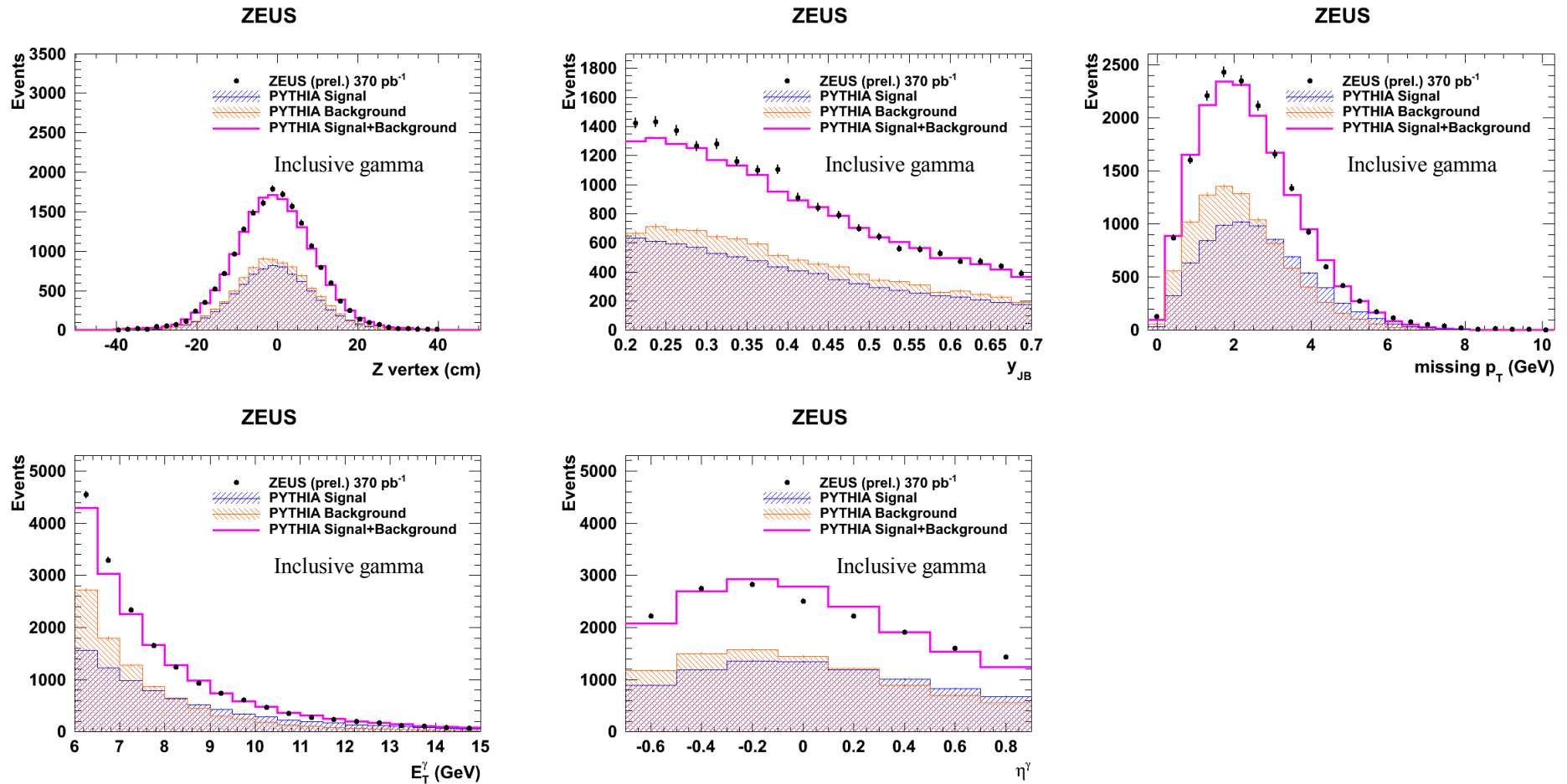


A model of  
50% PYTHIA Direct,  
40% Resolved,  
5% each of resolved and direct radiative (“fragmentation”).

The radiative events are obtained from the background.  
They are discarded from the background before it is  
used in the fits.

A systematic uncertainty allows for uncertainties in this model.

# Control plots. Inclusive photon

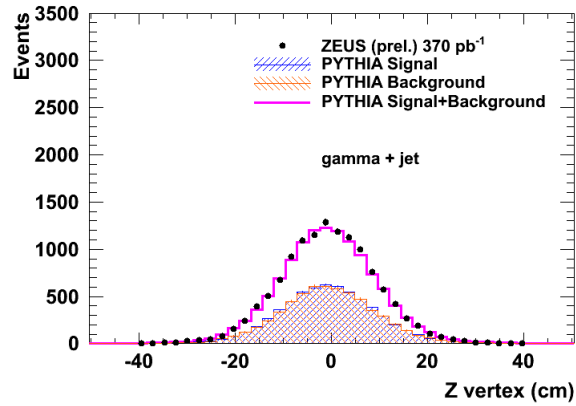


Shown are results of the deltaz fit to signal+background for Z vertex,  $y_{JB}$ , missing transverse momentum,  $E_T^\gamma$ ,  $\eta^\gamma$ .

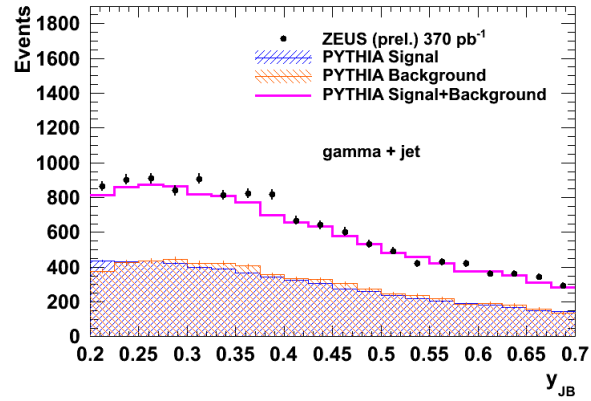
PYTHIA MC signal consists of 50% direct-prompt, 5% direct-fragmentation, 40% resolved-prompt, 5% resolved-fragmentation.

# Control plots. Photon + jet

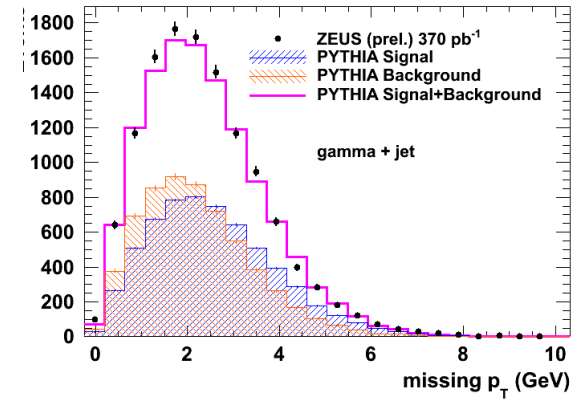
ZEUS



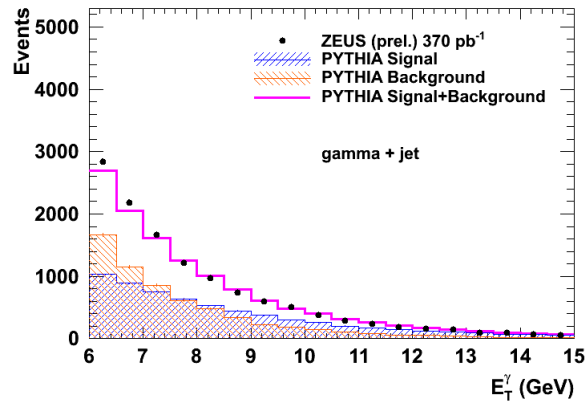
ZEUS



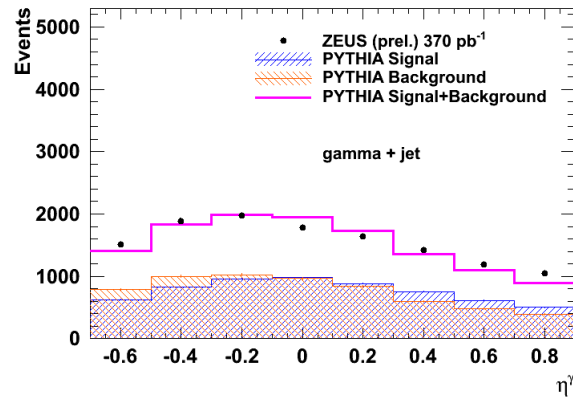
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ZEUS



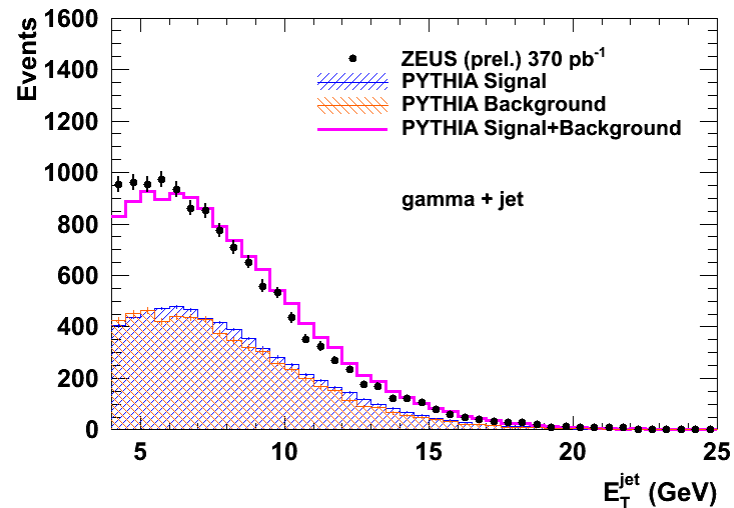
ZEUS



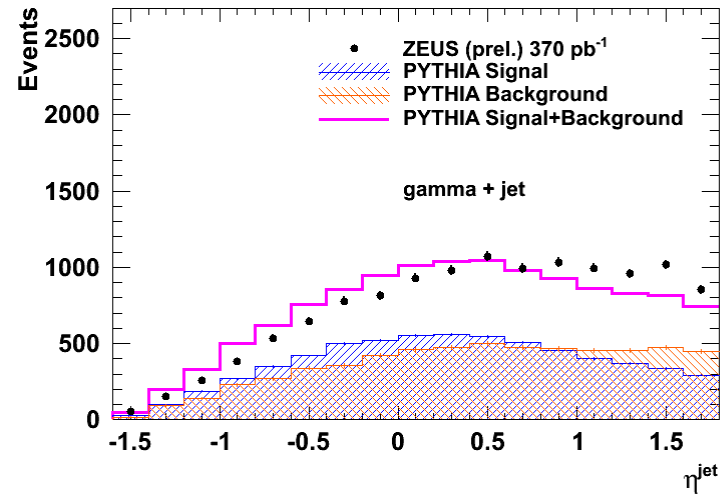
Description of data looks reasonable.

# Control plots. Photon + jet

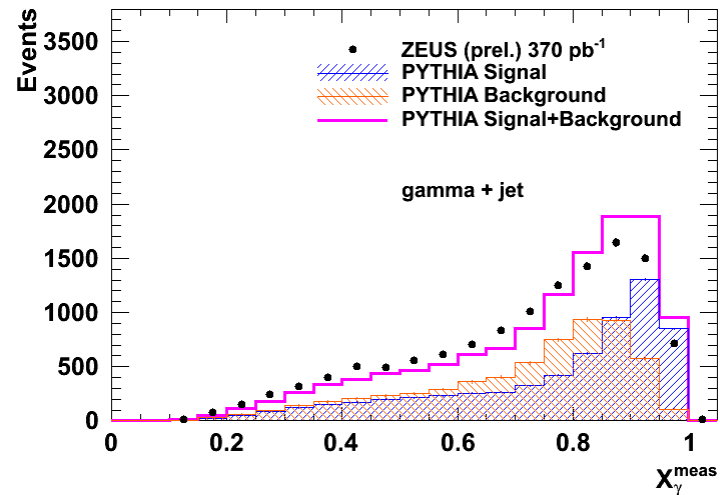
ZEUS



ZEUS



ZEUS



In view of the minor mismatch in the modelling of  $E_T^{\text{jet}}$ , a reweighting procedure was studied.

It made little difference to the results and is not used.

$X_\gamma^{\text{MEAS}}$  – the fraction of the incoming photon energy given to the final state photon and jet, at a lowest-order approximation

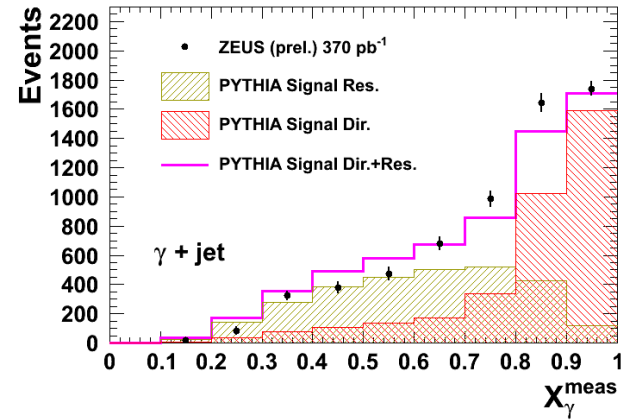
$$x_\gamma^{\text{meas}} = \frac{E^\gamma + E^{\text{jet}} - p_Z^\gamma - p_Z^{\text{jet}}}{E^{\text{all}} - p_Z^{\text{all}}}$$

# Definition of direct/resolved mix

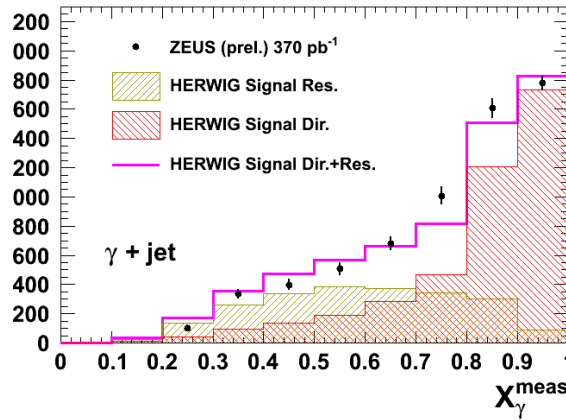
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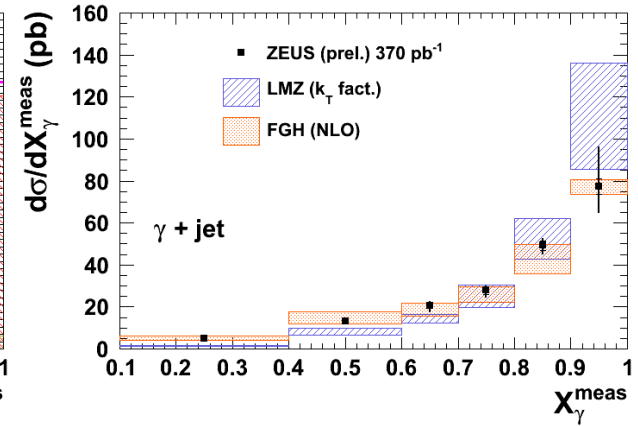
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a)



b)



c)

- Events detected for different values of  $X_\gamma^{\text{MEAS}}$ , compared to a mixture of 50% direct prompt, 40% resolved prompt, 5% dir frag and 5% res frag events generated using PYTHIA. No acceptance corrections have been applied at this point.
- b) HERWIG events. Mix is 60% dir prompt, 30% res prompt, 5% dir frag and 5% res frag.
- Cross sections as a function of  $X_\gamma^{\text{MEAS}}$ , for events containing an inclusive photon and a jet, compared to predictions from FGH (uncertainty from varying all scales simultaneously, changing only one scale gives uncertainties  $\pm 10\%$ ) and LMZ.

Here and in following hadronization corrections are applied to theory.

Both PYTHIA and HERWIG are not perfect. Using PYTHIA as central value.

# Trigger studies

# Trigger studies

Investigations of FLT and SLT.

PJB.

9 Aug 2013

1) FLT studies

2) SLT studies

# Trigger studies

## Procedure:

Following Katsuo's advice, take DIS events which contain high-ET photons (our DIS sample, thank you Oleg).

Select events with an inclusive isolated prompt photon satisfying broadened cuts in ET, but otherwise as in the analysis.

No jet requirement. Normal trigger requirement is not applied.

In the DIS, select events having an FLT that just takes RCAL information.

See what fraction of selected events fire the relevant photon FLT.

1) The DIS FLT triggers selected were 30, 36 and 46.

2) The HPP FLT uses triggers 28, 30, 39. 30 is a DIS trigger, and so look for 28 or 39. 28 seems included in 39 **so just look for 39**

Compare the DIS data to the php direct MC.

**We first look at track multiplicity and add a track multiplicity requirement to the MC events.**



# Trigger studies

**DIS** Plot the DIS flt39 efficiency as a function of number of tracks, i.e. **Trk\_ntracks**

Black=all selected, red = selected+flt39

Ratio red/black = flt eff.

Events

FLT eff.

Events

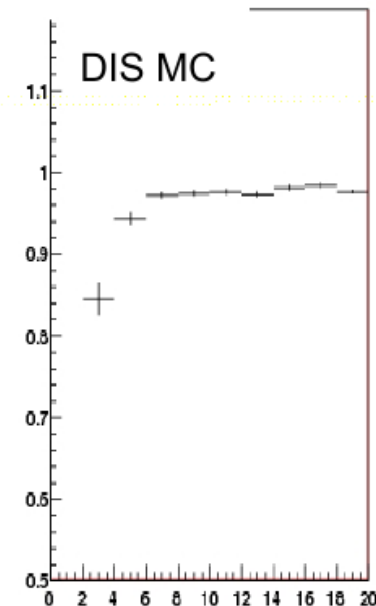
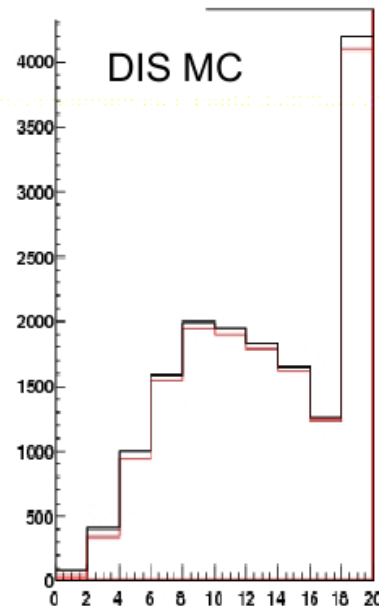
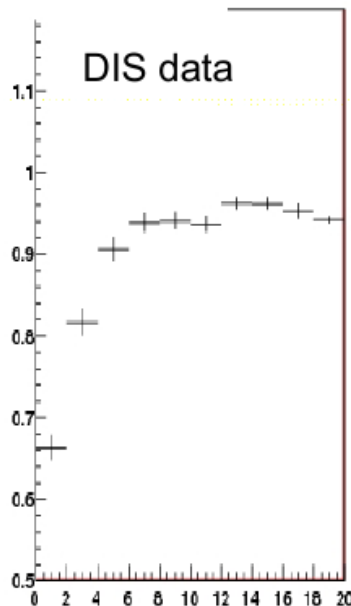
FLT eff.

**FLT-NTR0**

**FLT-NTR-rat**

**FLT-NTR0**

**FLT-NTR-rat**



NTR

Conclusion: There is a low-multiplicity component to the DIS data (DVCS??) which has very different properties from the rest. Otherwise the two distributions are similar but the MC at 98% is a bit higher than the data at 95%.

# Trigger studies

**Php** plot the php flt39 efficiency as a function of number of tracks, i.e. **Trk\_ntracks**

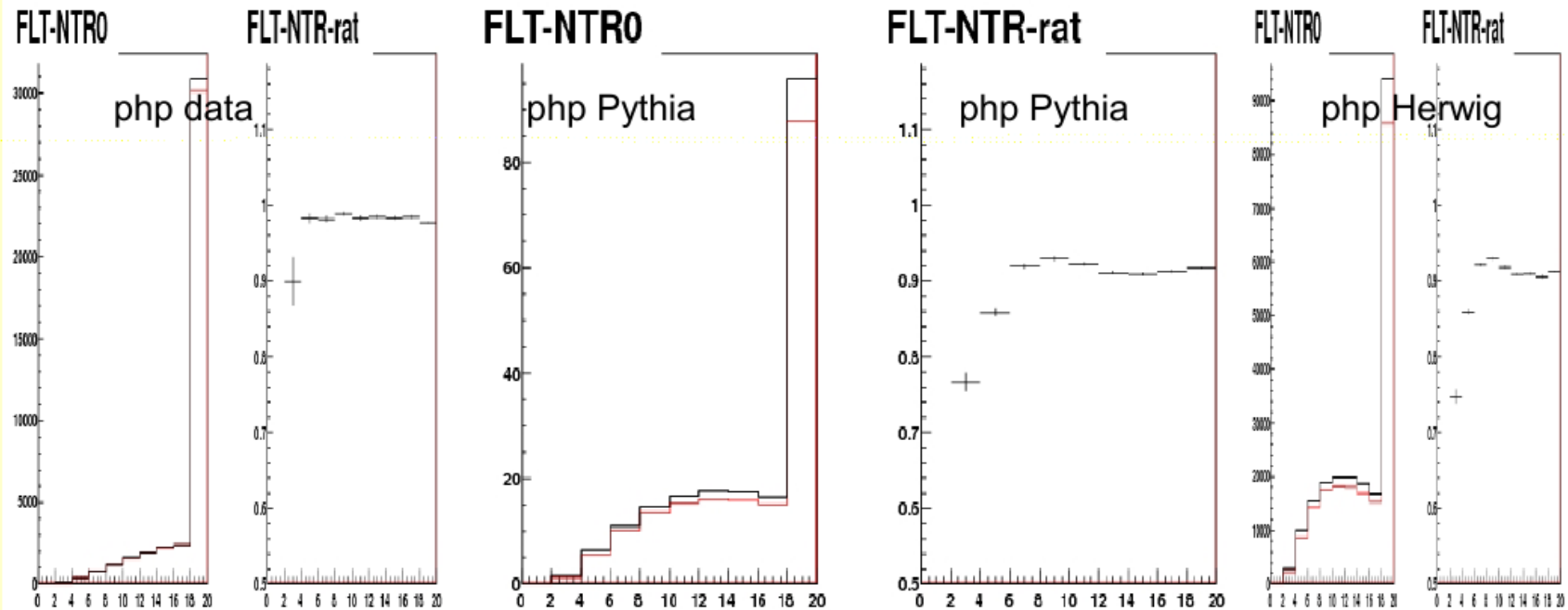
Black=all tagged, red = +flt39      Event numbers,      ratio red/black = flt eff.

Events

Events

FLT eff

FLT Eff

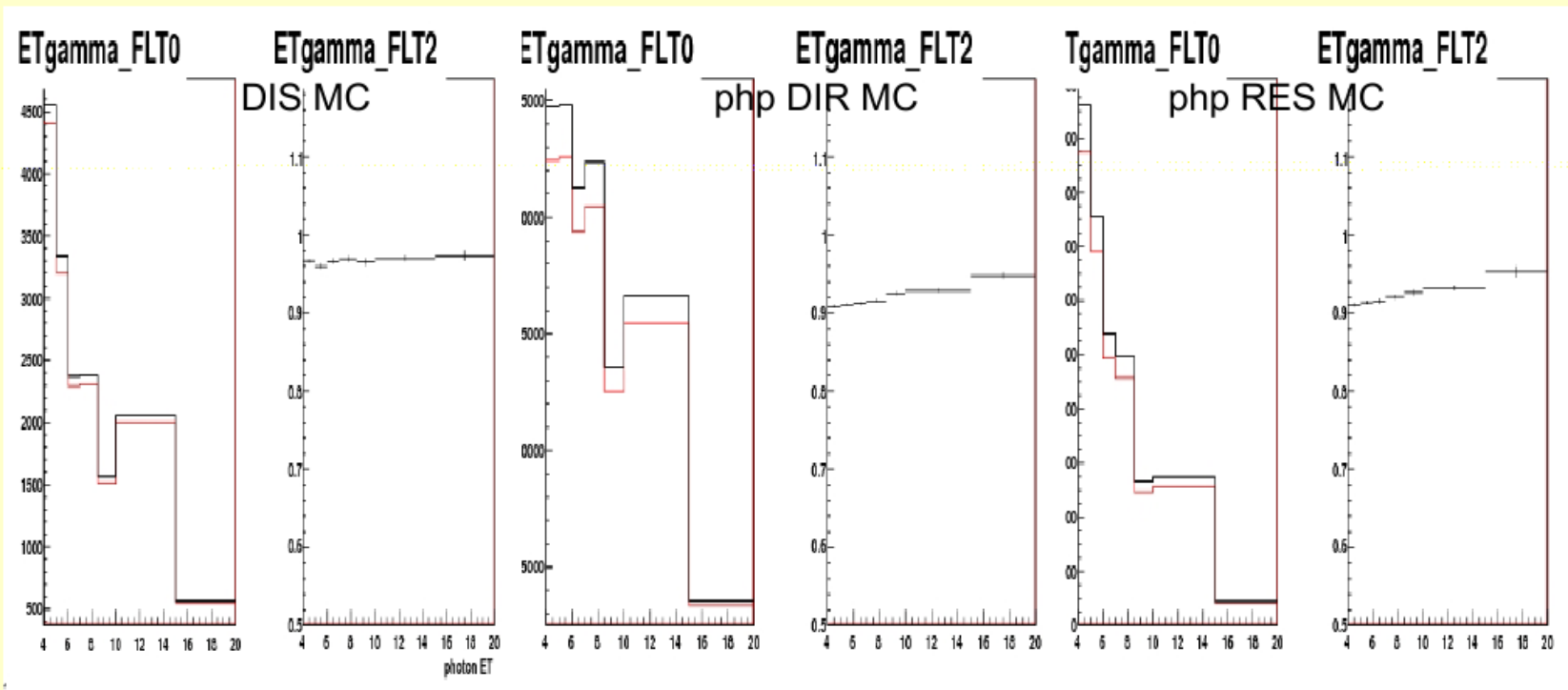


Conclusion: The php data have high track multiplicity and so we can ignore the low-multiplicity effect. Enough statistics in both MCs to show php is somewhat similar to DIS. The MC are a mixture of resolved and direct.

# Trigger studies

New: now compare the different **MCs** imposing the condition that  $\text{Trk\_ntracks} \geq 5$

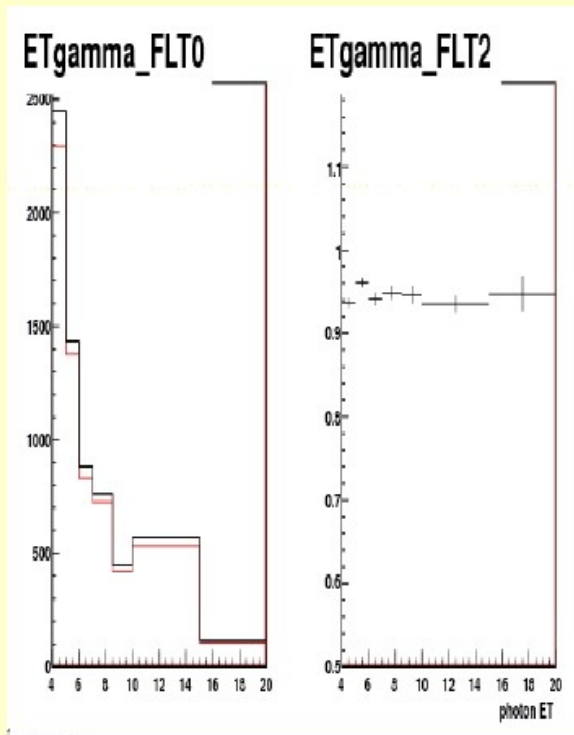
**flt39 efficiency as before. Plot as function of photon ET.**



Conclusion: the distributions are fairly flat and DIS is different from photoproduction.

# Trigger studies

Compare with **DIS data** with ntracks ge 5:



Between 6 and 15 GeV we have the following flt39 efficiencies:

DIS data: 94-95%

DIS MC: 97%

php MC: 91-93%

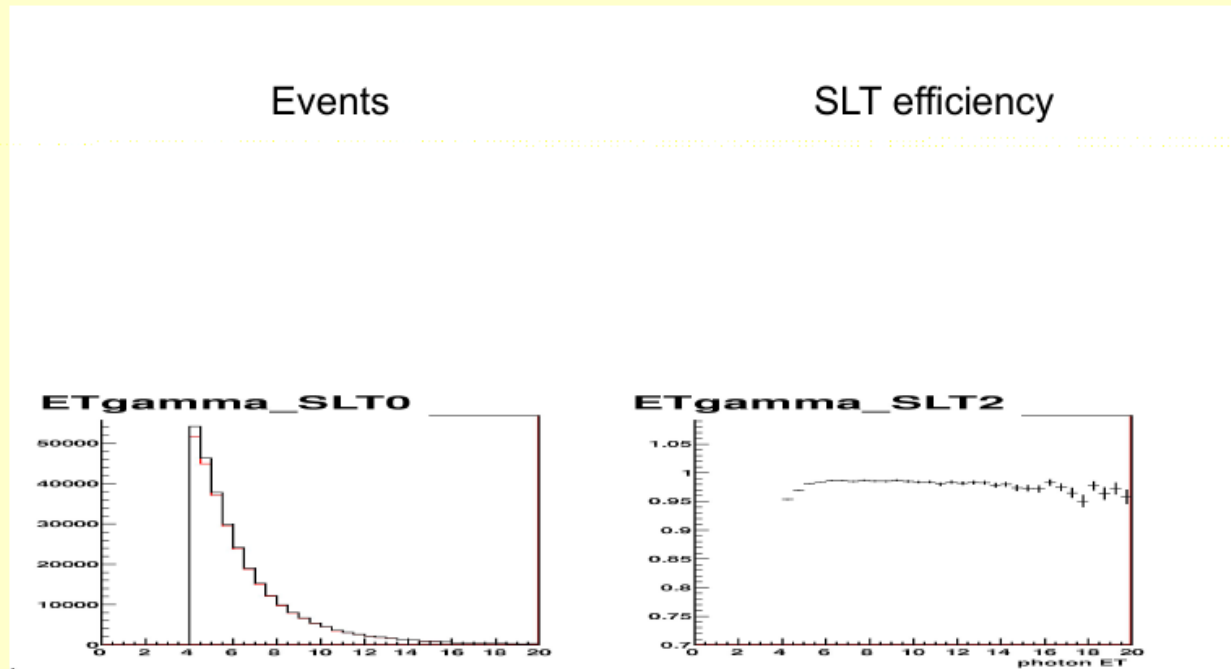
Advice from Katsuo:  
the present php MC efficiency should be corrected down by 2-3% to match the effect seen in the DIS samples.

**Selecting between 6-15 GeV  
the correction is 2.64%**

# Trigger studies

## Study of SLT efficiencies

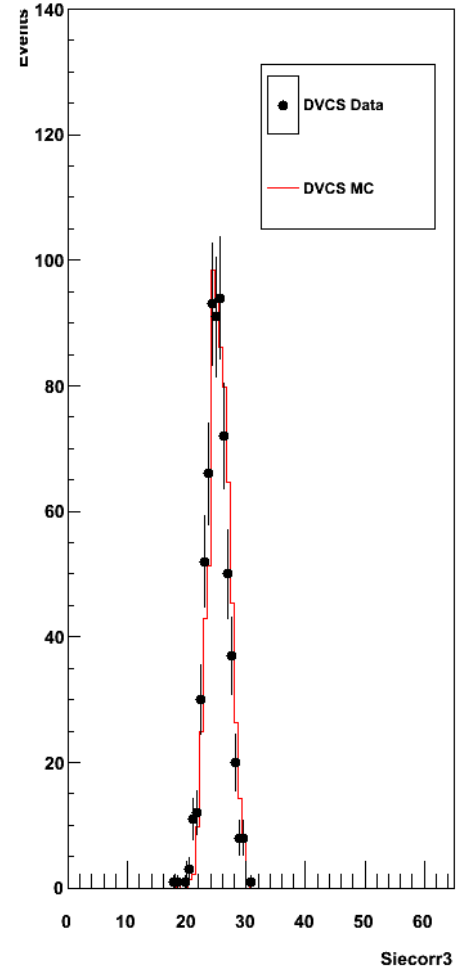
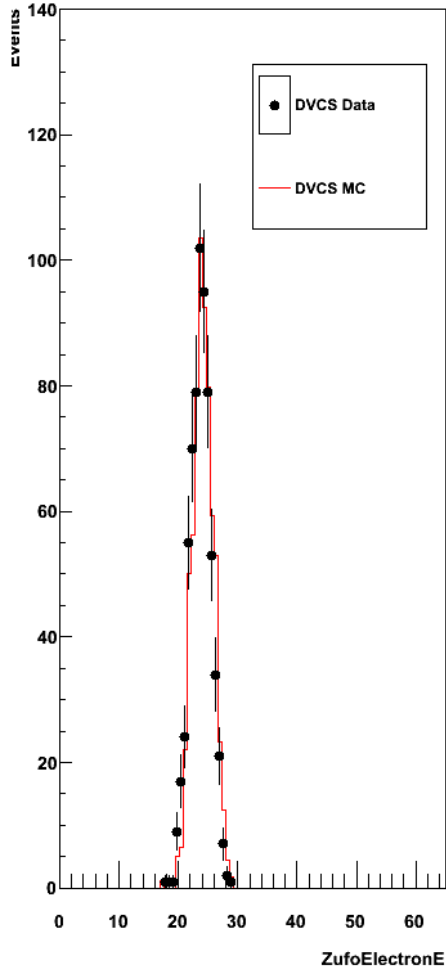
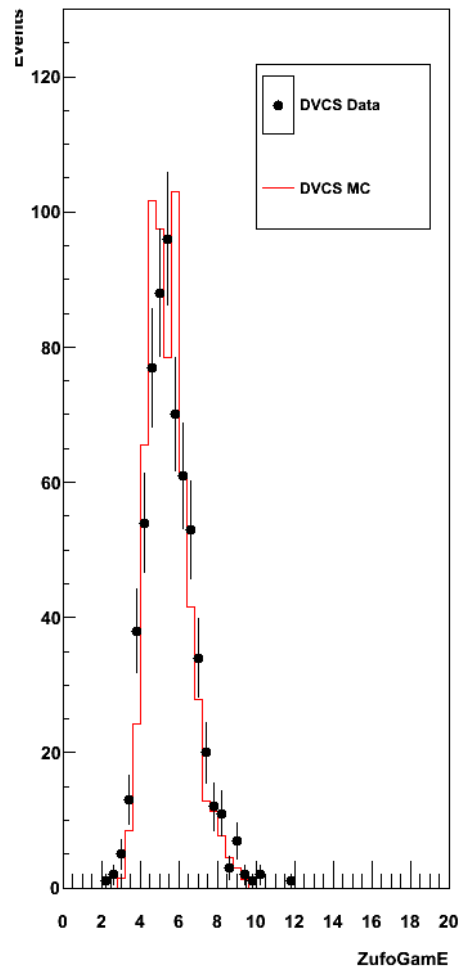
Take php MC events that pass the FLT trigger, and ask what fraction also pass the SLT trigger. Inclusive photons are selected with normal event requirements apart from the ET and trigger.



Conclusion: The SLT efficiency is rather flat at  $\sim 98\%$  within our normal accepted range. Above 8 GeV the “cone requirement” efficiency should be 100%. So the inefficiency seen must come from event cleaning requirements. We are well clear of the fall off at lower ET values, though the efficiency is quite high everywhere.

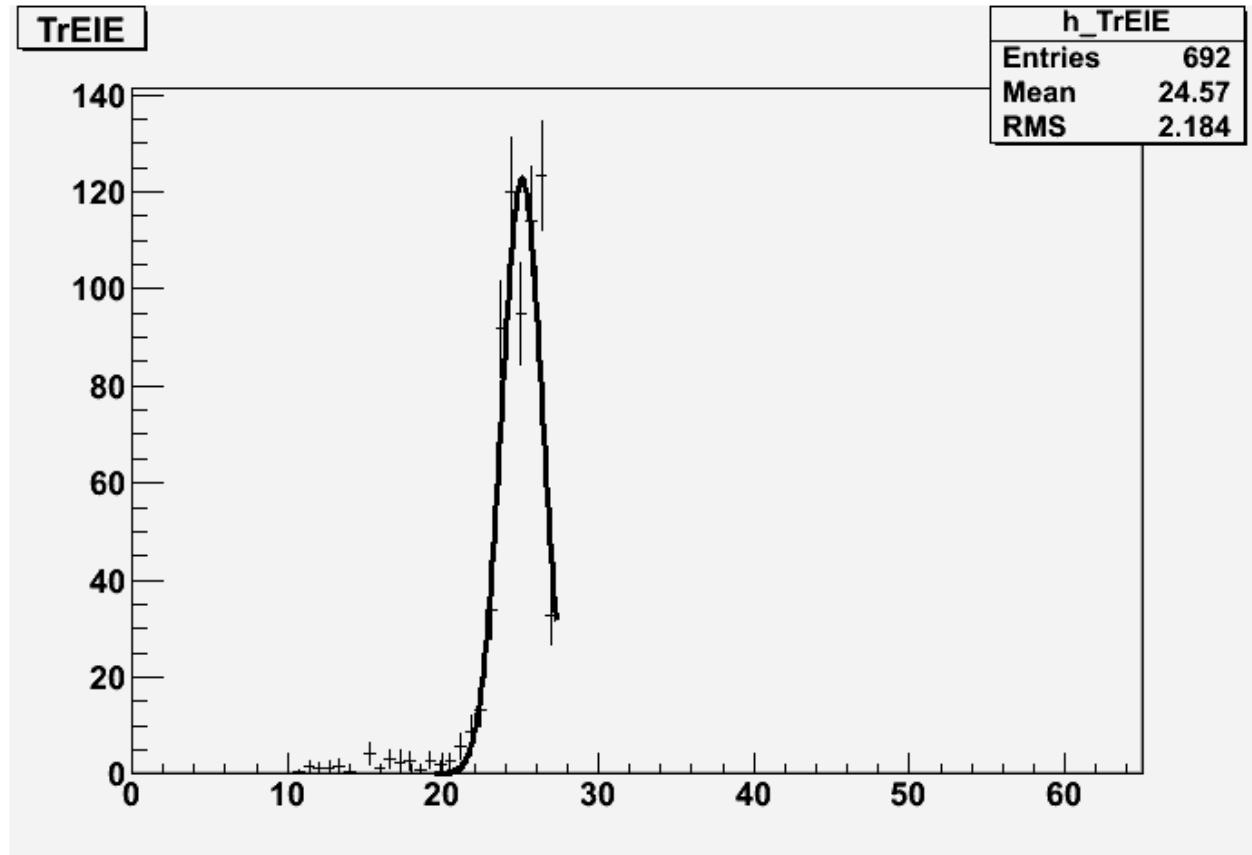
# Energy scale studies

# DVCS . Control plots after reweighting



Electron peaks agree within 2% for both Zupo and Sira electrons.

# DVCS Hadron level electron E

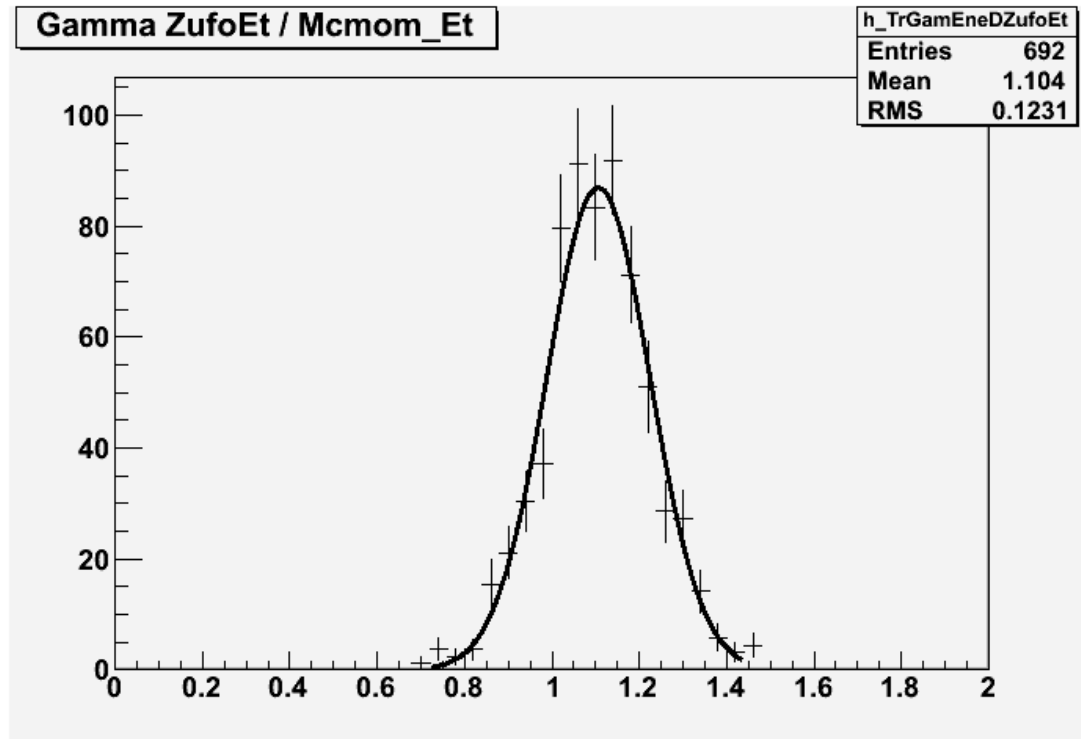


Hadron level electron energy

EXT	PARAMETER	STEP	FIRST	
NO.	NAME	VALUE	ERROR	SIZE DERIVATIVE
1	Constant	1.22490e+02	6.70833e+00	1.65887e-02 1.84283e-06
2	Mean	<b>2.50732e+01</b>	6.41946e-02	1.94258e-04 -8.07657e-04 - within 2% of S inistra in Data and detector level MC
3	Sigma	1.32999e+00	5.58002e-02	3.35839e-05 8.78500e-03



# DVCS MC Gamma Det Et / Had Et



MC EtGamDet / EtGamHad

EXT PARAMETER

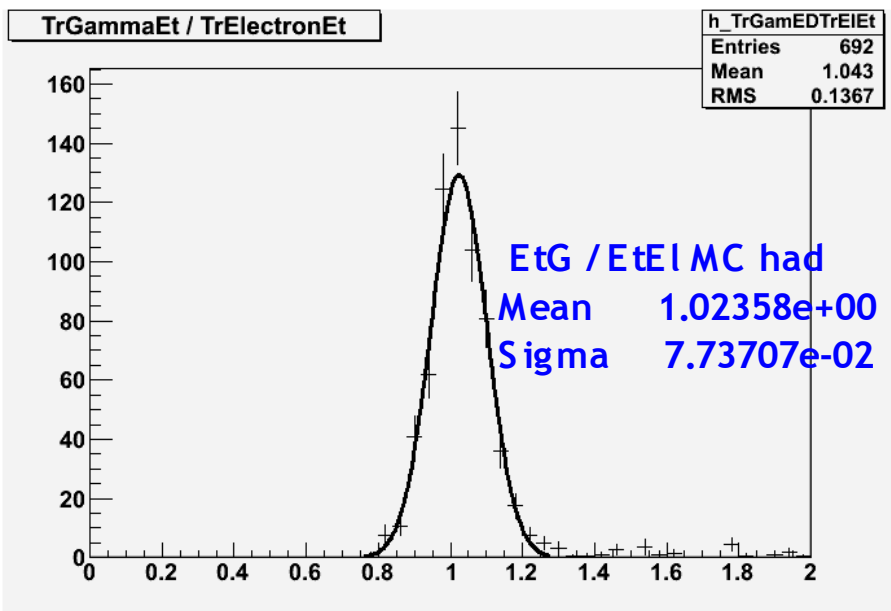
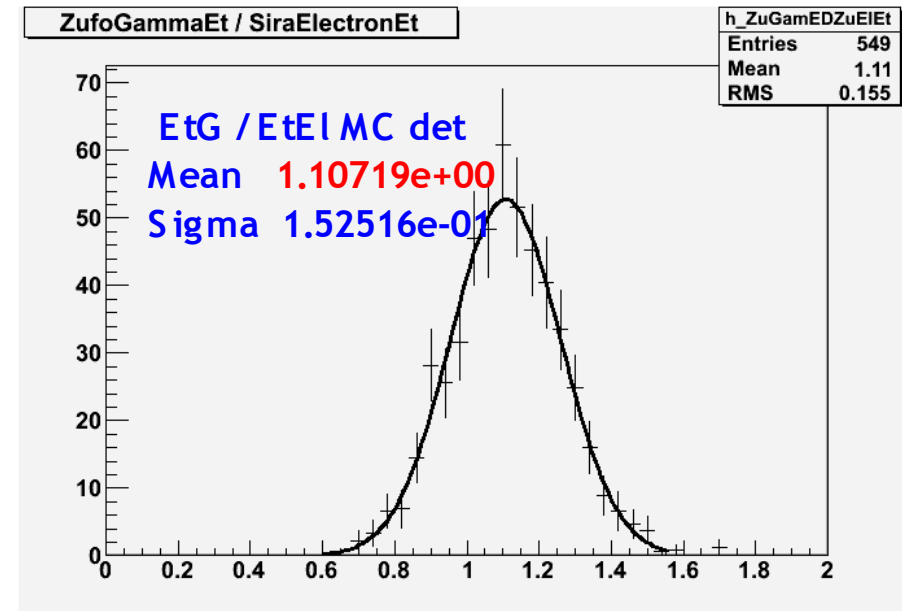
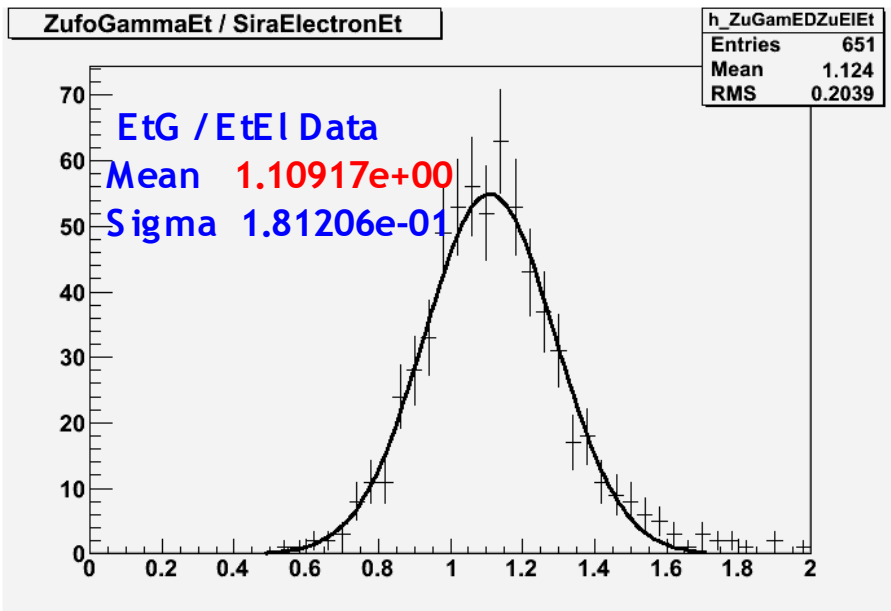
STEP FIRST

NO.	NAME	VALUE	ERROR	SIZE	DERIVATIVE
1	Constant	8.69706e+01	4.58659e+00	7.14497e-03	7.45215e-08
2	Mean	<b>1.10550e+00</b>	4.83577e-03	9.73969e-06	3.26266e-05
3	Sigma	1.18495e-01	3.96685e-03	1.72827e-05	-2.52587e-05

Reweighting applied.

A 10% effect similar to PHP.

# DVCS Zufo Gamma Et / Ele Et

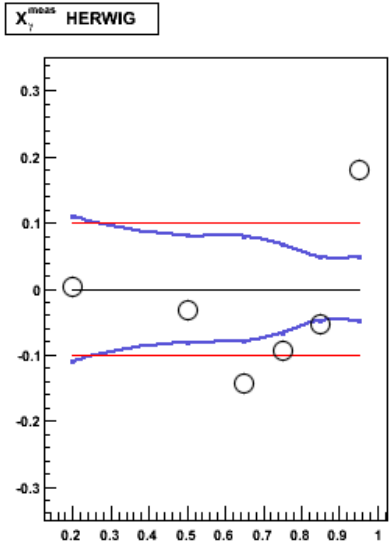
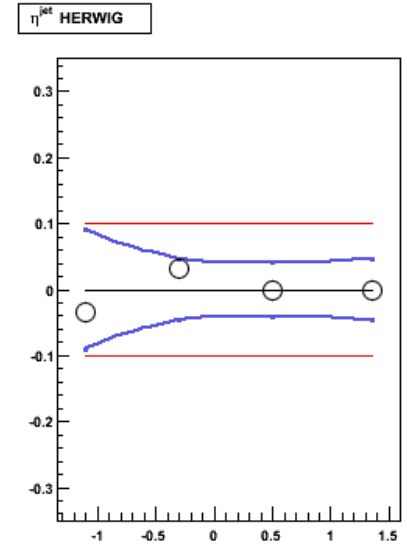
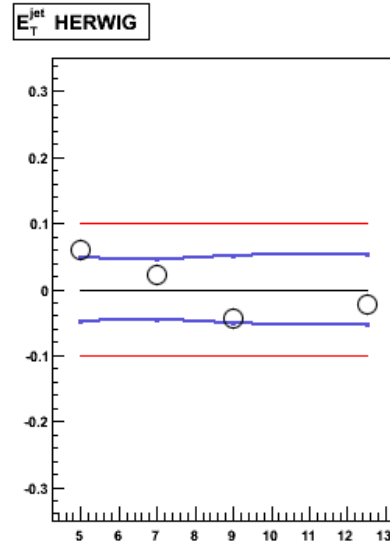
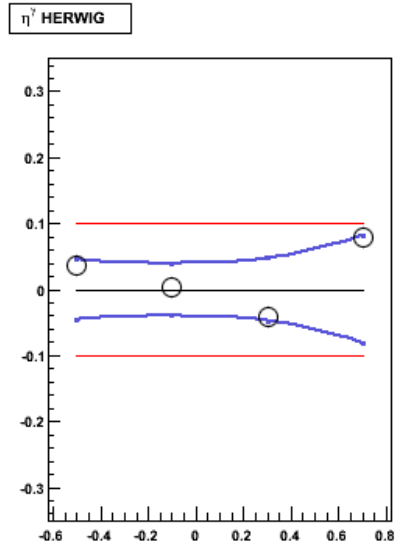
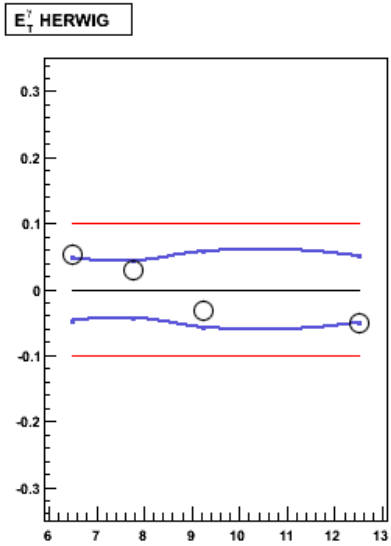


Reweighting applied.

Gamma and electron Et difference in Data is close to MC detector level value.

Factor 1.1 is applicable to MC and Data.

# Systematic uncertainties: HERWIG

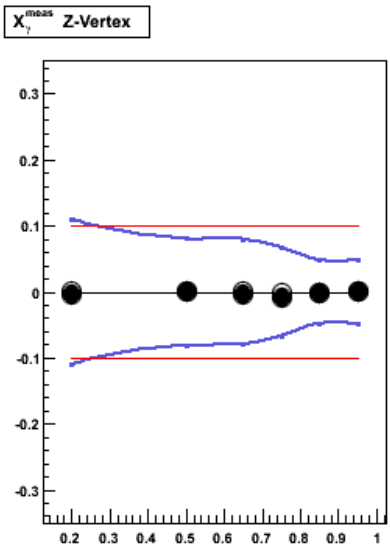
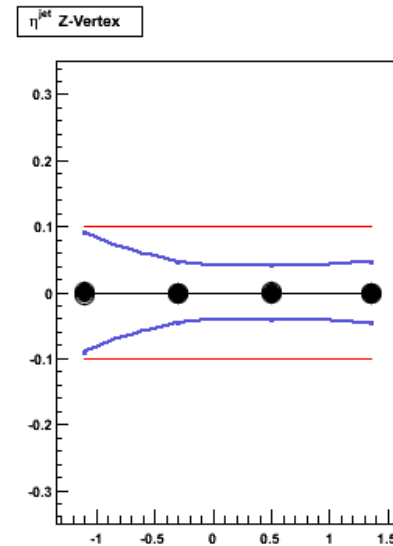
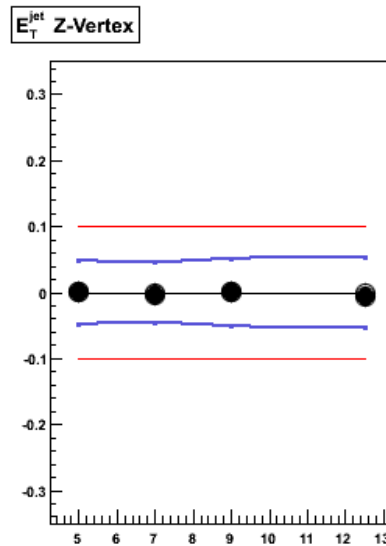
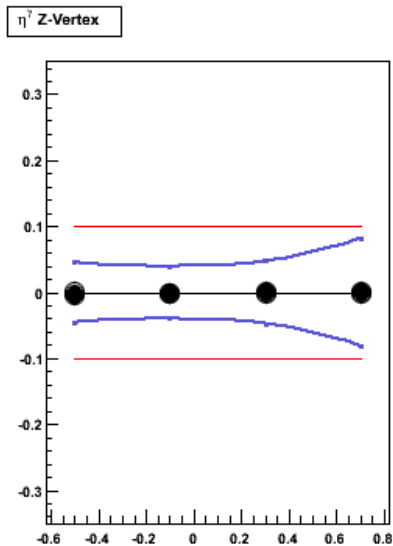
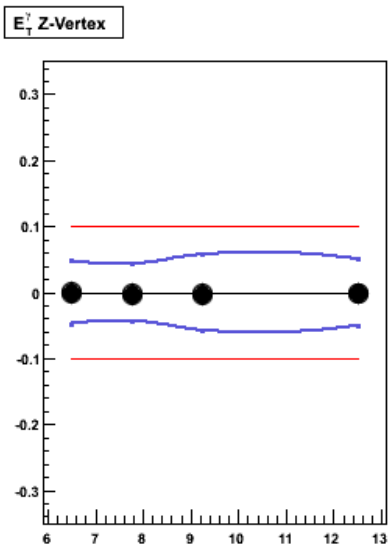


Use HERWIG signal and background instead of PYTHIA

$\gamma$ +jet

- *Rel. statistical uncertainties*
- *10% line*
- *HERWIG*

# Systematic uncertainties: Zvtx



**Standard cut:**

- $|Z_{vtx}| < 40$  cm

**Vary z-vertex cut by  $\pm 5$  cm**

From now on systematics are averaged in every bin independently, without the shift of the central value.

**$\gamma$ +jet**

— Rel. statistical uncertainties

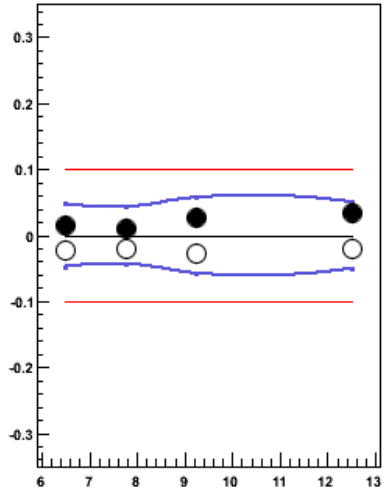
— 10% line

○  $|Z_{vertex}| < 45$

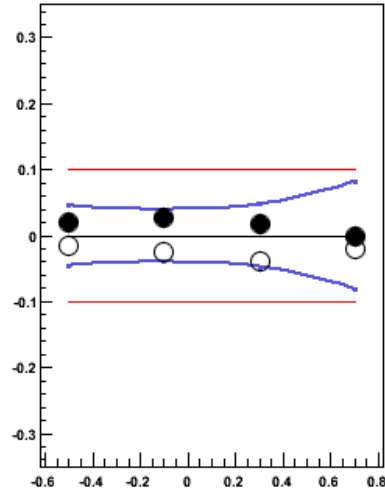
●  $|Z_{vertex}| < 35$

# Systematic uncertainties: deltaR

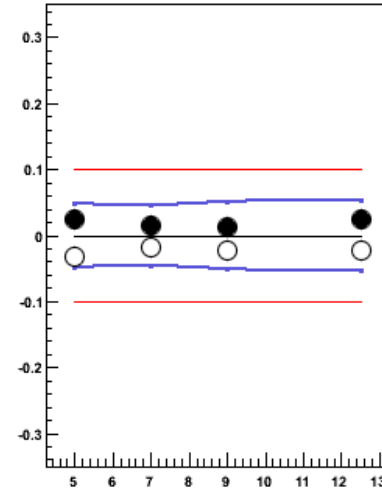
$E_T^\gamma \delta R$



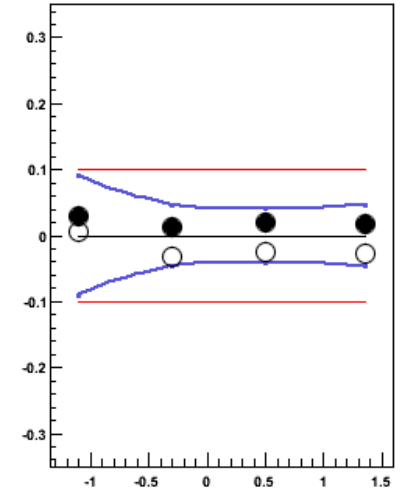
$\eta^\gamma \delta R$



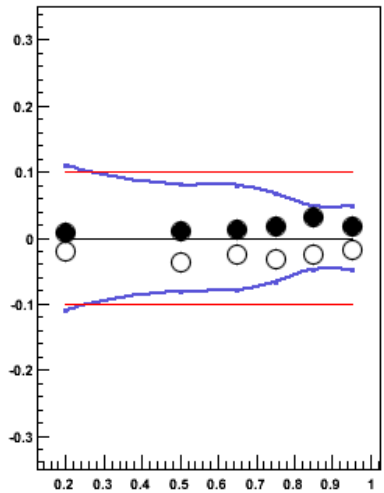
$E_T^{\text{jet}} \delta R$



$\eta^{\text{jet}} \delta R$



$x_\gamma^{\text{meas}} \delta R$



**Standard cut:**

- deltaR track isolation in cone 0.2

**Vary cone radius by  $\pm 0.1$**

**$\gamma + \text{jet}$**

— Rel. statistical uncertainties

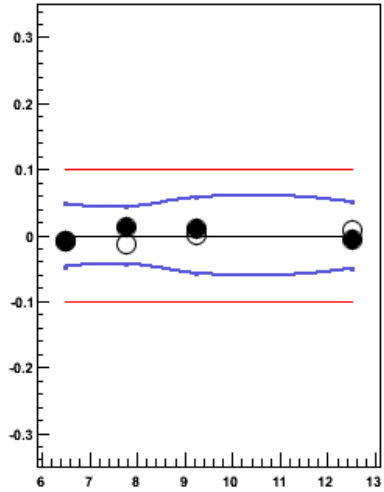
— 10% line

○  $\delta R 0.3$

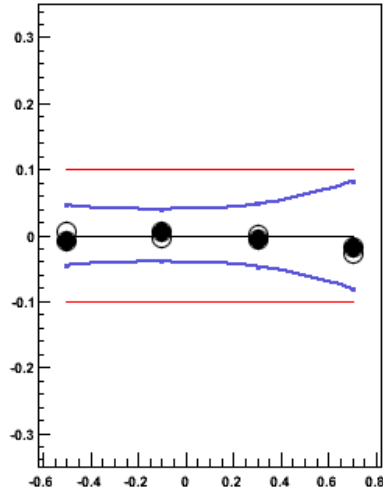
●  $\delta R 0.1$

# Systematic uncertainties: deltaZ fit range

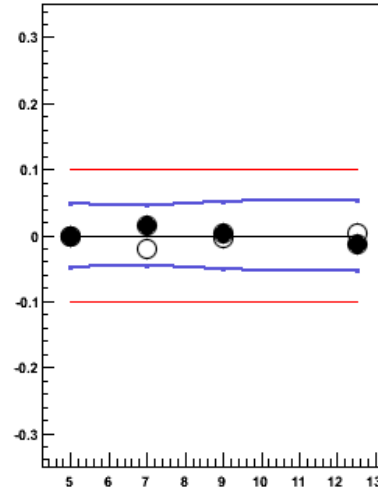
$E_T^\gamma \delta Z$



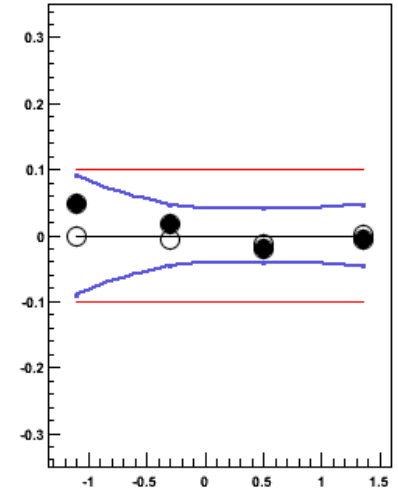
$\eta^\gamma \delta Z$



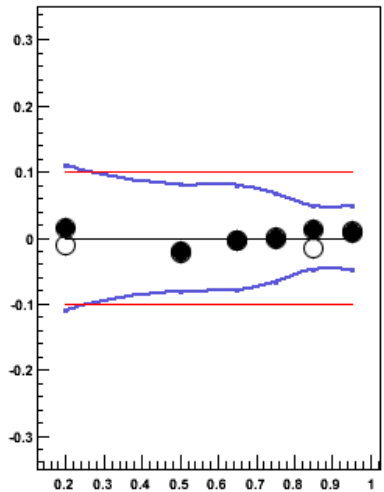
$E_T^{\text{jet}} \delta Z$



$\eta^{\text{jet}} \delta Z$



$x_\gamma^{\text{meas}} \delta Z$



Standard range:

0.05 – 0.8

deltaZ upper limit variation  
between 0.6 and 1.0

$\gamma$ +jet

— Rel.statistical uncertainties  $\delta Z$

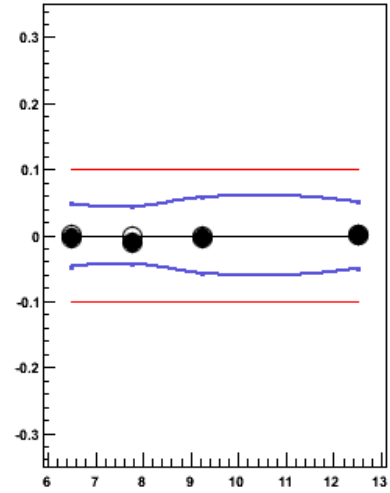
— 10% line

○  $\delta Z$  fit range 1.0

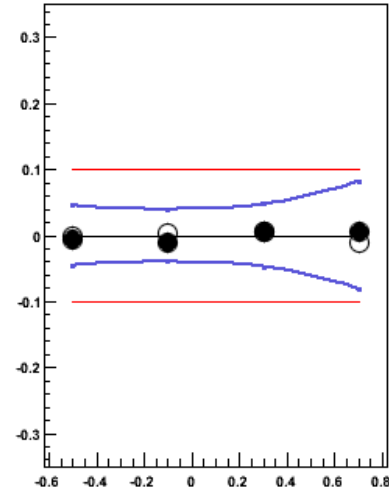
●  $\delta Z$  fit range 0.6

# Systematic uncertainties: Fraction EMC

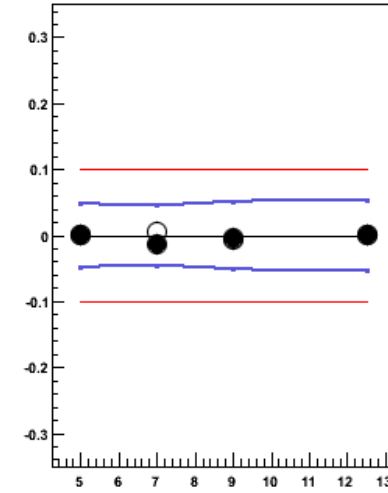
$E_T^{\gamma}$  fraction EMC



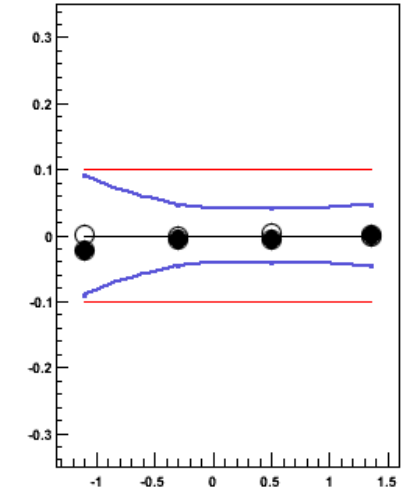
$\eta^{\gamma}$  fraction EMC



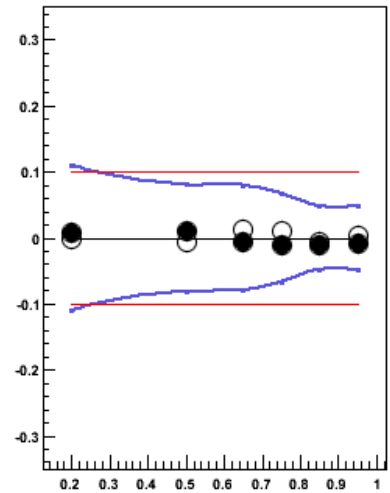
$E_T^{\text{jet}}$  fraction EMC



$\eta^{\text{jet}}$  fraction EMC



$x_T^{\text{meas}}$  fraction EMC



Standard cut:

- $Z_{\text{ufoEemc}}/Z_{\text{ufoEcal}} > 0.9$

Vary the em fraction in the photon  
zufe by  $\pm 2.5\%$

$\gamma + \text{jet}$

— Rel. statistical uncertainties

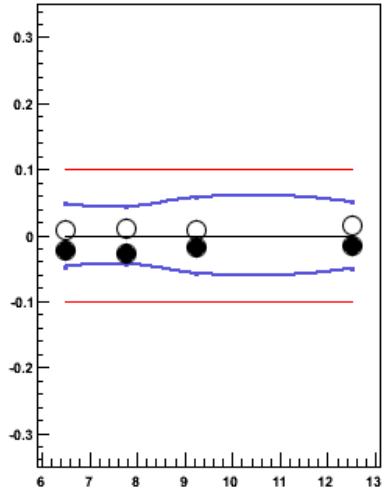
— 10% line

○ Fraction EMC +0.025

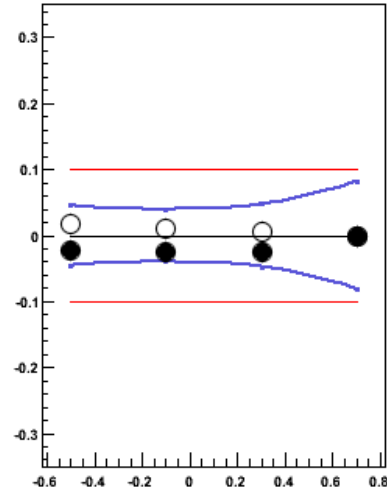
● Fraction EMC -0.025

# Systematic uncertainties: Track momentum

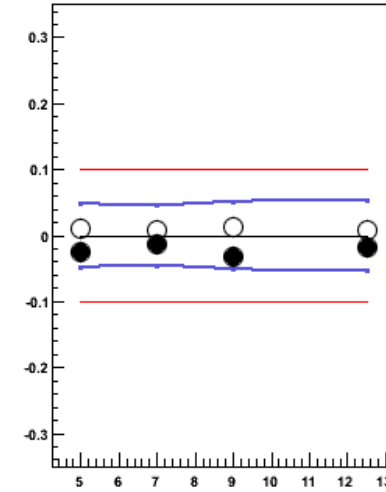
$E_T^{\gamma}$  Track Magnitude



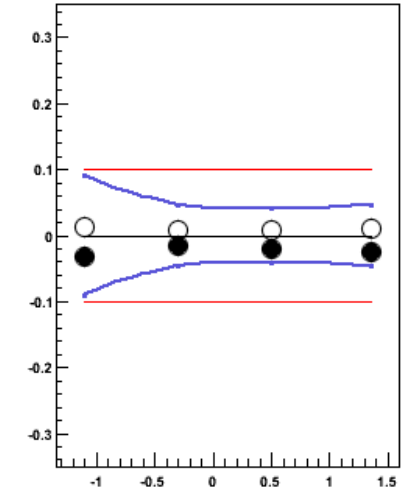
$\eta^{\gamma}$  Track Magnitude



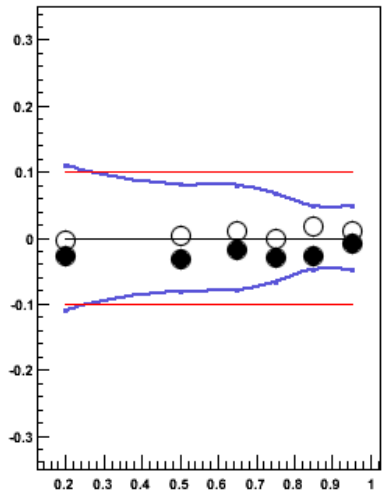
$E_T^{\text{jet}}$  Track Magnitude



$\eta^{\text{jet}}$  Track Magnitude



$X_T^{\text{meas}}$  Track Magnitude



Standard cut:

- Track momentum > 250 MeV

Vary track momentum by 100 MeV

$\gamma + \text{jet}$

— Rel. statistical uncertainties

— 10% line

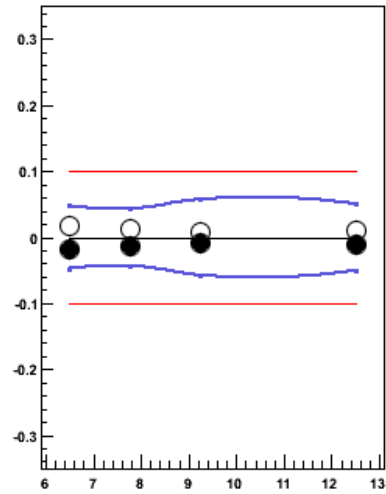
○  $p_{\text{track}} > 350 \text{ MeV}$

●  $p_{\text{track}} > 150 \text{ MeV}$

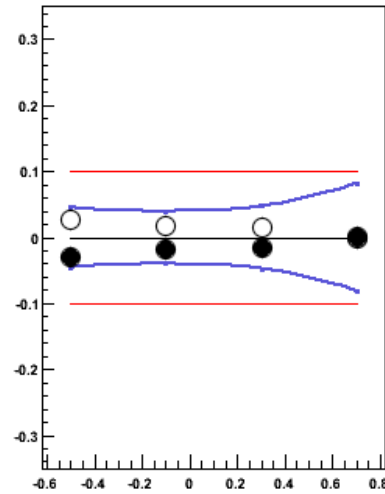


# Systematic uncertainties: PYTHIA dir / res

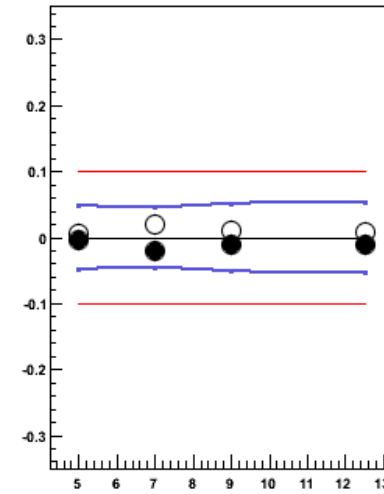
$E_T^\gamma$  Dir / Res ratio



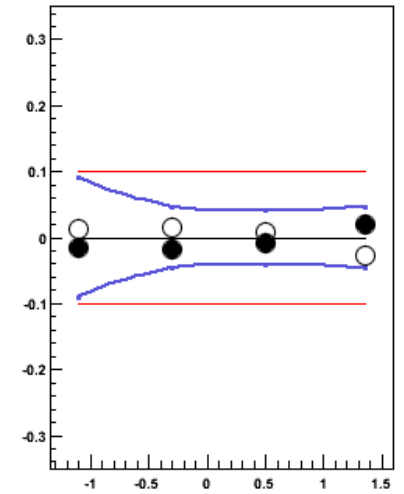
$\eta^\gamma$  Dir / Res ratio



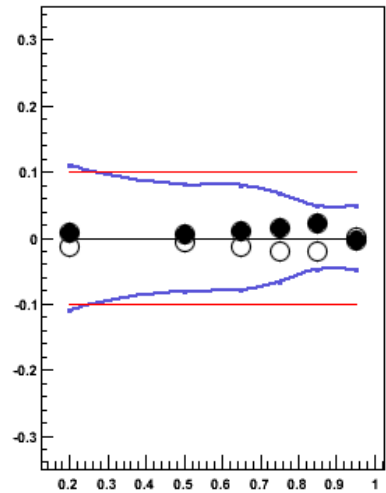
$E_T^{\text{jet}}$  Dir / Res ratio



$\eta^{\text{jet}}$  Dir / Res ratio



$X_\gamma^{\text{meas}}$  Dir / Res ratio



Standard direct/ resolved ratio:  
50% / 40%

Vary fraction of resolved by  $\pm 15\%$ :

35% / 55%  
65% / 25%

$\gamma + \text{jet}$

— Rel. statistical uncertainties

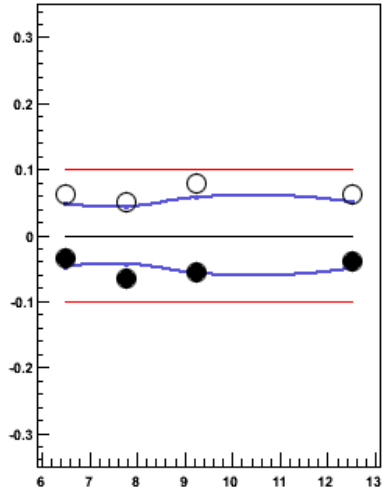
— 10% line

○ -15% resolved

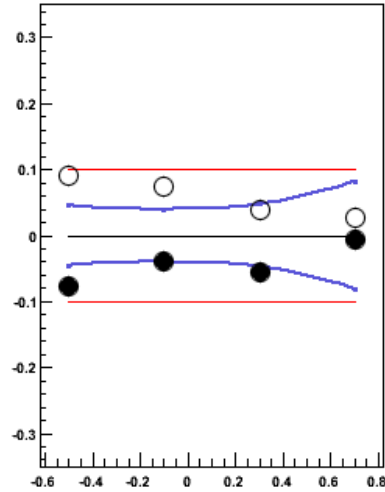
● +15% resolved

# Systematic uncertainties: E- $\gamma$ , E-jet variation

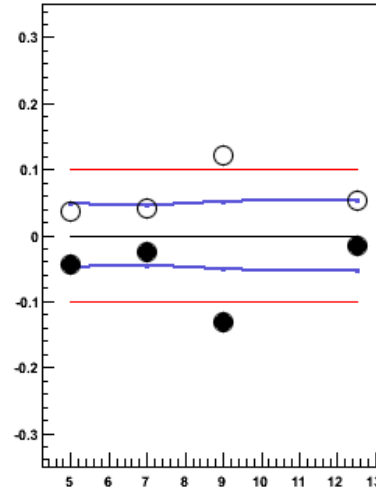
$E_T^\gamma E_{\gamma, \text{Jet}}$



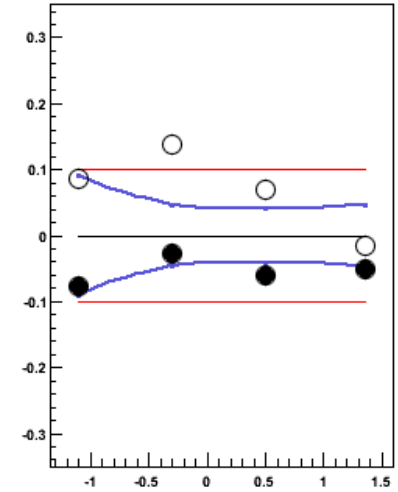
$\eta^\gamma E_{\gamma, \text{Jet}}$



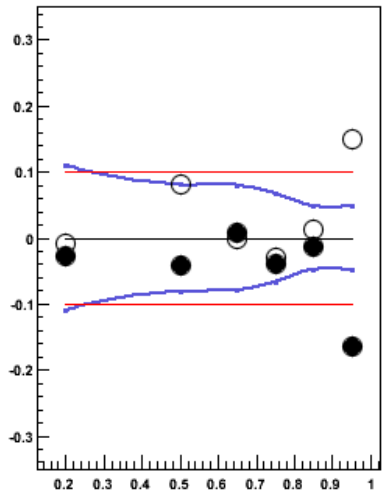
$E_T^{\text{jet}} E_{\gamma, \text{Jet}}$



$\eta^{\text{jet}} E_{\gamma, \text{Jet}}$



$x_\gamma^{\text{meas}} E_{\gamma, \text{Jet}}$



**Standard cuts:**

- $6 < E_T^{\text{zifo}} < 15$  GeV
- $4 < E_T^{\text{jet}} < 35$  GeV

**Vary  $E_\gamma$  by  $\pm 2\%$  simultaneously with E-jet depending on its value:  $\pm 4\%$  if  $E_T^{\text{jet}} < 6$  GeV,  $\pm 2.5\%$  if  $6 < E_T^{\text{jet}} < 10$  GeV,  $\pm 1.5\%$  if  $E_T^{\text{jet}} > 10$  GeV**

**$\gamma + \text{jet}$**

— Rel. statistical uncertainties

— 10% line

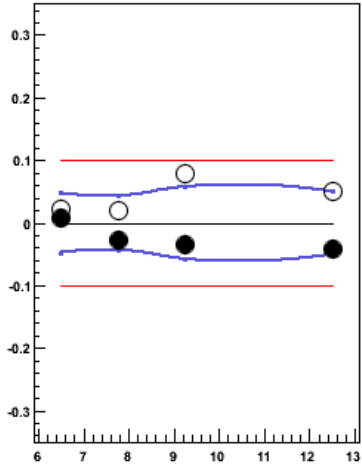
○ up. limit

● low. limit

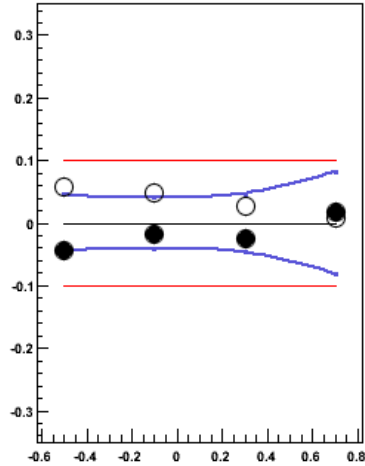
**Dominating systematic**

# Systematic uncertainties: independent E- $\gamma$ variation

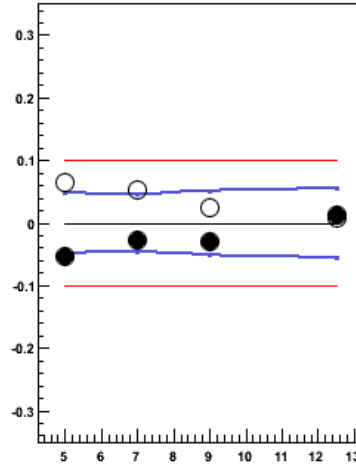
$E_T^\gamma E_\gamma$



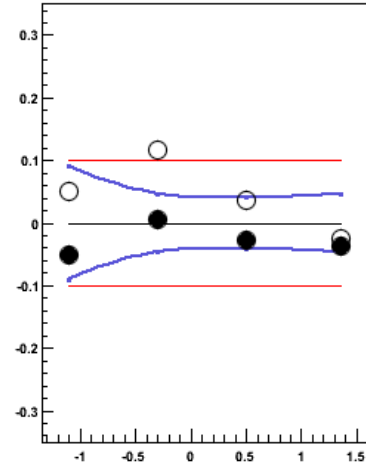
$\eta^\gamma E_\gamma$



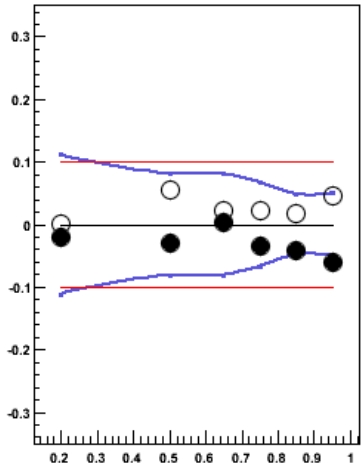
$E_T^{\text{jet}} E_\gamma$



$\eta^{\text{jet}} E_\gamma$



$X_\gamma^{\text{meas}} E_\gamma$



**Standard cuts:**

$$\bullet 6 < E_T^{\text{z UFO}} < 15 \text{ GeV}$$

**Vary  $E_\gamma$  by  $\pm 2\%$**

— *Rel. statistical uncertainties*

— *10% line*

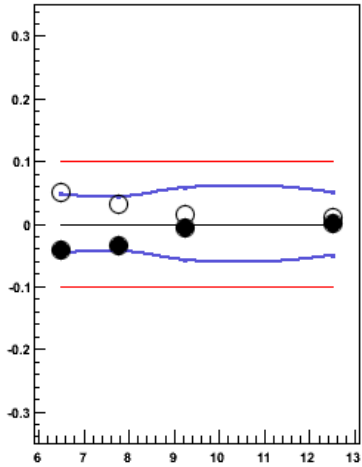
○  $E_\gamma + 2\%$

●  $E_\gamma - 2\%$

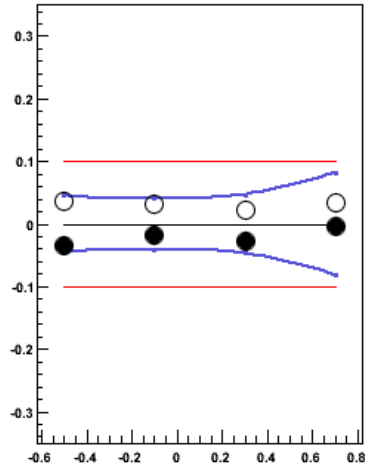
**Not used for current cross-section pictures**

# Systematic uncertainties: independent E-jet variation

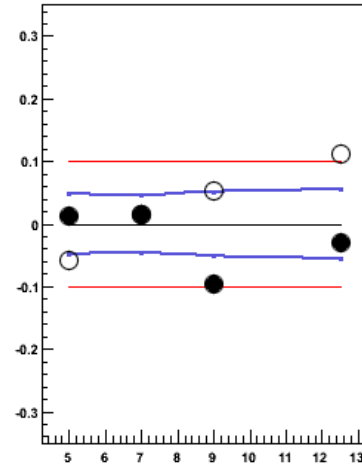
$E_T^{\gamma}$  UncorJE



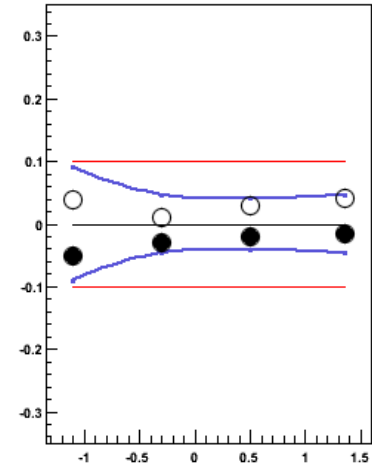
$\eta^{\gamma}$  UncorJE



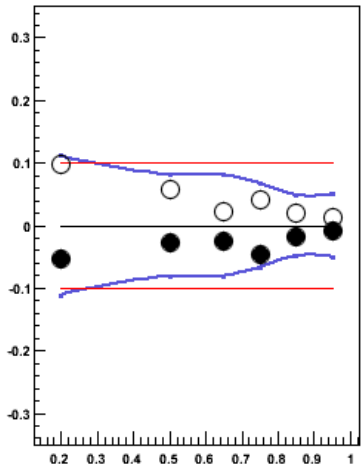
$E_T^{\text{jet}}$  UncorJE



$\eta^{\text{jet}}$  UncorJE



$x_{\gamma}^{\text{mass}}$  UncorJE



**Standard cuts:**

$$\bullet 4 < E_T^{\text{jet}} < 35 \text{ GeV}$$

**Vary Et-jet depending on its value:**

**+/-4% if  $E_T^{\text{jet}} < 6 \text{ GeV}$ , +/-2.5% if**

**$6 < E_T^{\text{jet}} < 10 \text{ GeV}$**

**+/-1.5% if  $E_T^{\text{jet}} > 10 \text{ GeV}$**

**AND add 2% electron uncertainty i.e.**

**+/-6% if  $E_T^{\text{jet}} < 6 \text{ GeV}$**

**Not used for current cross-section pictures**

— *Rel. statistical uncertainties*

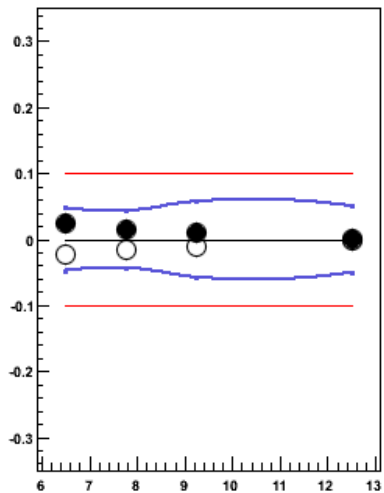
— *10% line*

○ *variation up*

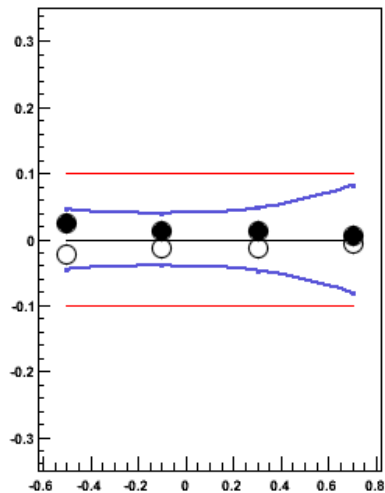
● *variation down*

# Systematic uncertainties: PYTHIA fragmentation

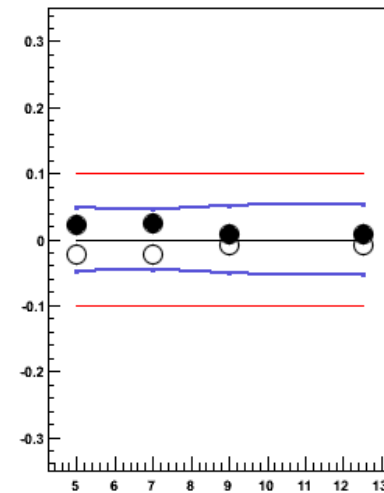
$E_T^{\gamma}$  Fragmentation



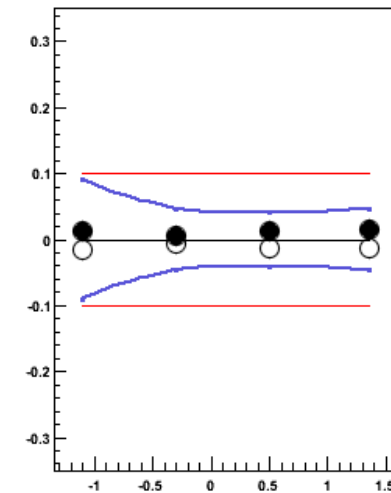
$\eta^{\gamma}$  Fragmentation



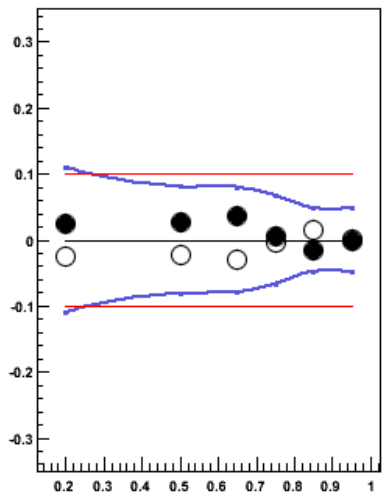
$E_T^{\text{jet}}$  Fragmentation



$\eta^{\text{jet}}$  Fragmentation



$X_T^{\text{meas}}$  Fragmentation



Standard value:

- 10%

Vary fraction of fragmentation direct and resolved simultaneously by  $\pm 5\%$

$\gamma + \text{jet}$

— Rel. statistical uncertainties

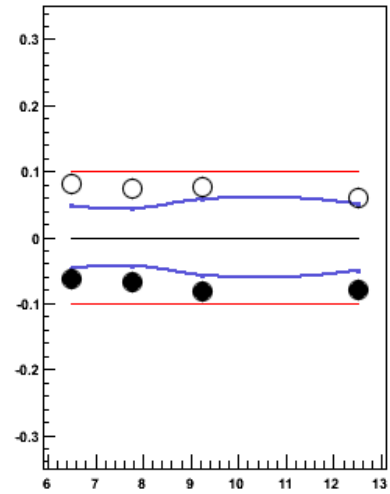
— 10% line

○ +5% Fragmentation

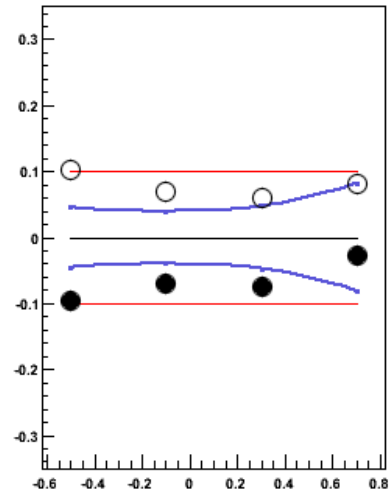
● -5% Fragmentation

# Overall Systematic uncertainties

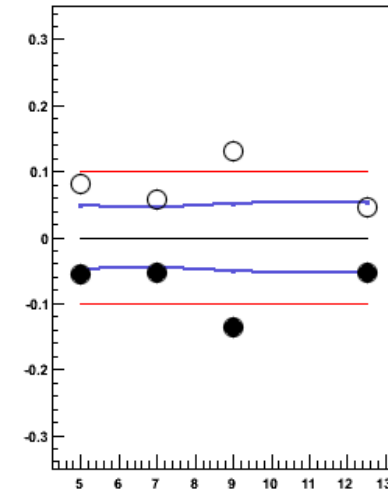
$E_T^{\gamma}$  Overall



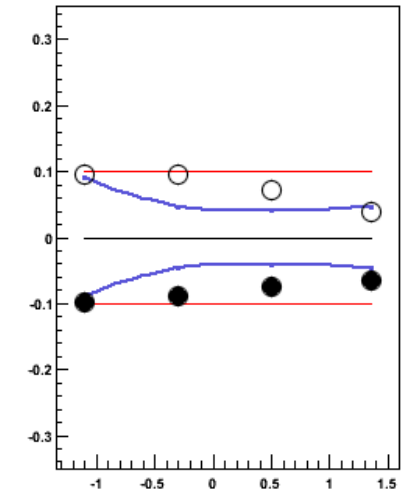
$\eta^{\gamma}$  Overall



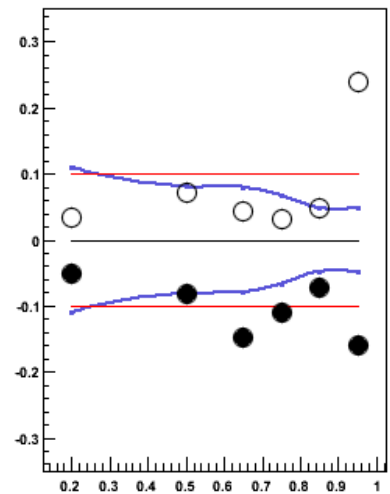
$E_T^{\text{jet}}$  Overall



$\eta^{\text{jet}}$  Overall



$X_{\gamma}^{\text{meas}}$  Overall

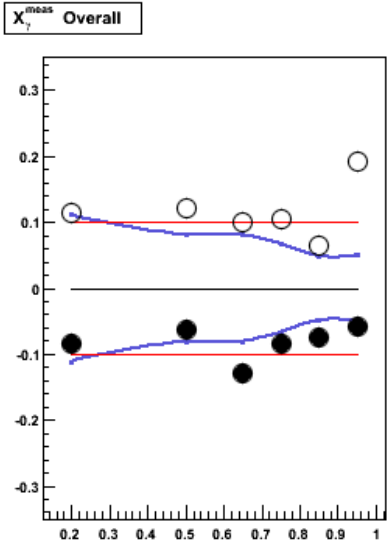
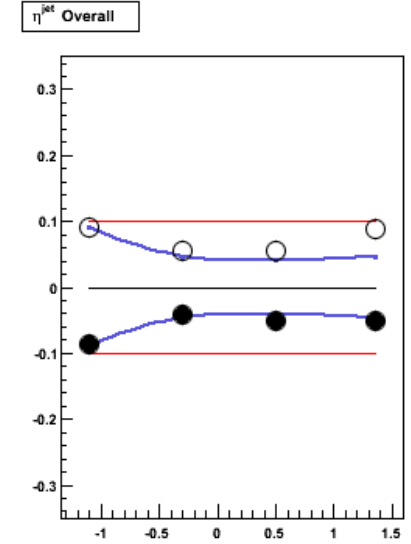
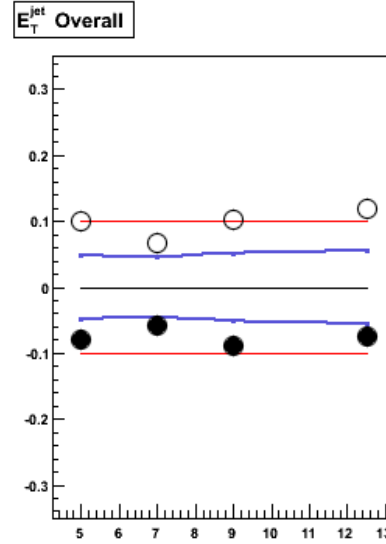
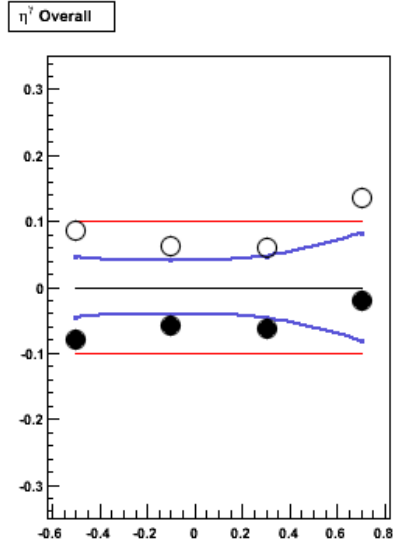
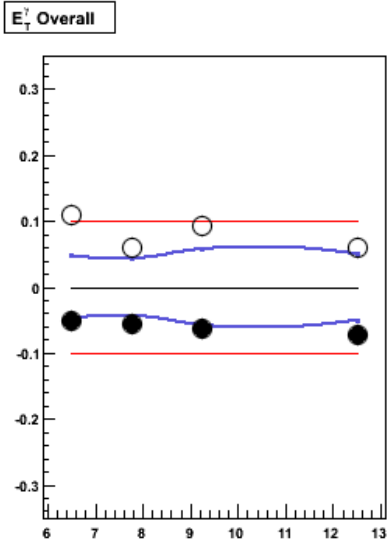


Systematics are close to statistical errors.  
Largest contributions: from E variation overall  
and from HERWIG to X-gamma.

$\gamma$ +jet

- *Rel. statistical uncertainties*
- *10% line*
- *upper sum*
- *lower sum*

# Overall Systematic uncertainties with independent E variations

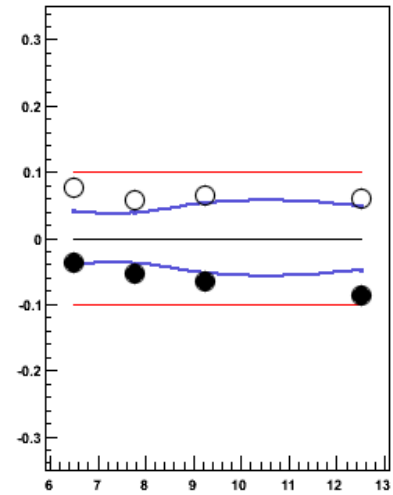


Not used for current cross-section pictures

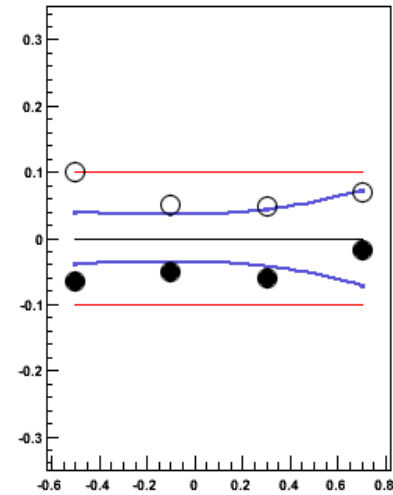
- *Rel. statistical uncertainties*
- *10% line*
- *upper sum*
- *lower sum*

# Overall Systematic uncertainties: Inclusive $\gamma$

$E_T^{\gamma}$  Overall



$\eta^{\gamma}$  Overall



— *Rel. statistical uncertainties*

— *10% line*

○ *upper sum*

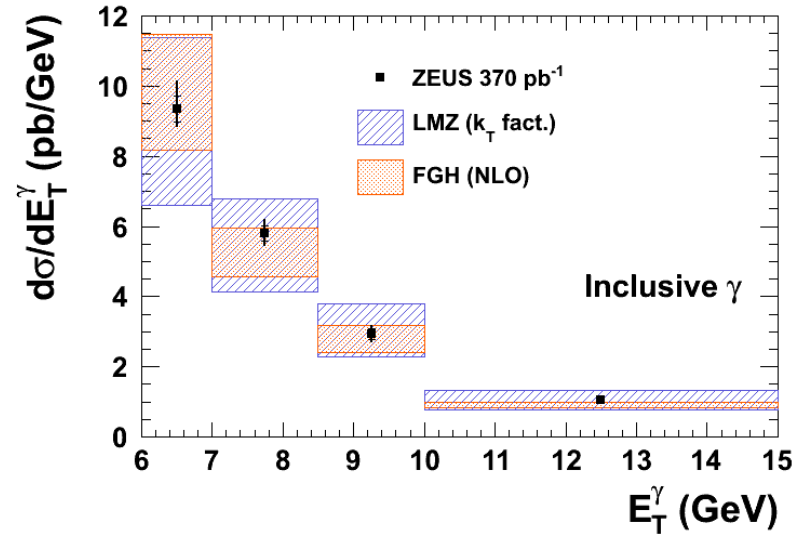
● *lower sum*

Systematics are comparable with the statistical errors.



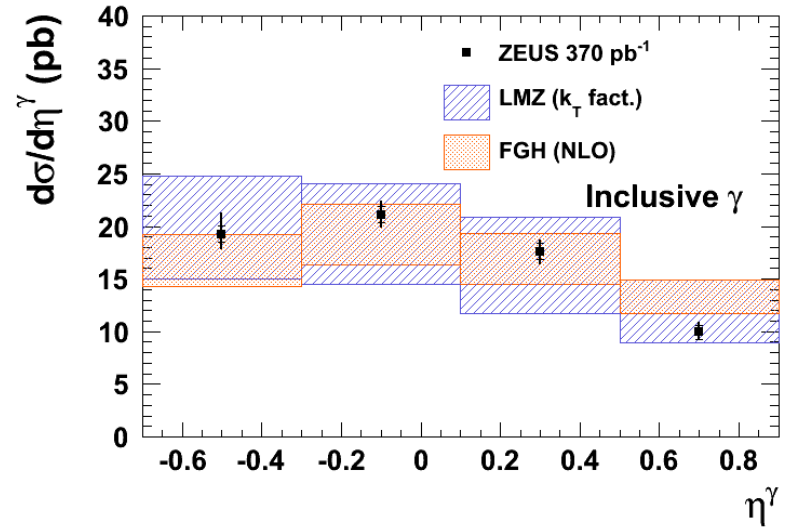
# Inclusive cross-sections

ZEUS



(a)

ZEUS



(b)

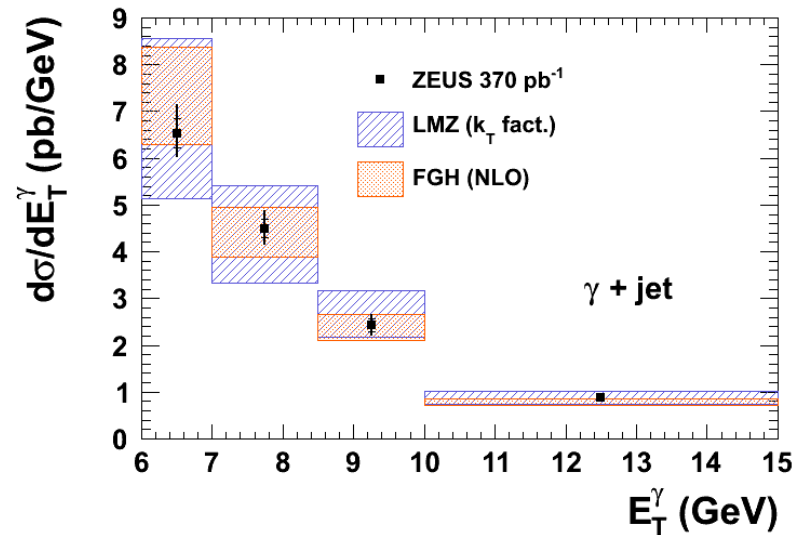
Cross sections as a function of (a)  $E_T^\gamma$  and (b)  $\eta^\gamma$ , for events containing an isolated photon compared to predictions from FGH and LMZ.

$$\frac{d\sigma}{dY} = \frac{\mathcal{A} N(\gamma)}{\mathcal{L} \Delta Y}$$

$N(\gamma)$  is the number of photons extracted from the fit,  $\Delta Y$  is the bin width,  $\mathcal{L}$  is the total integrated luminosity, and  $\mathcal{A}$  is the acceptance correction and was calculated using Monte Carlo from the ratio of the number of events generated to those reconstructed in a given bin. Its value was typically around 1.2.

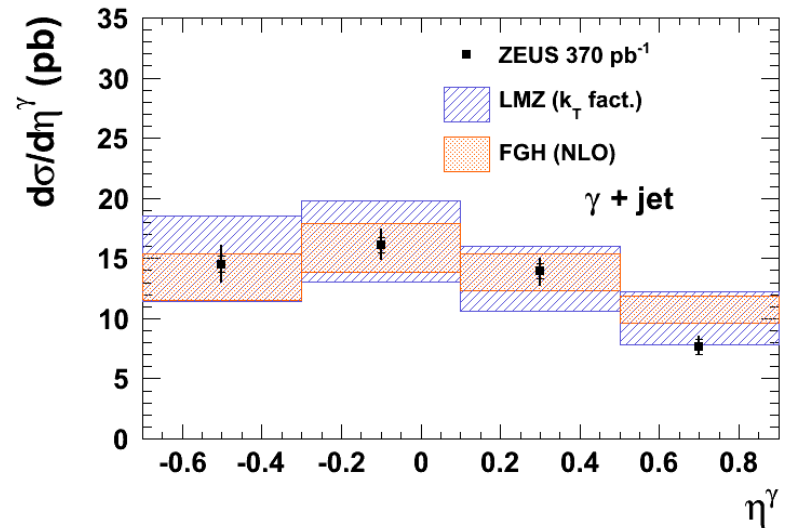
# Cross sections for photon plus jet

ZEUS



(a)

ZEUS



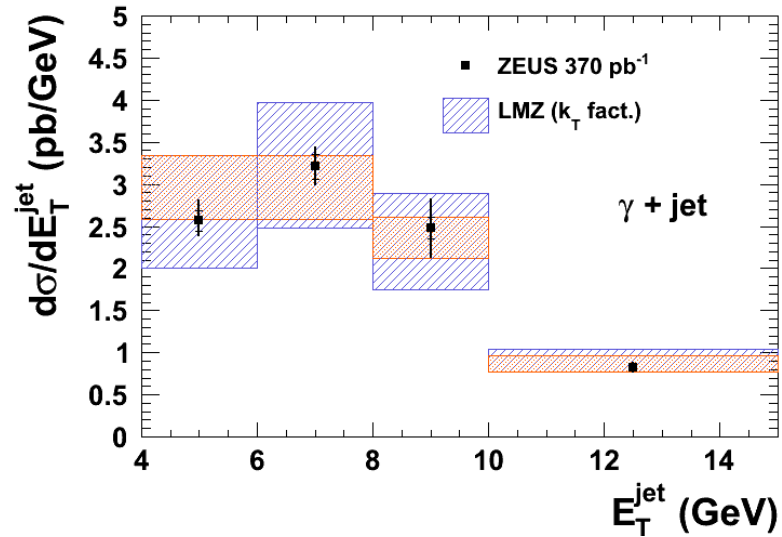
(b)

Cross sections as a function of (a)  $E_T^\gamma$  and (b)  $\eta^\gamma$ , for events containing an isolated photon accompanied by jet compared to predictions from FGH and LMZ.

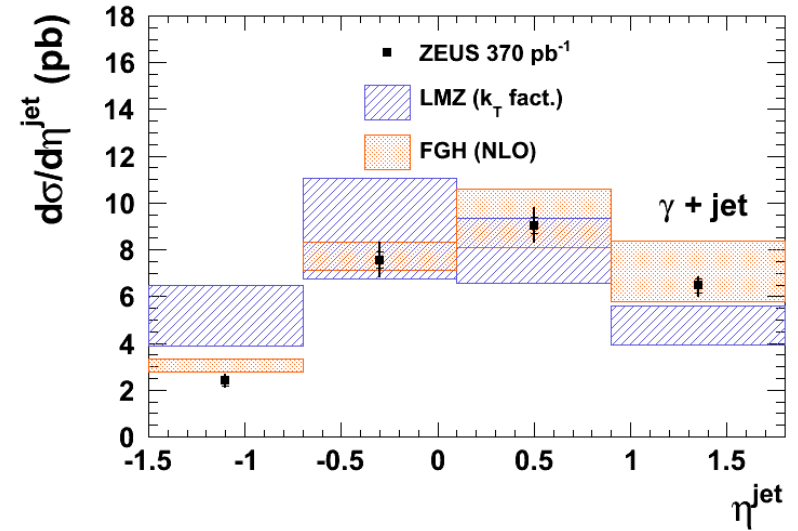
All acceptance factors were calculated using a model containing equal numbers of direct and resolved PYTHIA events and a 10% admixture of fragmentation events.

# Cross sections for photon plus jet

ZEUS



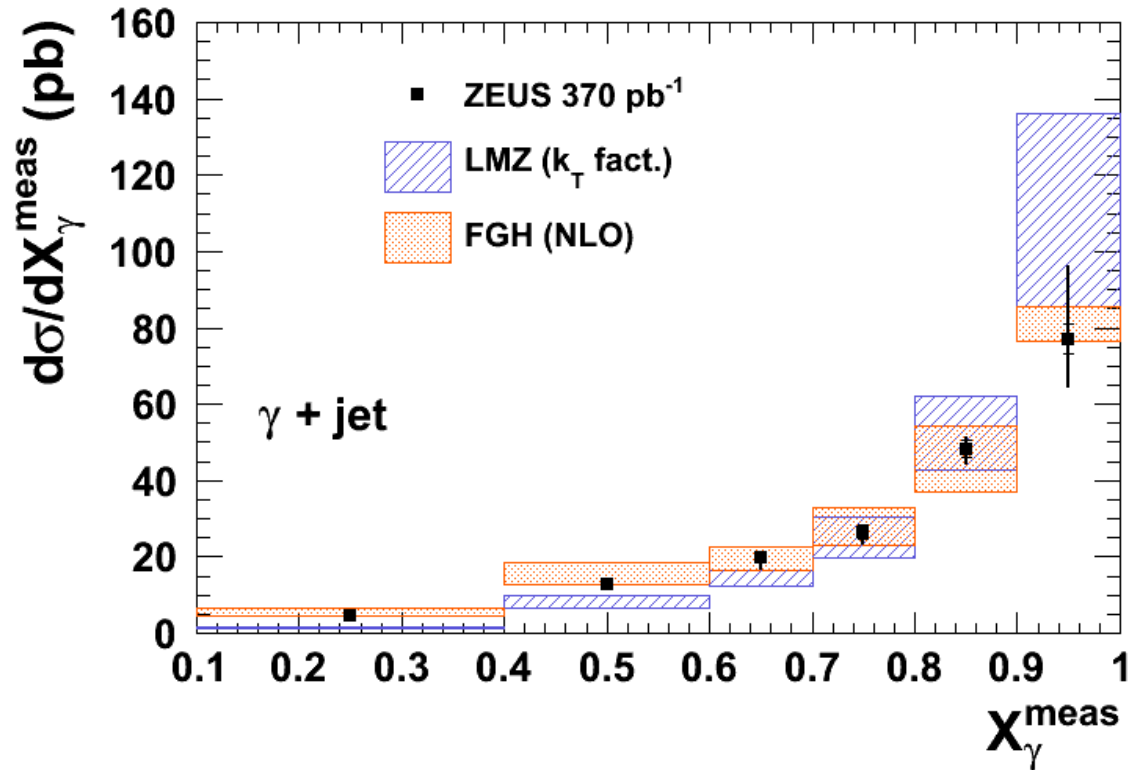
ZEUS



Cross sections as a function of (a)  $E_T^{\text{jet}}$  and (b)  $\eta^{\text{jet}}$ , for events containing an isolated photon accompanied by jet compared to predictions from FGH and LMZ.

# Cross sections for photon plus jet

## ZEUS



Cross sections as a function of  $X_\gamma^{\text{MEAS}}$ , for events containing an isolated photon accompanied by jet compared to predictions from FGH and LMZ.

## Summary

A complete set of photoproduction cross sections has been calculated, for inclusive prompt photons and prompt photons plus a jet.

Inclusive: Good description within errors by both theories.

Photon+jet: LMZ is systematically low at small  $x$ -gamma and the central value is high for the  $x$ -gamma 0.9-1.0 bin.

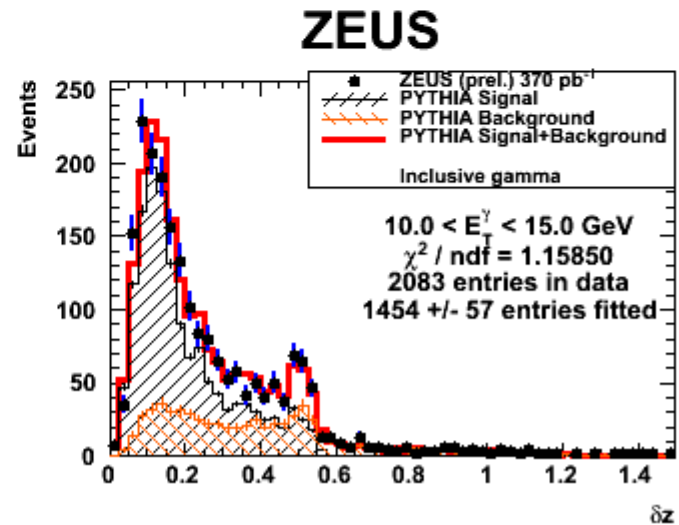
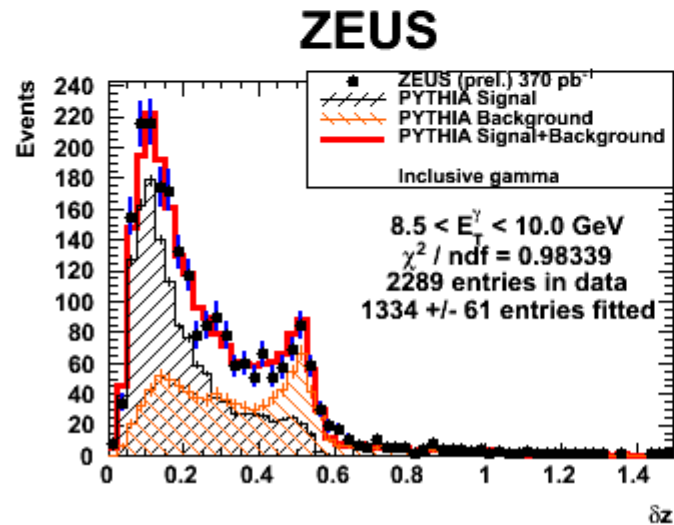
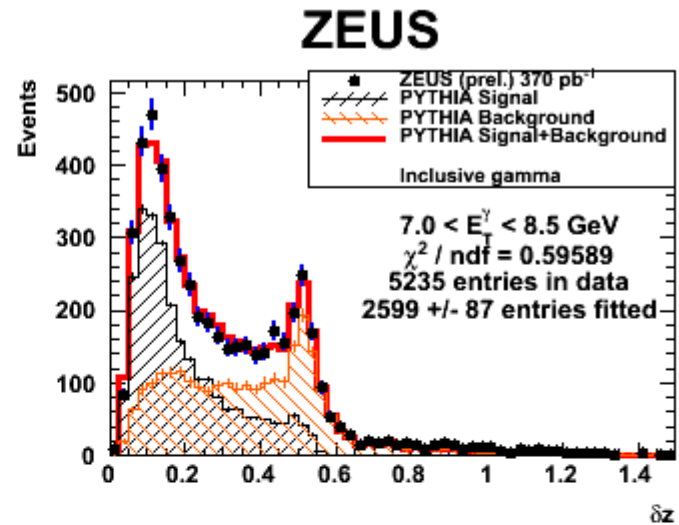
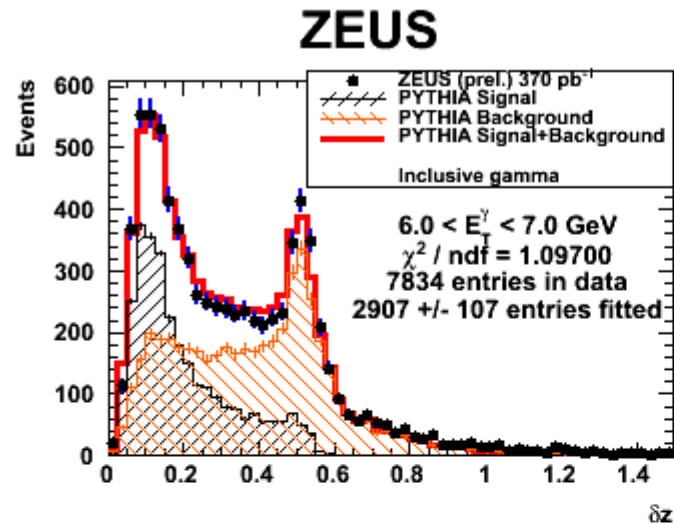
The trend for  $\eta^{\text{jet}}$  is better described by FGH than LMZ.

# Backups

- Fits in bins of studied variables
- Inclusive systematics
- Hadronisation correction factors
- Acceptance
- E scale systematic
- Post-fitting comparison with the PYTHIA model
- PYTHIA and HERWIG X-gamma
- Comparison between analyses

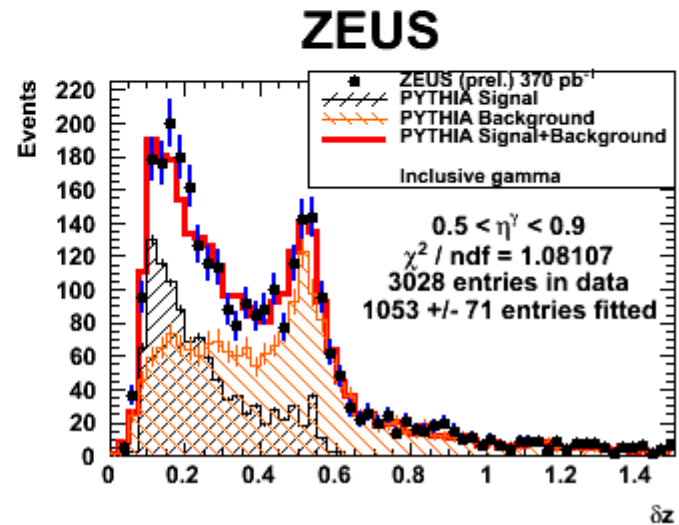
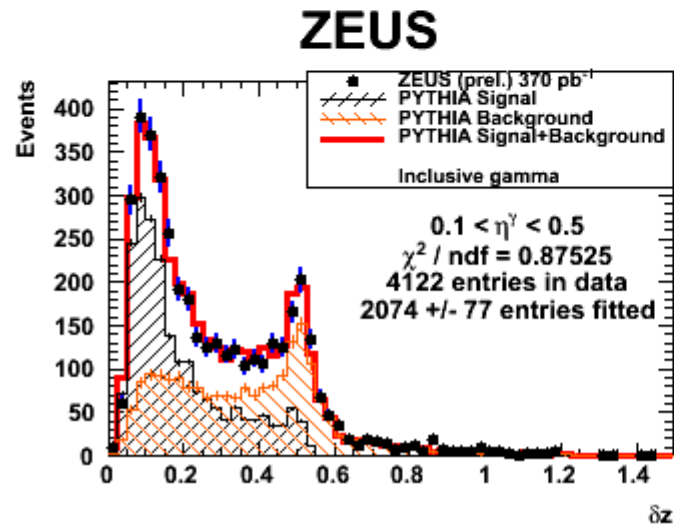
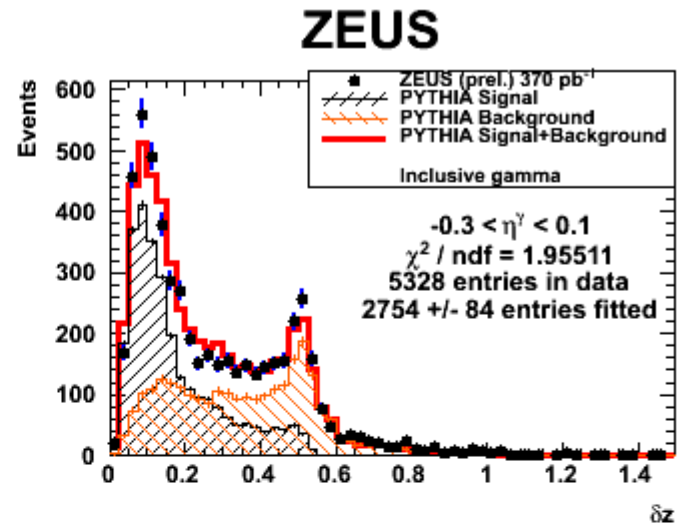
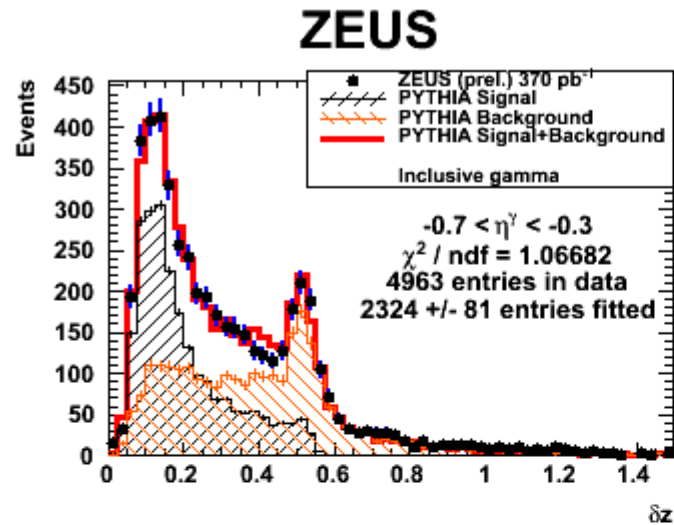
# **Fits in bins of studied variables**

# Fits in bins (Inclusive)

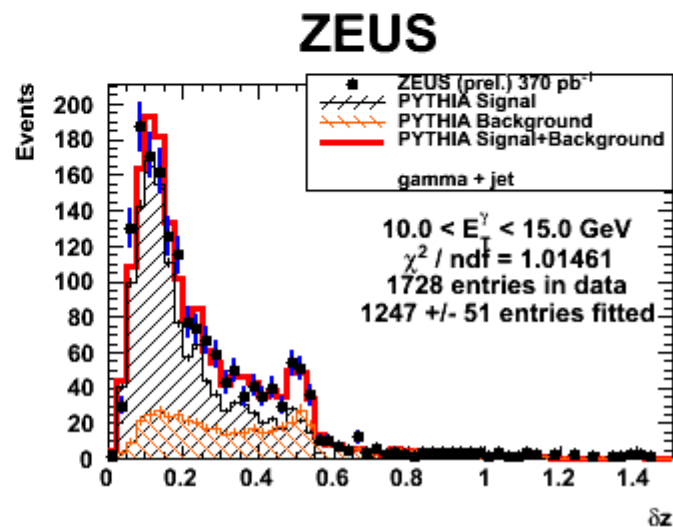
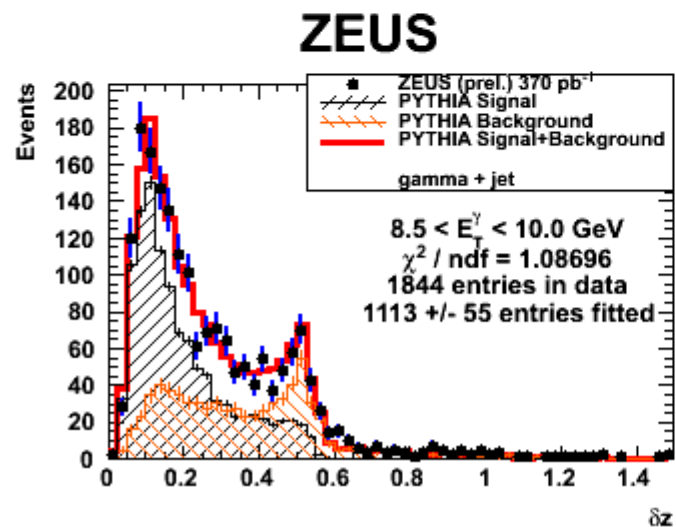
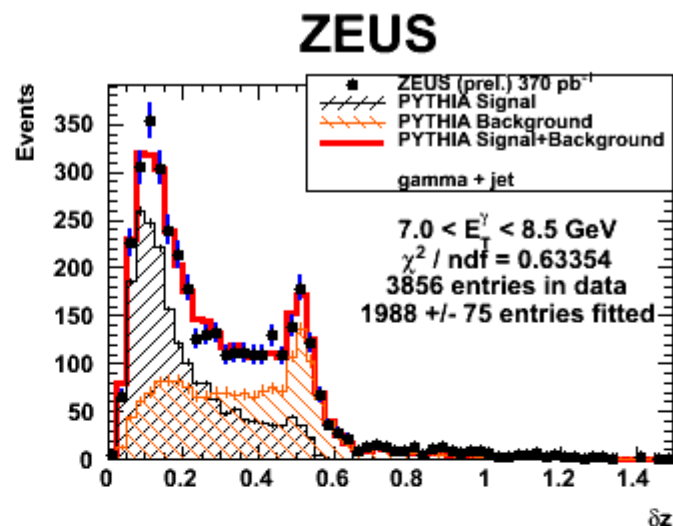
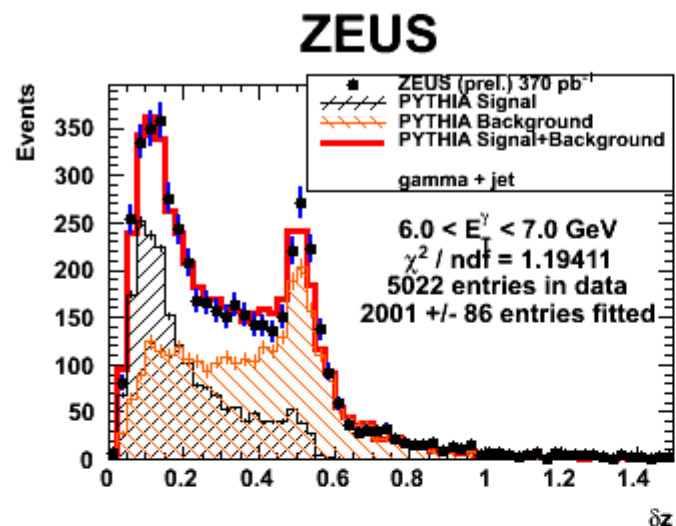




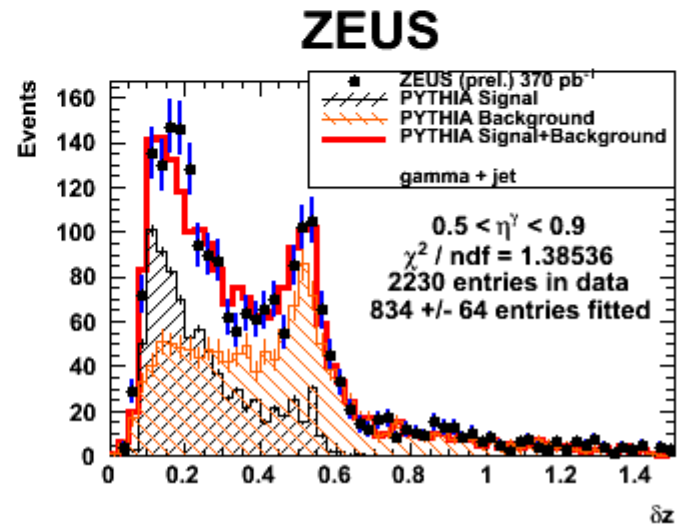
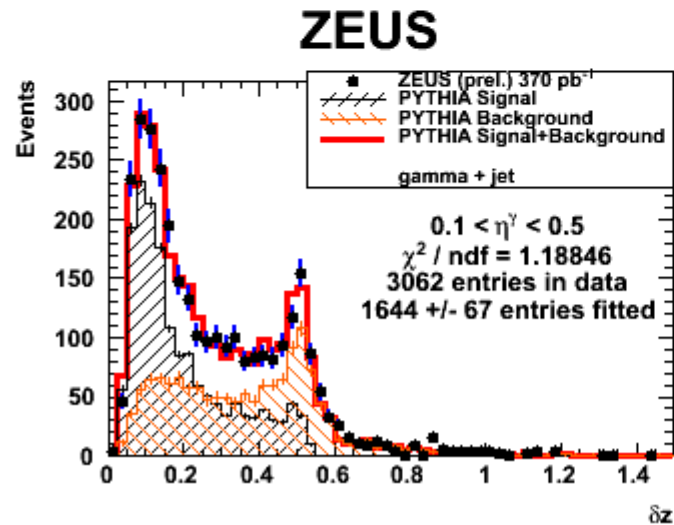
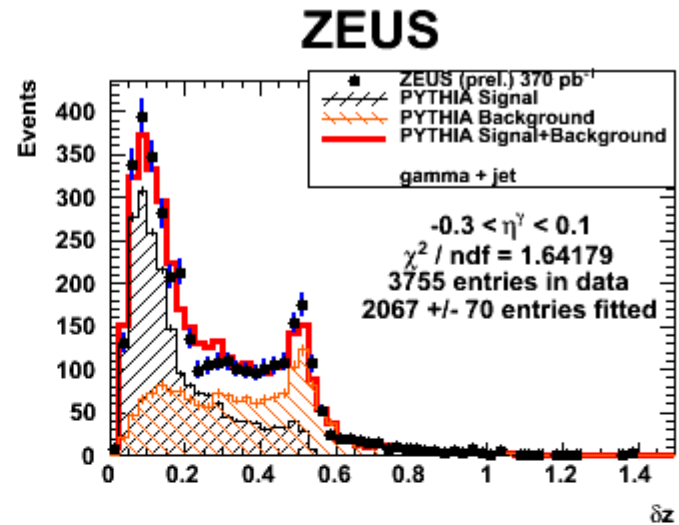
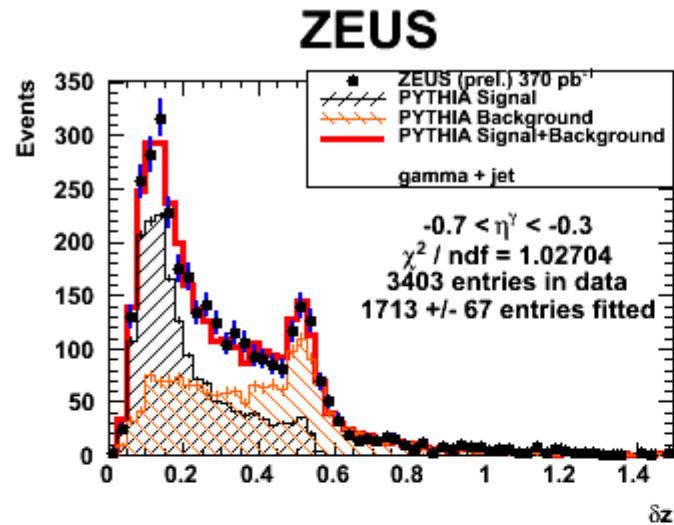
# Fits in bins (Inclusive)



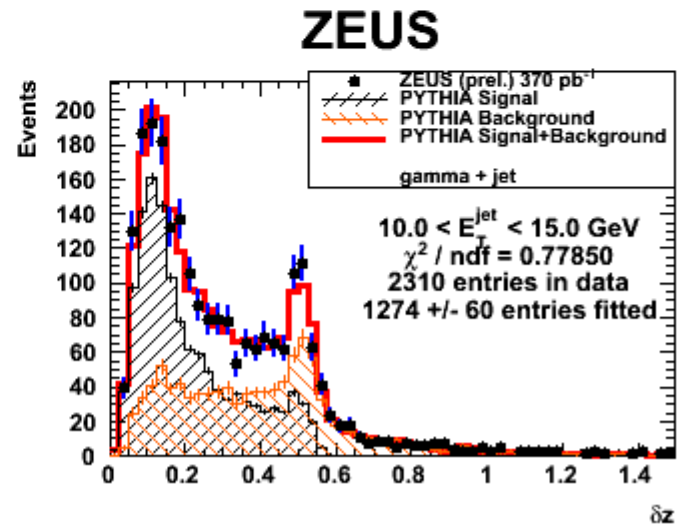
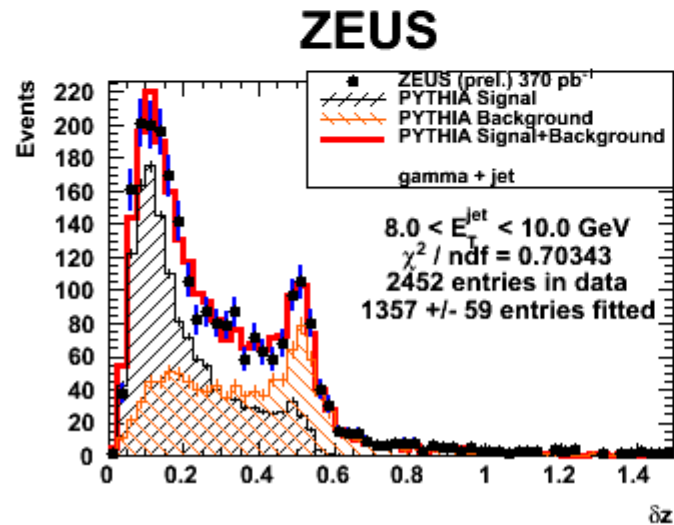
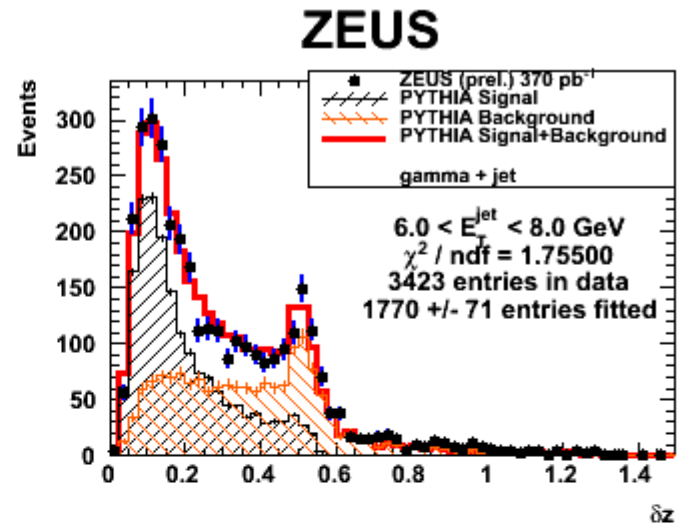
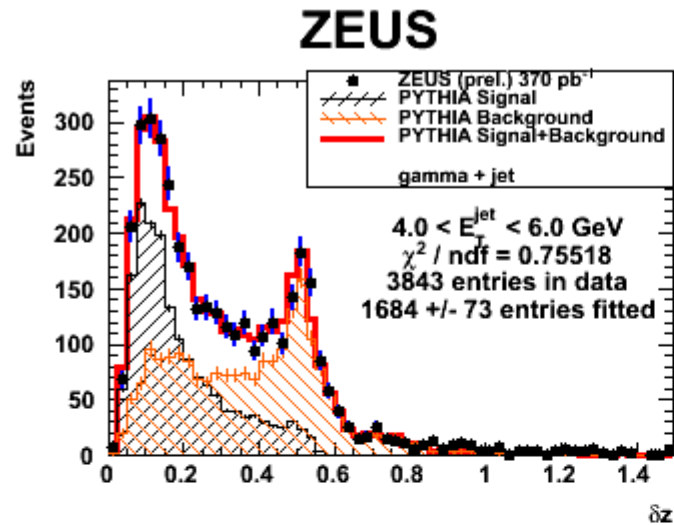
# Fits in bins (photon+jet)



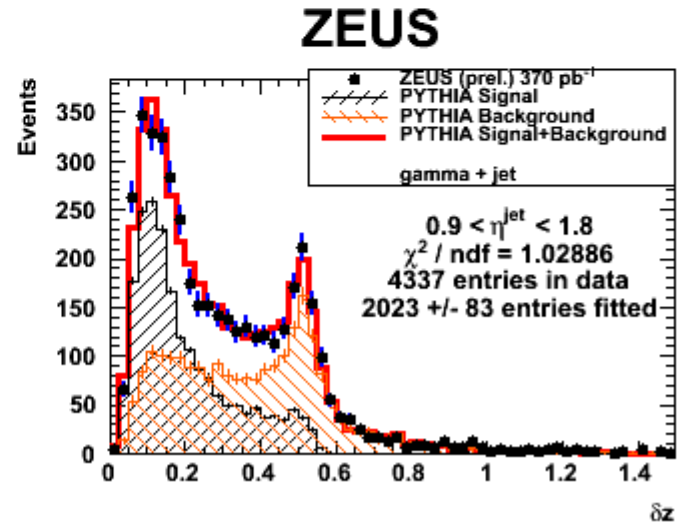
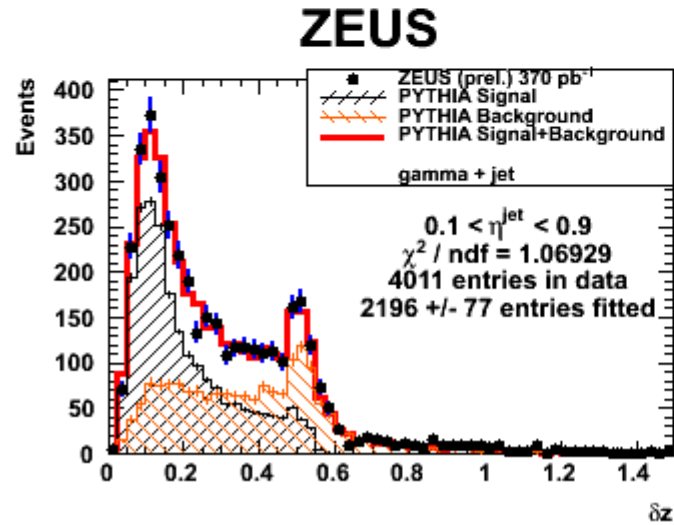
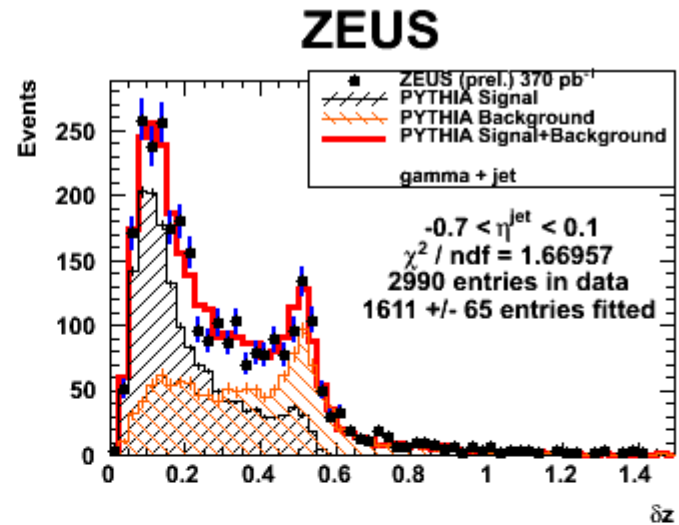
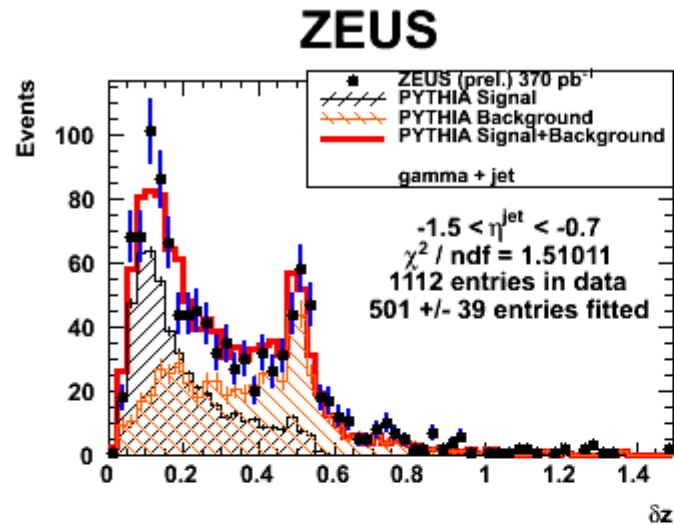
# Fits in bins (photon+jet)



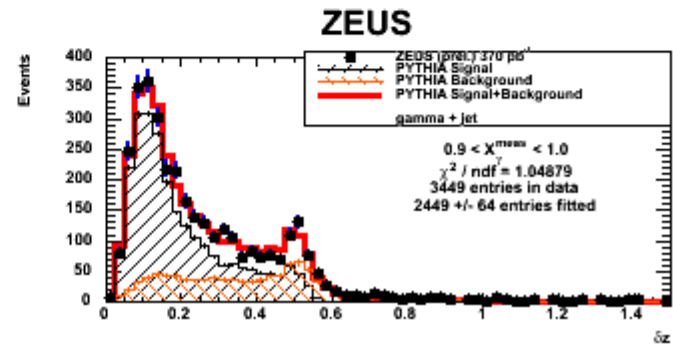
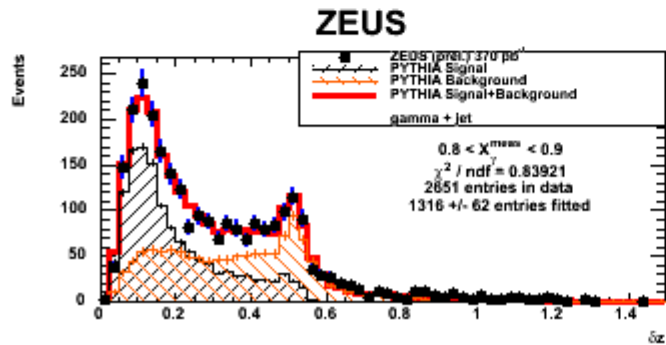
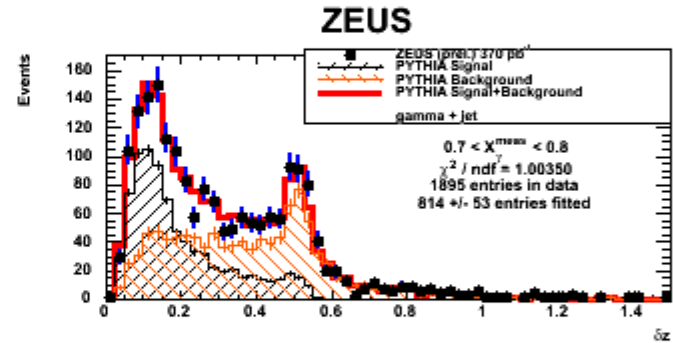
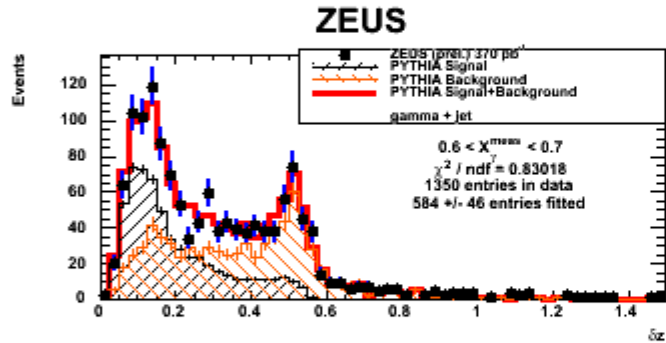
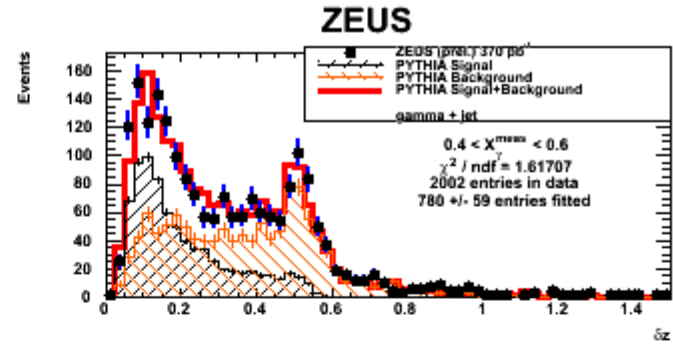
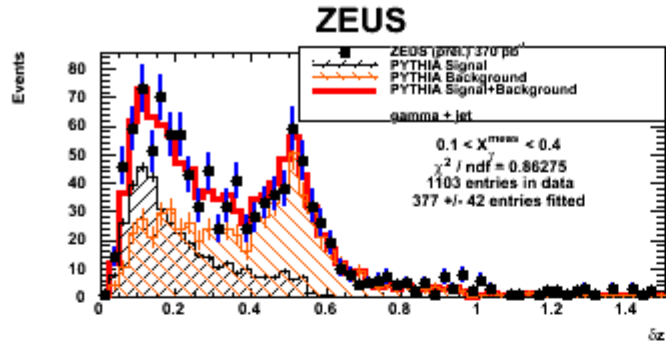
# Fits in bins (photon+jet)



# Fits in bins (photon+jet)

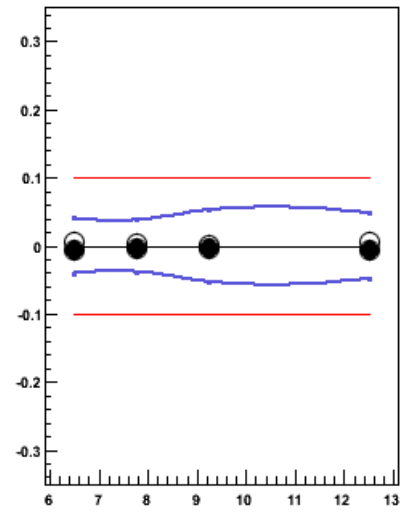


# Fits in bins (photon+jet)

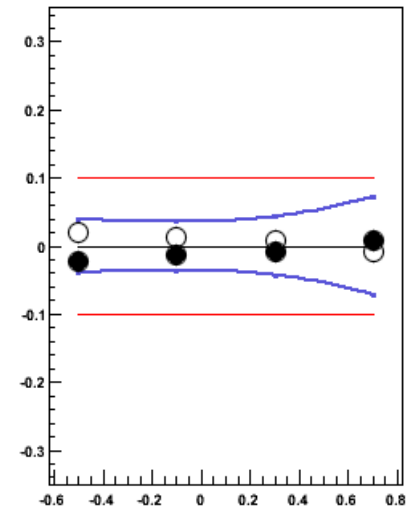


# **Inclusive systematics**

$E_T^{\gamma}$  Dir / Res ratio



$\eta^{\gamma}$  Dir / Res ratio



— *Rel. statistical uncertainties*

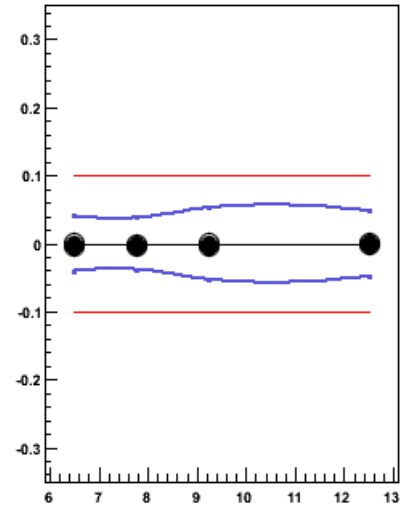
— *10% line*

○ *-15% resolved*

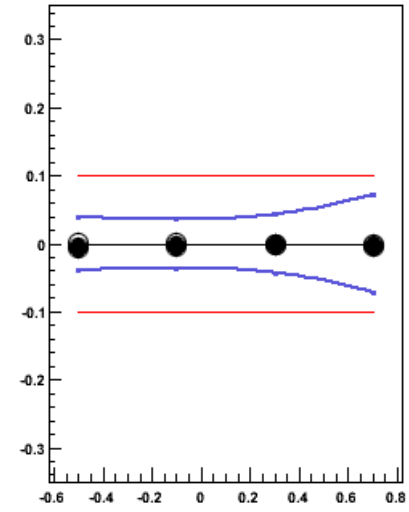
● *+15% resolved*



$E_{\tau}^{\gamma}$  Z-Vertex



$\eta^{\gamma}$  Z-Vertex



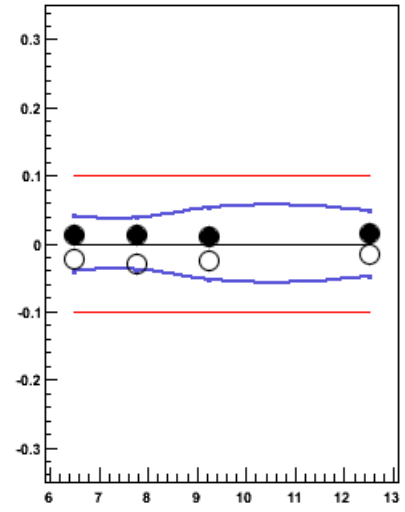
— *Rel.statistical uncertainties*

— *10% line*

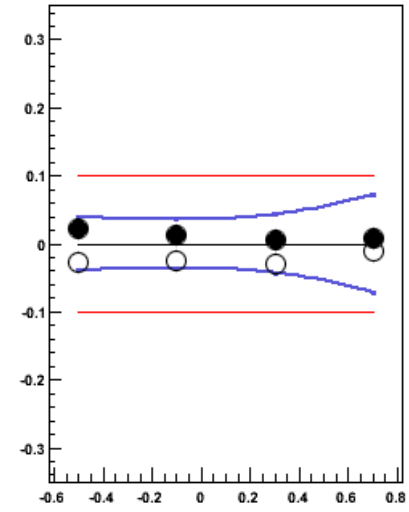
○  $|Z_{vertex}| < 45$

●  $|Z_{vertex}| < 35$

$E_T^\gamma$  Track Magnitude



$\eta^\gamma$  Track Magnitude



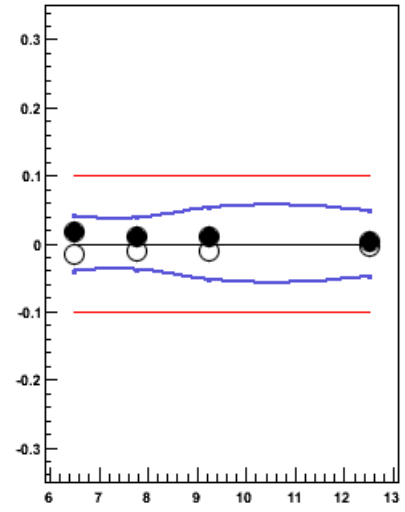
— Rel. statistical uncertainties

— 10% line

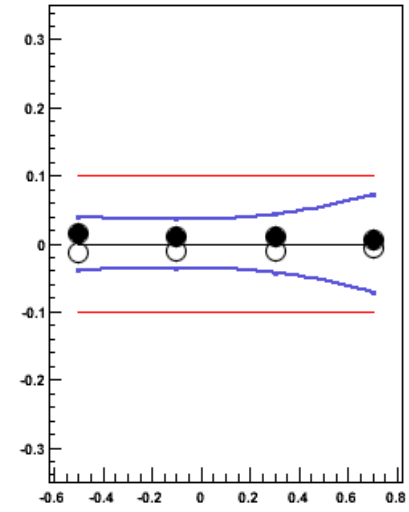
○  $p_{track} > 350$  MeV

●  $p_{track} > 150$  MeV

$E_T^\gamma$  Fragmentation



$\eta^\gamma$  Fragmentation



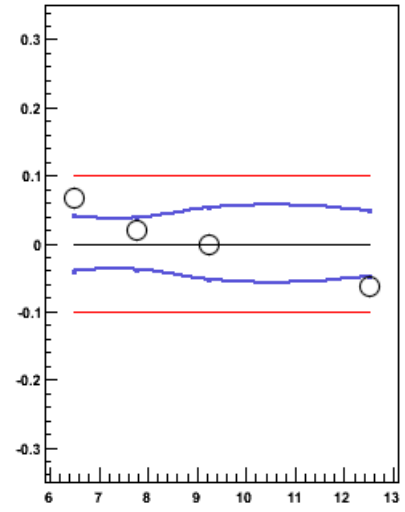
— *Rel. statistical uncertainties*

— *10% line*

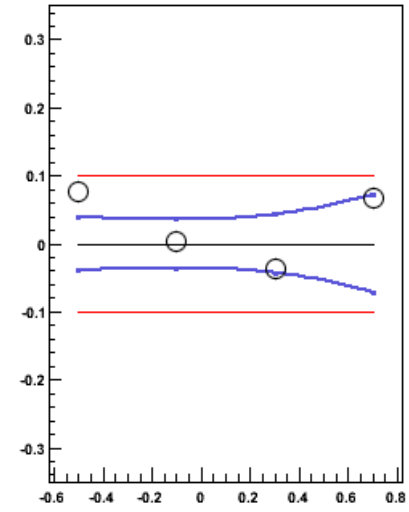
○ *+5% Fragmentation*

● *-5% Fragmentation*

$E_T^{\gamma}$  HERWIG



$\eta^{\gamma}$  HERWIG

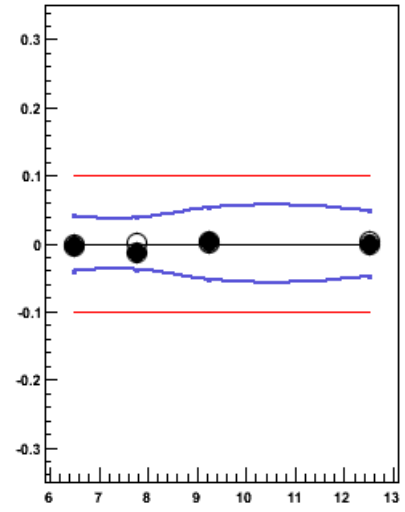


— *Rel.statistical uncertainties*

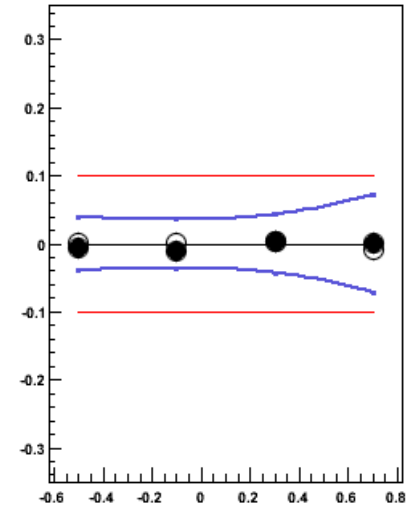
— *10% line*

○ *HERWIG*

$E_1^+$  fraction EMC



$\eta^+$  fraction EMC



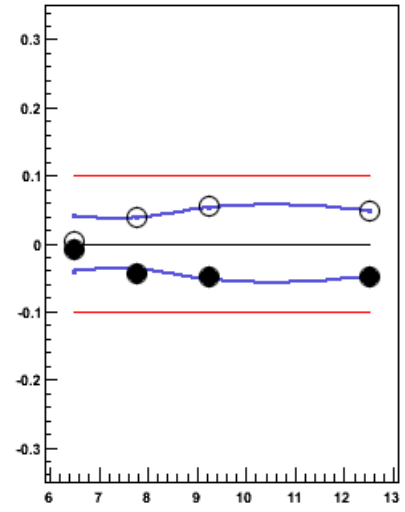
— Rel. statistical uncertainties

— 10% line

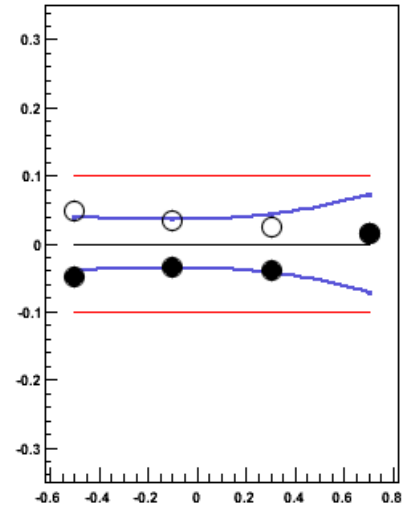
○ Fraction EMC +0.025

● Fraction EMC -0.025

$E_T^{\gamma} E_{\gamma, \text{Jet}}$



$\eta^{\gamma} E_{\gamma, \text{Jet}}$

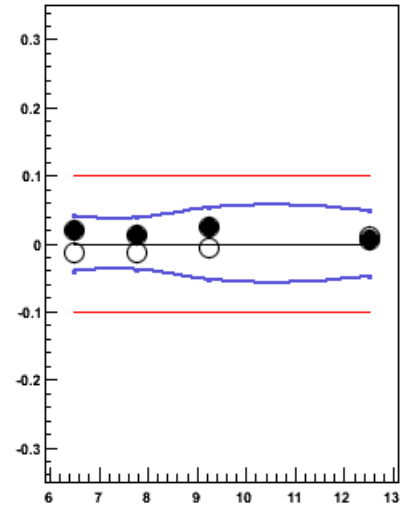
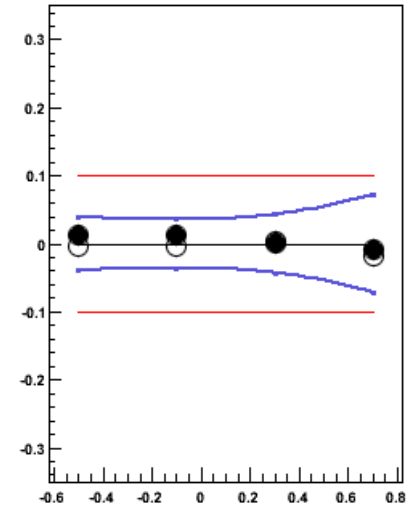


— *Rel. statistical uncertainties*

— *10% line*

○ *up. limit*

● *low. limit*

$E_T^\gamma \delta Z$  $\eta^\gamma \delta Z$ 

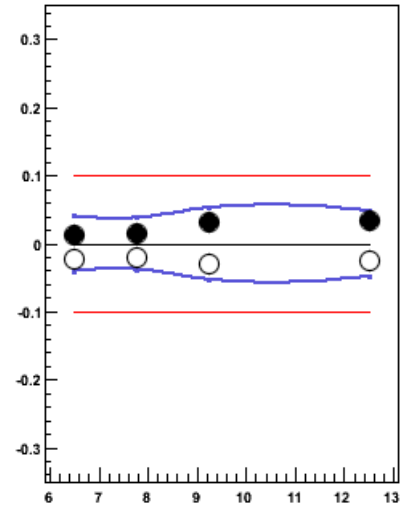
— *Rel. statistical uncertainties  $\delta Z$*

— *10% line*

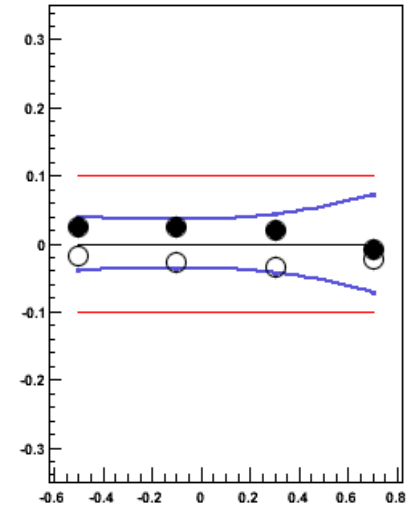
○  *$\delta Z$  fit range 1.0*

●  *$\delta Z$  fit range 0.6*

$E_T^\gamma \delta R$



$\eta^\gamma \delta R$



— *Rel. statistical uncertainties*

— *10% line*

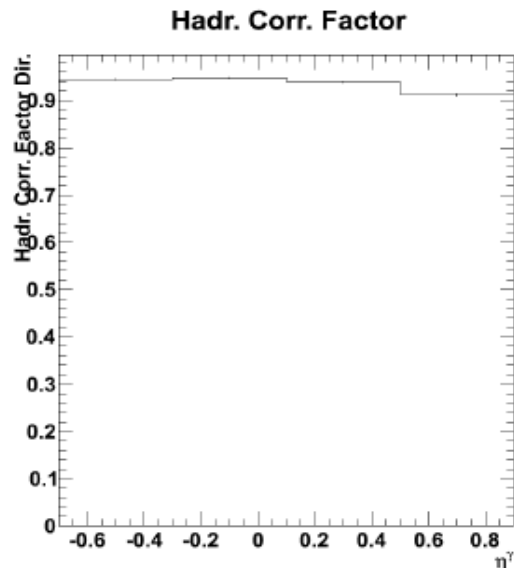
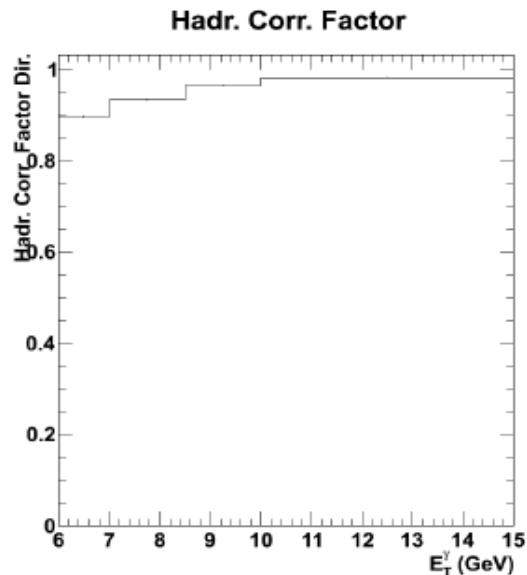
○  $\delta R 0.3$

●  $\delta R 0.1$

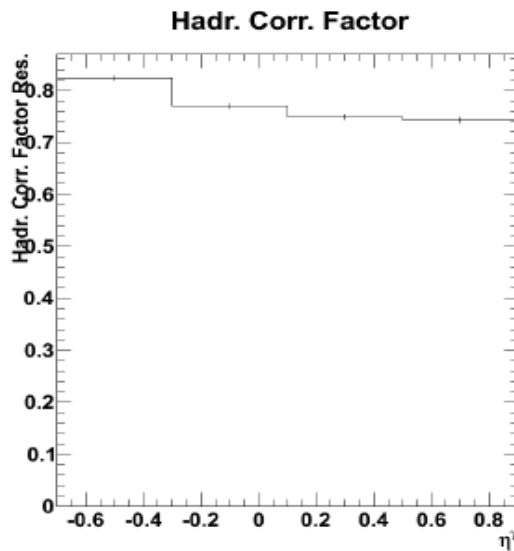
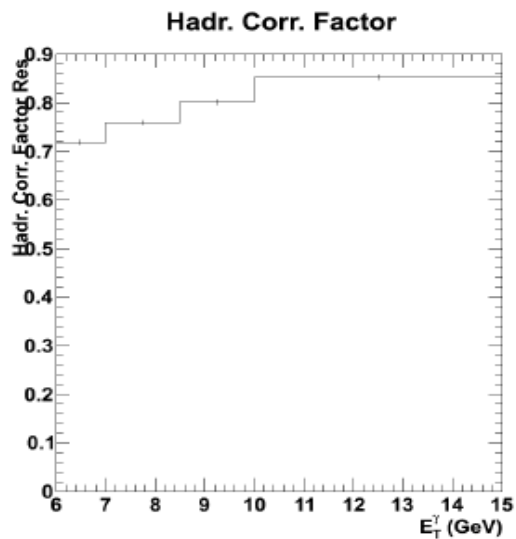
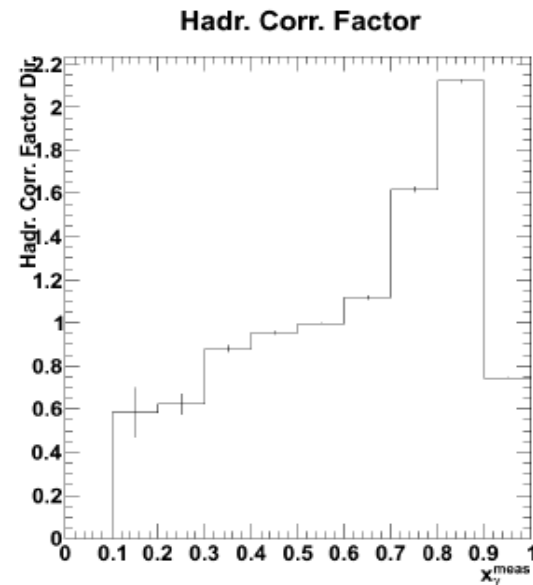


# **Hadronisation correction factors**

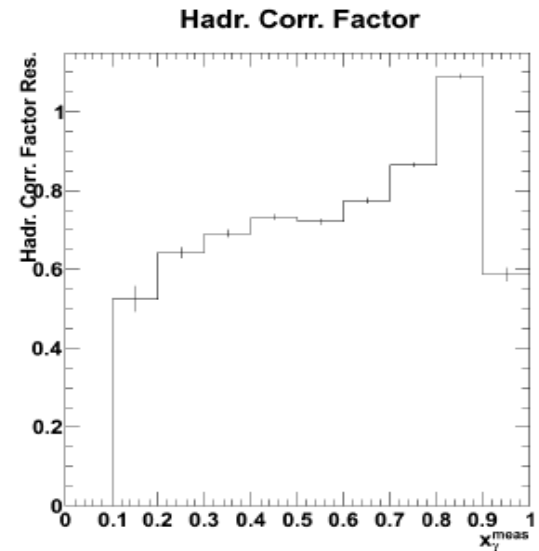
# PYTHIA Correction factors



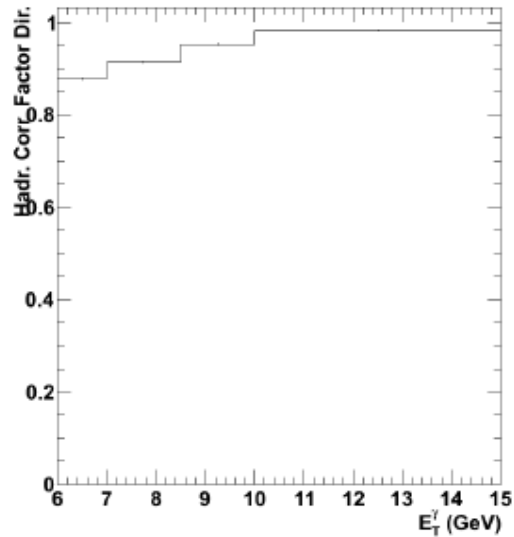
**Direct photon + jet**



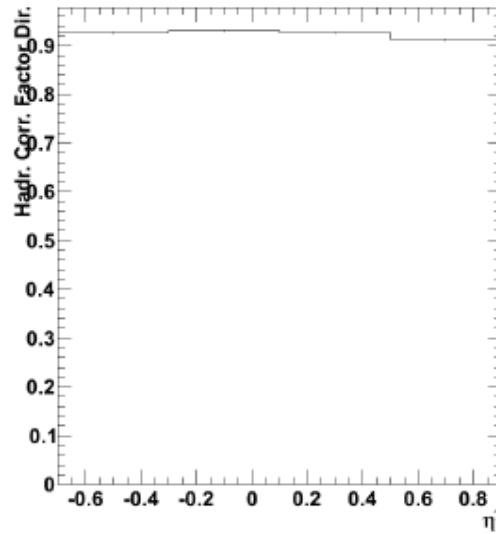
**Resolved photon + jet**



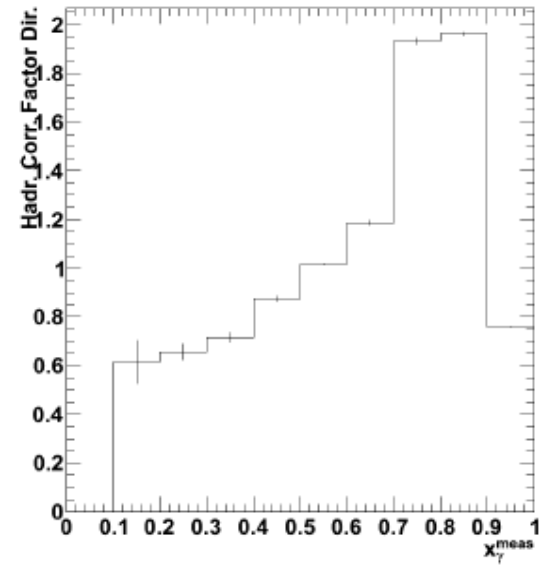
### Hadr. Corr. Factor



### Hadr. Corr. Factor

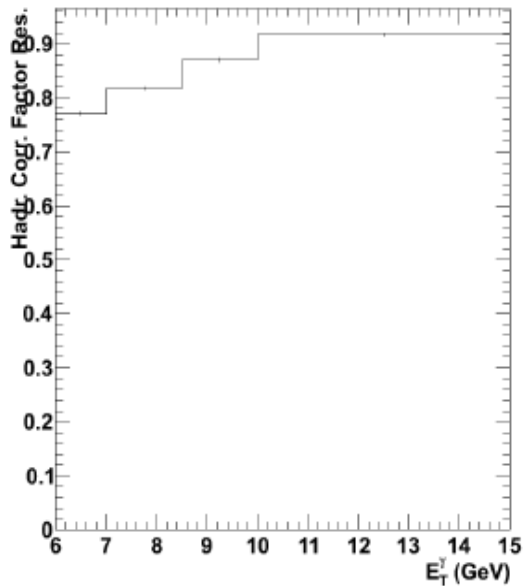


### Hadr. Corr. Factor

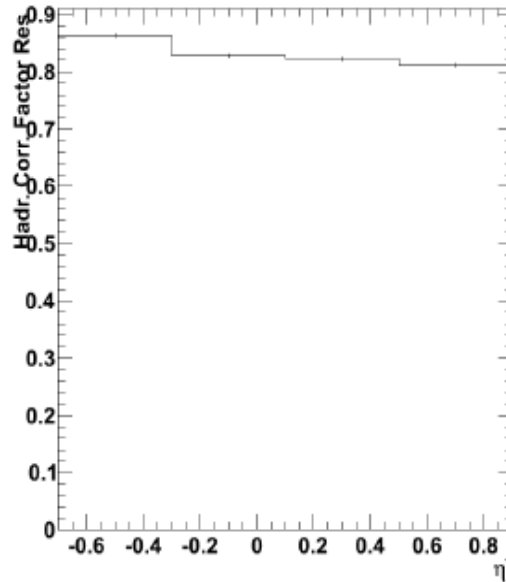


## Direct photon + jet

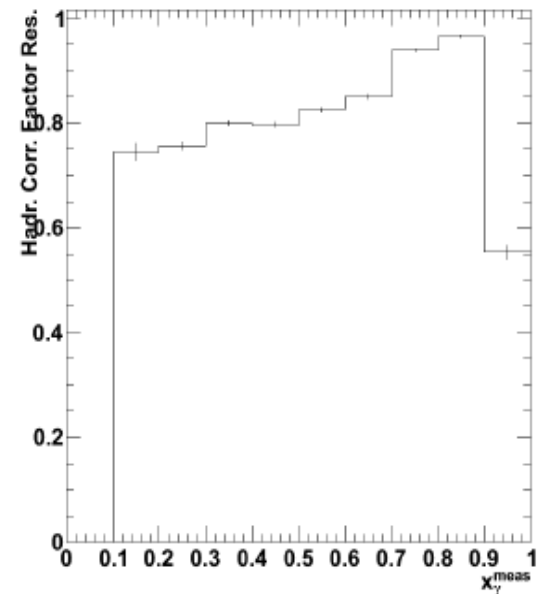
### Hadr. Corr. Factor



### Hadr. Corr. Factor



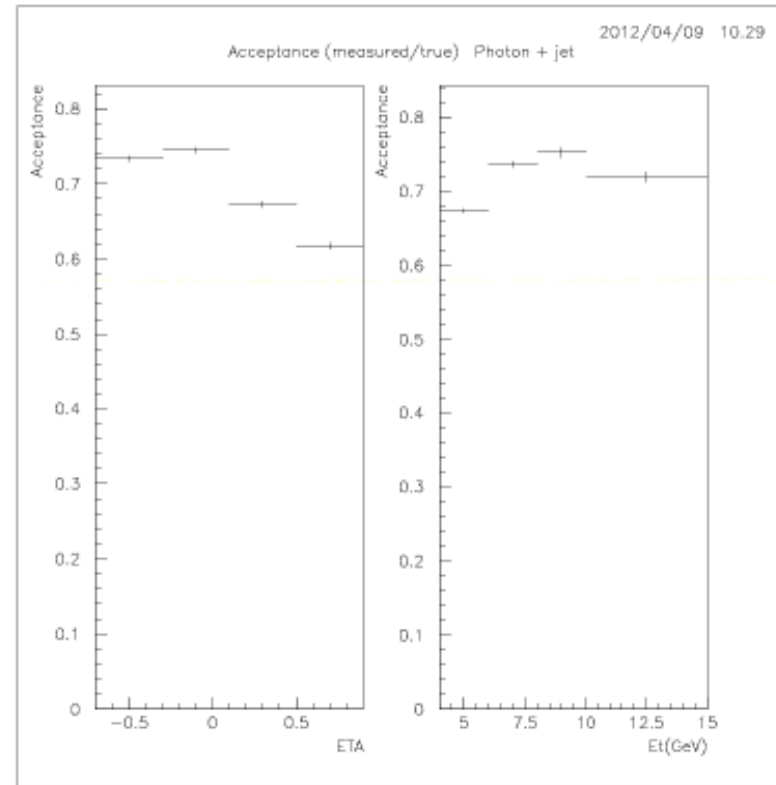
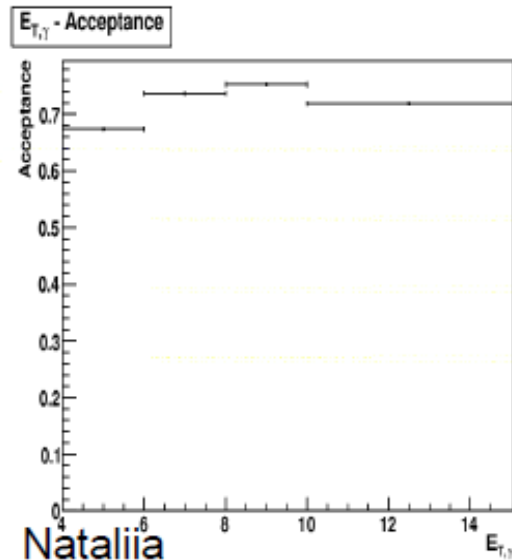
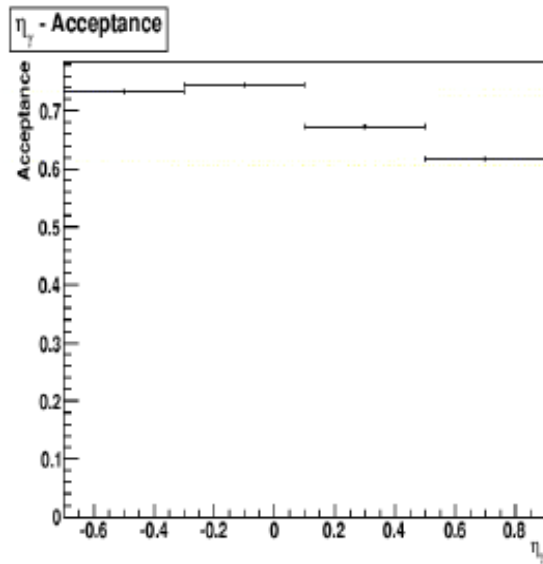
### Hadr. Corr. Factor



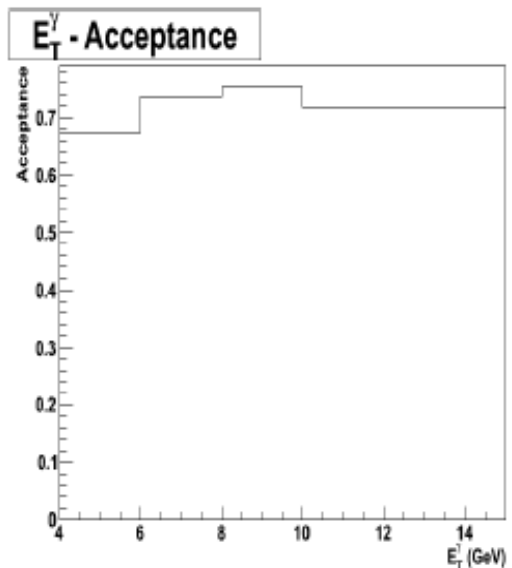
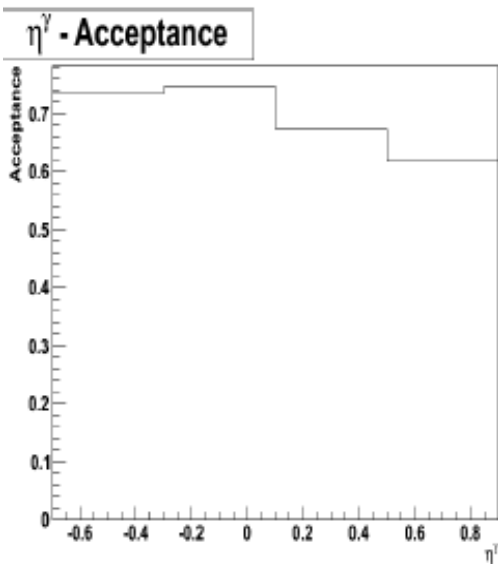
## Resolved photon + jet

# Acceptance

# MC signal direct acceptance (photon + jet)



Ian

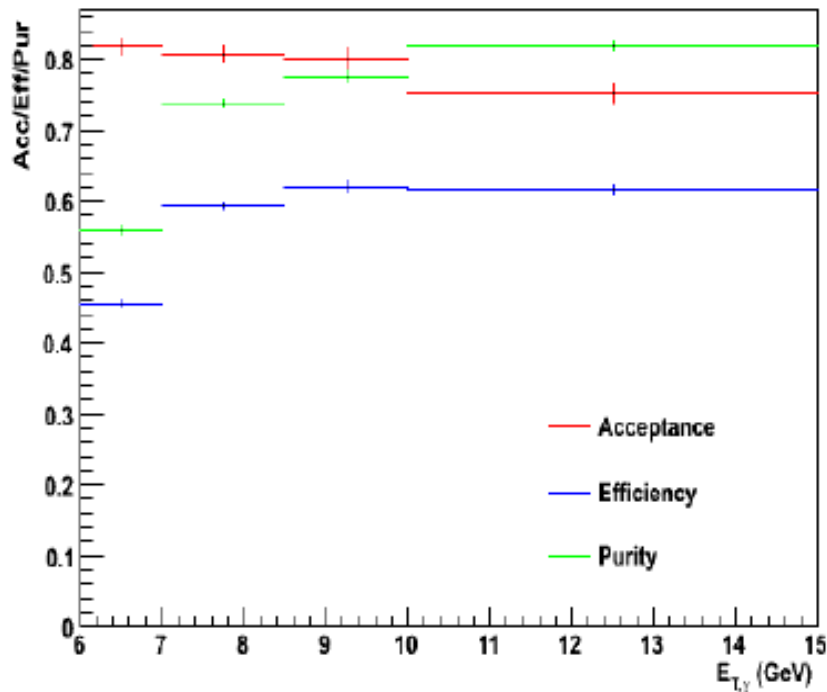


Andrii

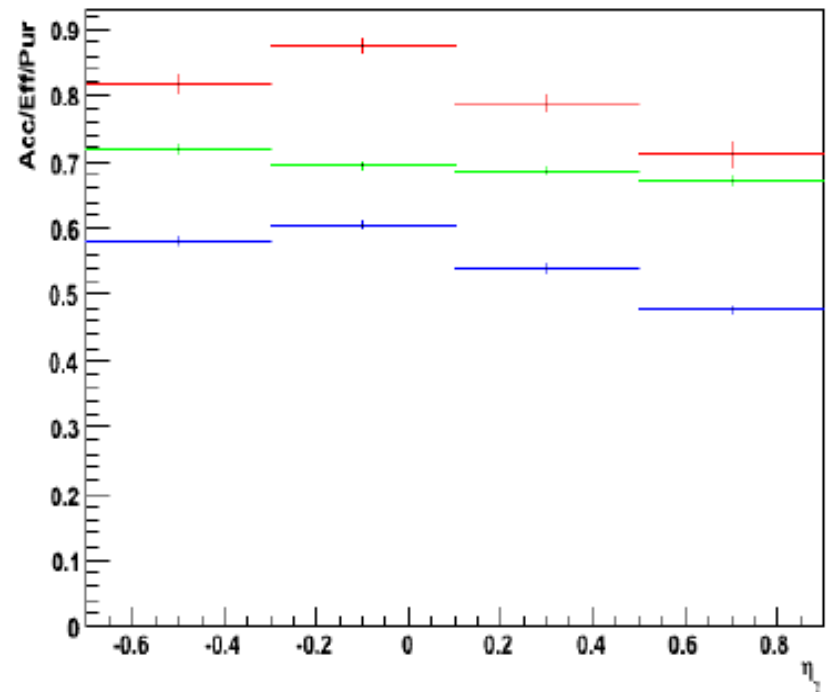
Truth level difference has no significant effect

# Inclusive Acceptance, Efficiency, Purity

Inclusive  $E_{T,\gamma}$  Acceptance, Efficiency, Purity

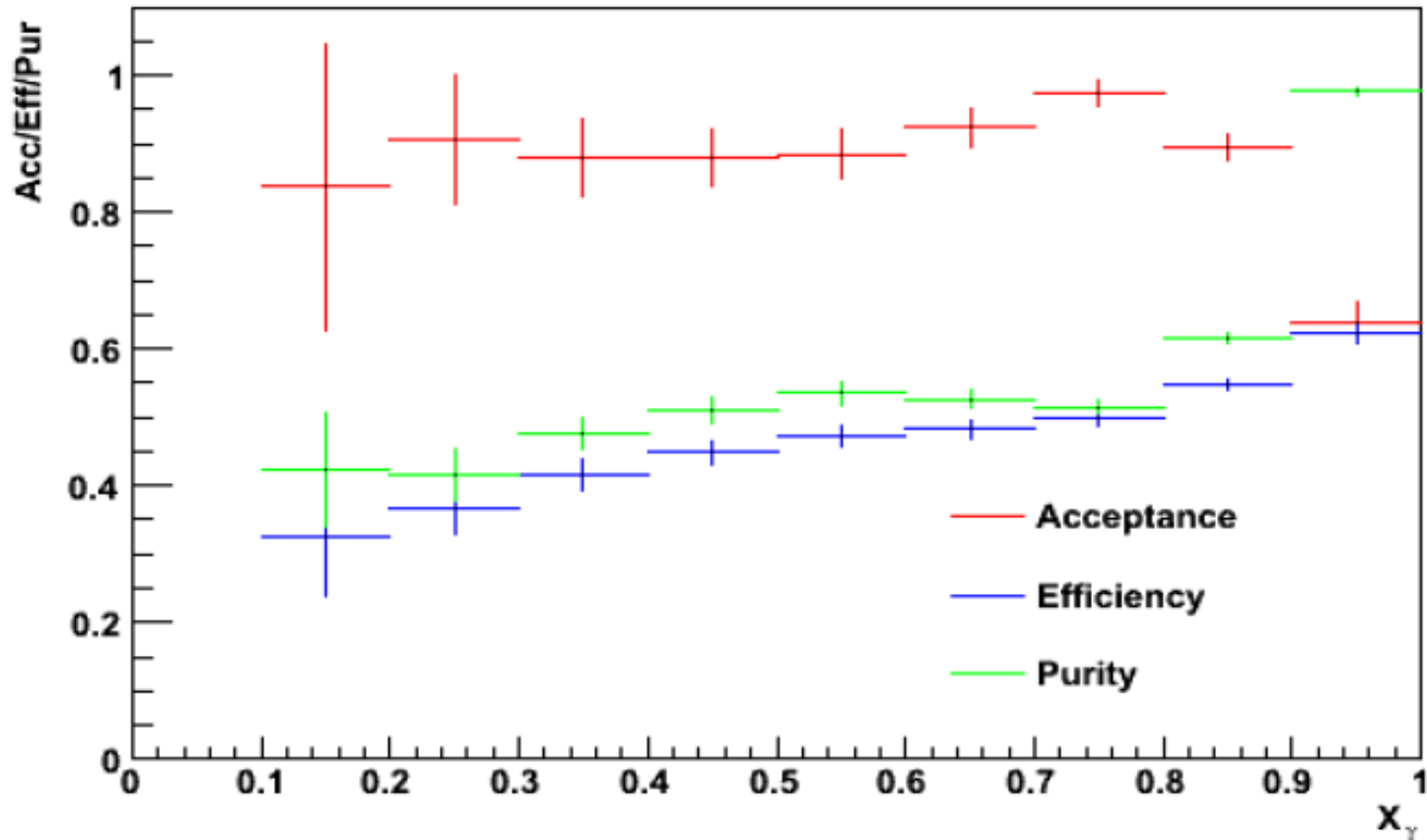


Inclusive  $\eta_\gamma$  Acceptance, Efficiency, Purity



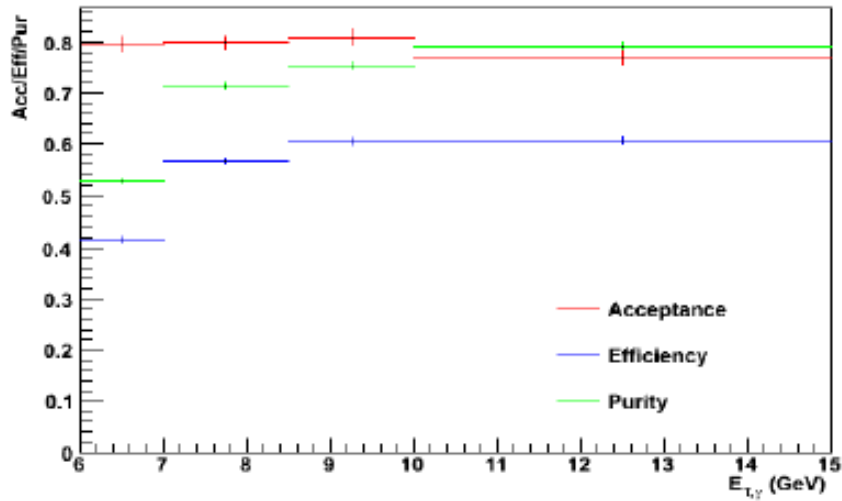
# Exclusive Acceptance, Efficiency, Purity

$X_Y$  Acceptance, Efficiency, Purity

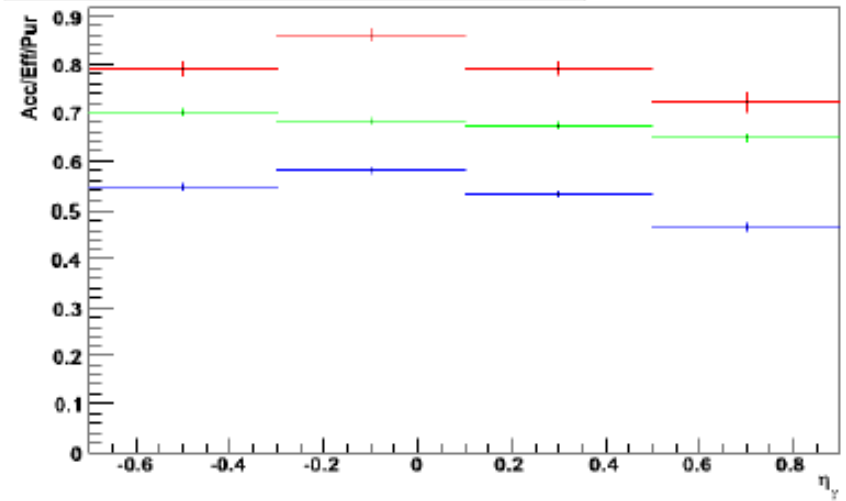


# Exclusive Acceptance, Efficiency, Purity

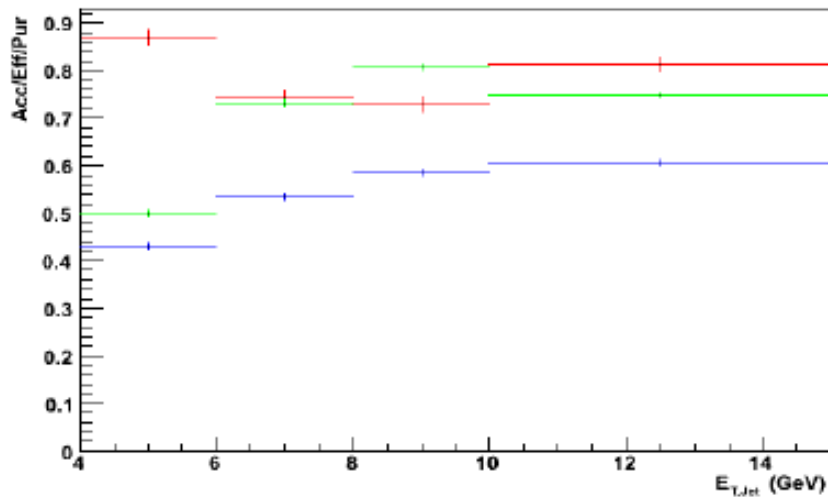
Exclusive  $E_{T,\gamma}$  Acceptance, Efficiency, Purity



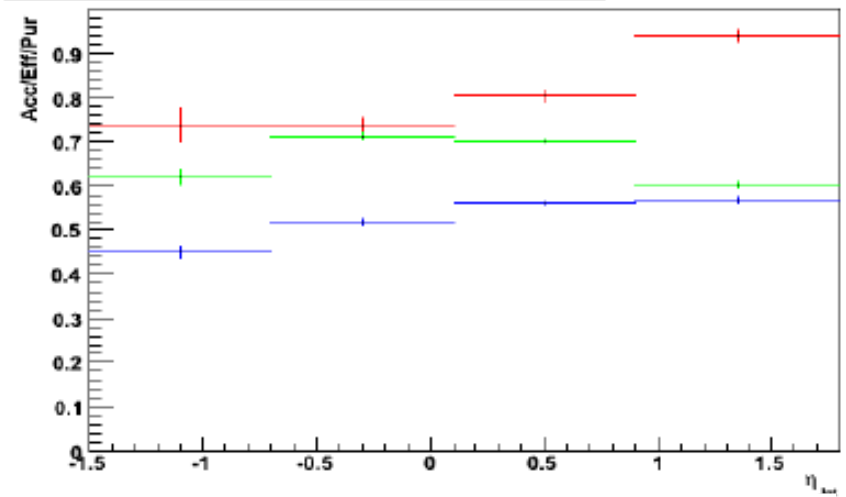
Exclusive  $\eta_\gamma$  Acceptance, Efficiency, Purity



Exclusive  $E_{T,\text{jet}}$  Acceptance, Efficiency, Purity



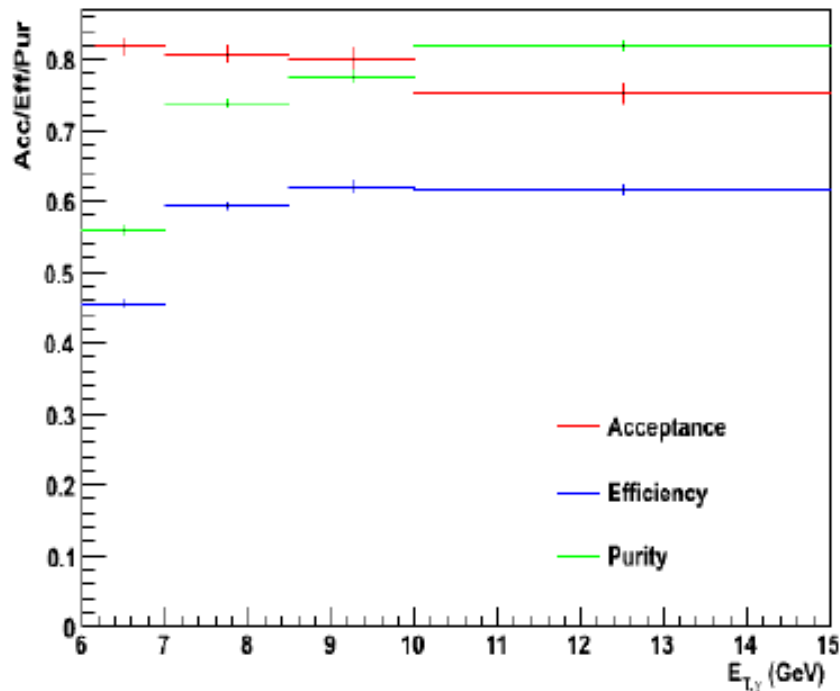
Exclusive  $\eta_{\text{jet}}$  Acceptance, Efficiency, Purity



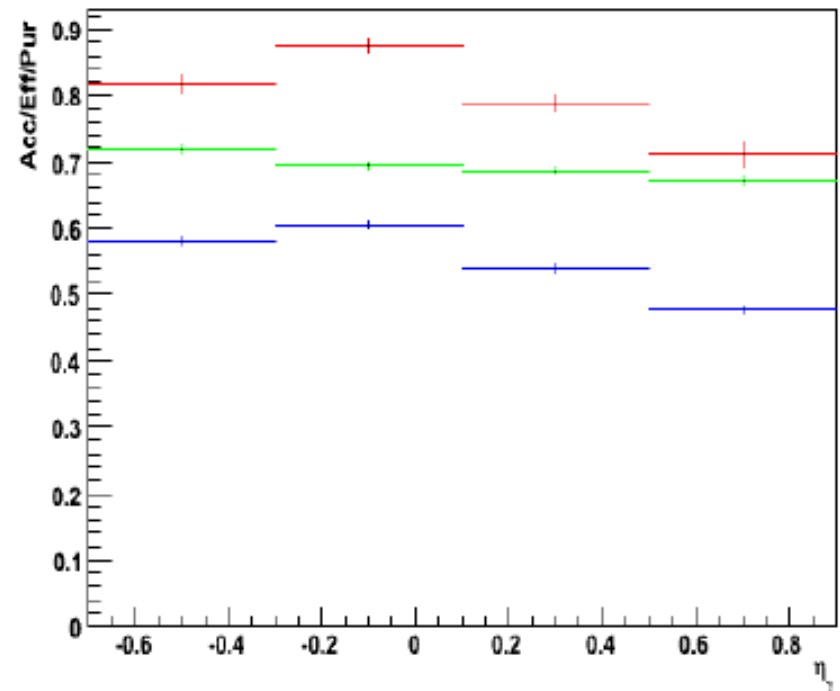


# Inclusive Acceptance, Efficiency, Purity

Inclusive  $E_{T,\gamma}$  Acceptance, Efficiency, Purity

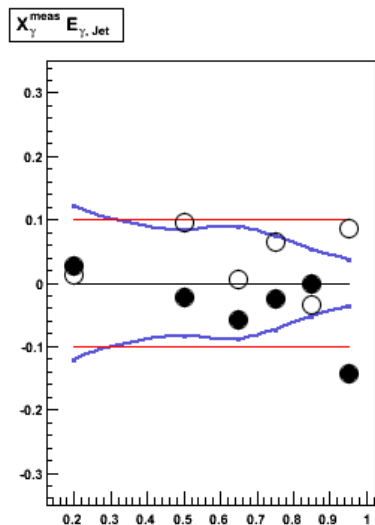
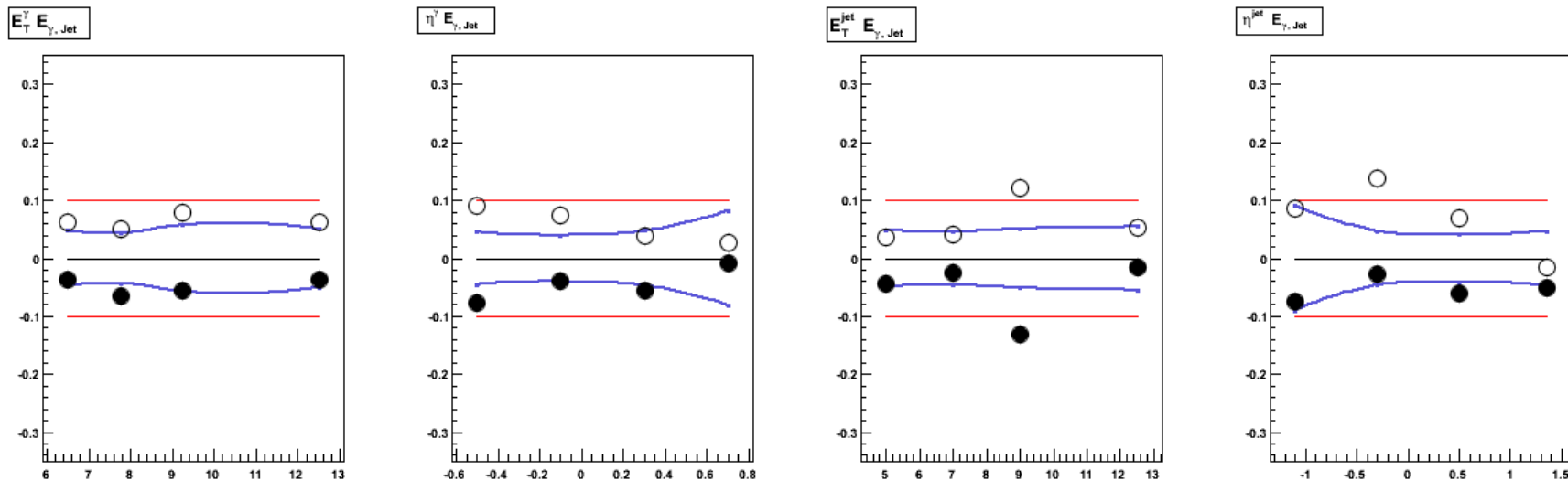


Inclusive  $\eta_\gamma$  Acceptance, Efficiency, Purity



# **E scale systematic**

# Energy scale systematics



Change photon E by 2% and

if JetEt > 10 GeV - jet E by 1.5 %.

if JetEt < 10 GeV - jet E by 2 %.

In addition:

JetEt < 6 GeV - jet E by  $\sqrt{4*4-2*2}$  %.

$6 < \text{JetEt} < 10$  GeV - jet E by  $\sqrt{2.5*2.5-2*2}$  %.

— Rel.statistical uncertainties

— 10% line

○ up. limit

● low. limit

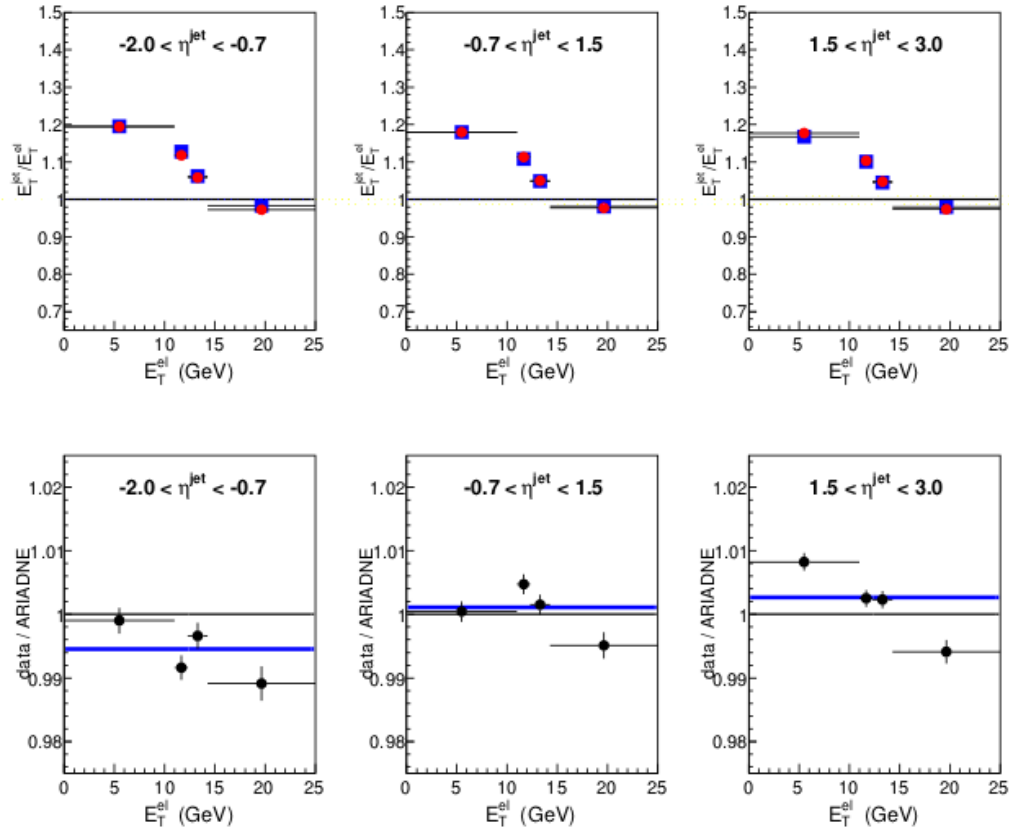
These systematics are to be symmetrised when calculating total systematics

# Estimation of jet energy uncertainty: jets with

$$E_T^{jet} > 10 \text{ GeV}$$

- 04-05e, 06e, 06-07p
  - **inclusive DIS events**
  - $E_T^{e'} > 10 \text{ GeV}$
- We consider:
- strictly one jet with  $E_T^{jet} > 10 \text{ GeV}$ ;
  - $-1.5 < \eta_{jet} < 1.8$  (our usual phase space);
  - no further jets in  $2.5 < E_T^{jet} < 10 \text{ GeV}$  range.
  - Build the ratio of  $E_T^{jet}$  to  $E_T^{e'}$  in bins of  $E_T^{e'}$  (4 bins) and  $\eta^{jet}$  (3 bins) for data and Ariadne.

# Jets with $E_T^{jet} > 10$ GeV. $E_T^{jet} / E_T^{e'}$ ratio.



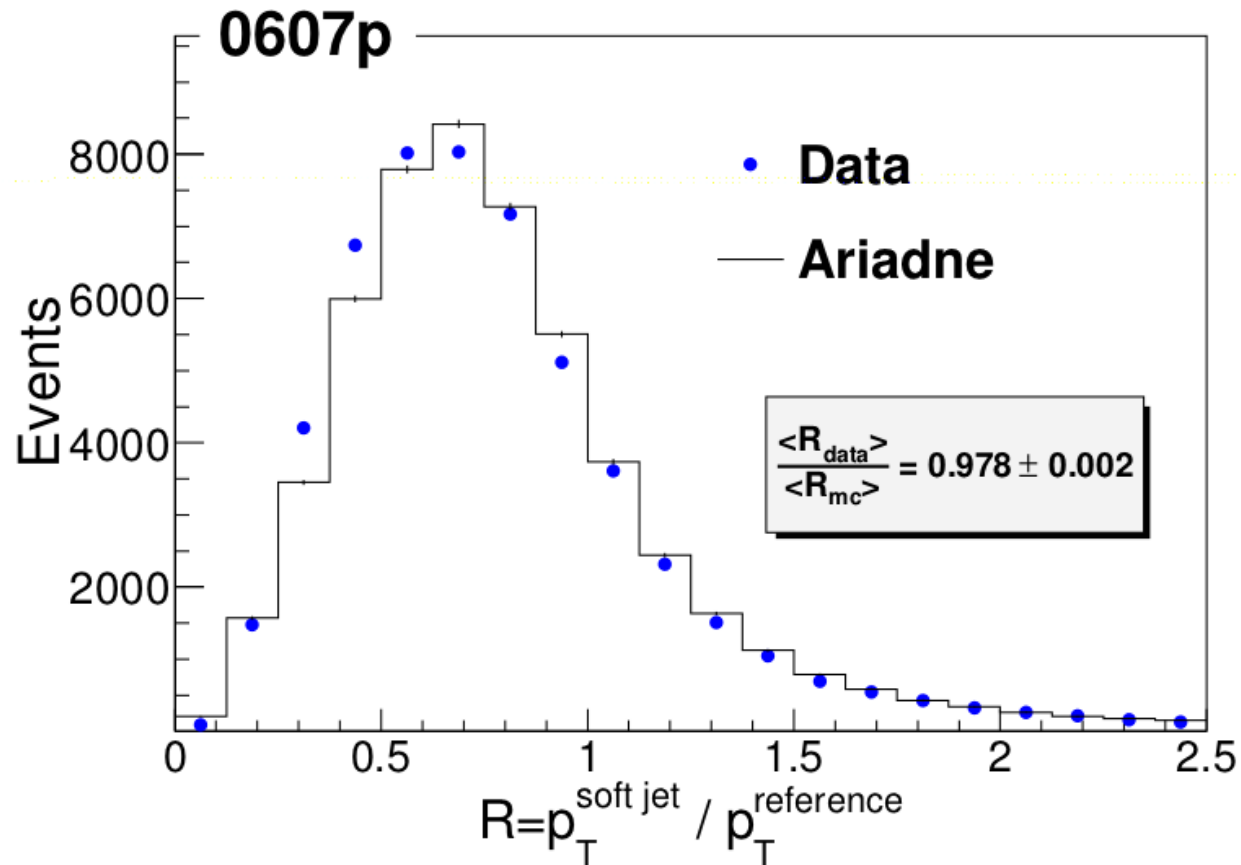
**Circles** are data, **squares** are ARIADNE

Difference between ratios in data and Monte Carlo is smaller than  $\pm 1.5\%$

# Estimation of jet energy uncertainty: jets with $2.5 < E_T < 10$ GeV

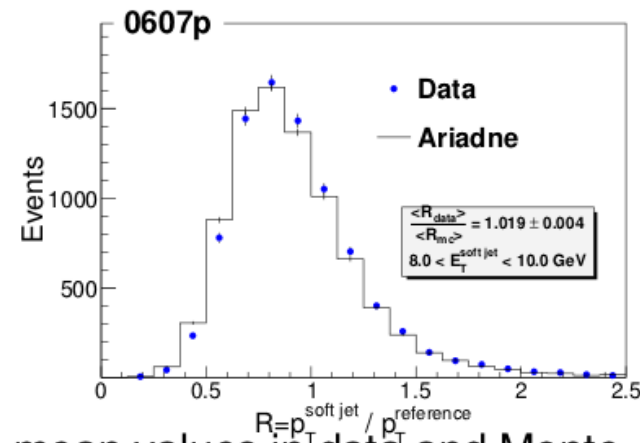
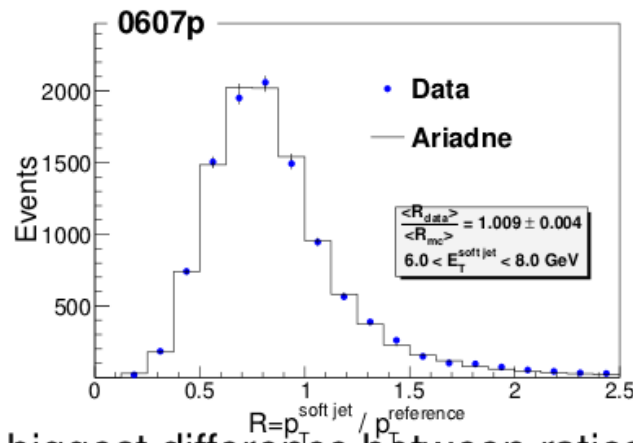
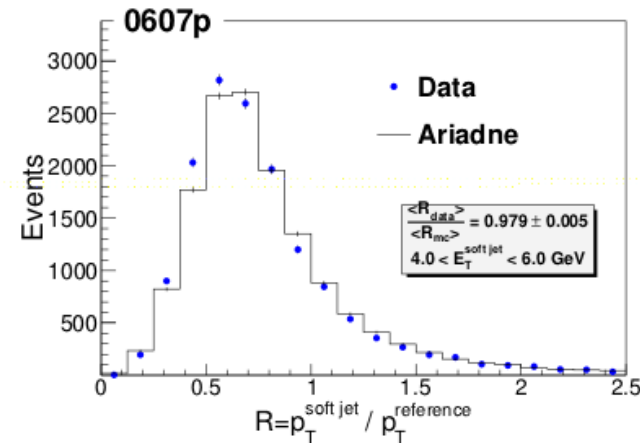
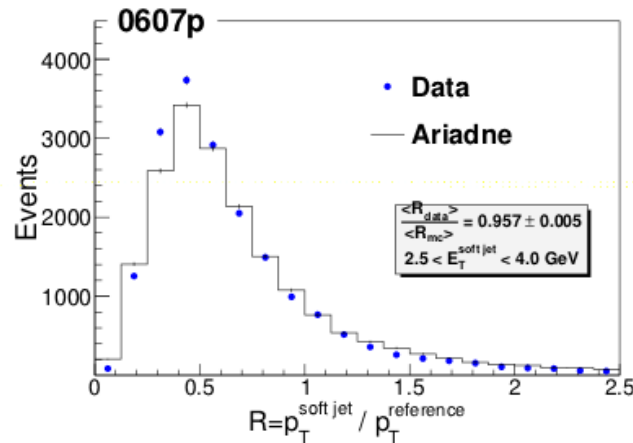
- 04-05e, 06e, 06-07p
- inclusive DIS events
- We consider:
  - strictly one jet with  $E_T^{jet} > 10$  GeV;
  - and strictly one jet with  $2.5 < E_T < 10$  GeV;
  - $-1.5 < \eta_{jet} < 1.8$  (our usual phase space).
  - Combine the 4-vectors of the high-energetic jet and of the scattered electron into reference 4-vector;
  - look into the ratio  $\frac{E_T^{softjet}}{E_T^{reference}}$  for data and Ariadne.

Jets with  $2.5 < E_T < 10$  GeV. Ratio  $\frac{E_T^{\text{softjet}}}{E_T^{\text{reference}}}$ .



Agreement within 2%

# Jets with $2.5 < E_T < 10$ GeV. Ratio $\frac{E_T^{\text{softjet}}}{E_T^{\text{reference}}}$ .



The biggest difference between ratios of the mean values in data and Monte Carlo is about 4% (for lowest  $E_T^{\text{jet}}$ ).

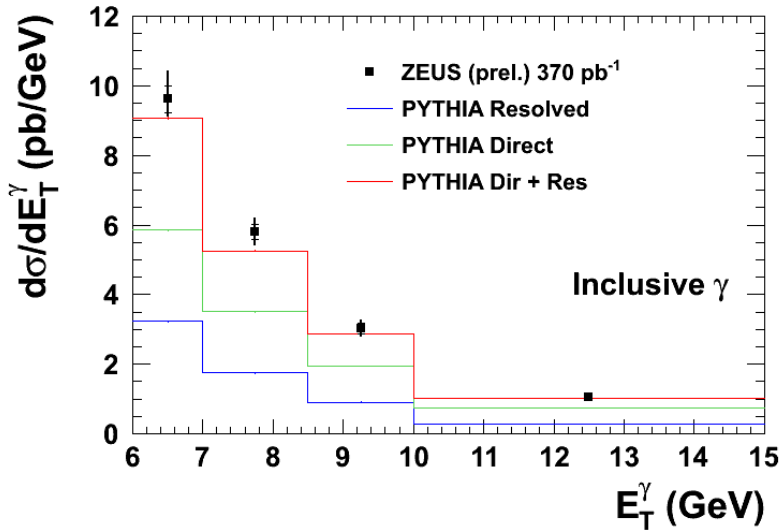
We conclude that the jets are well understood at this level.



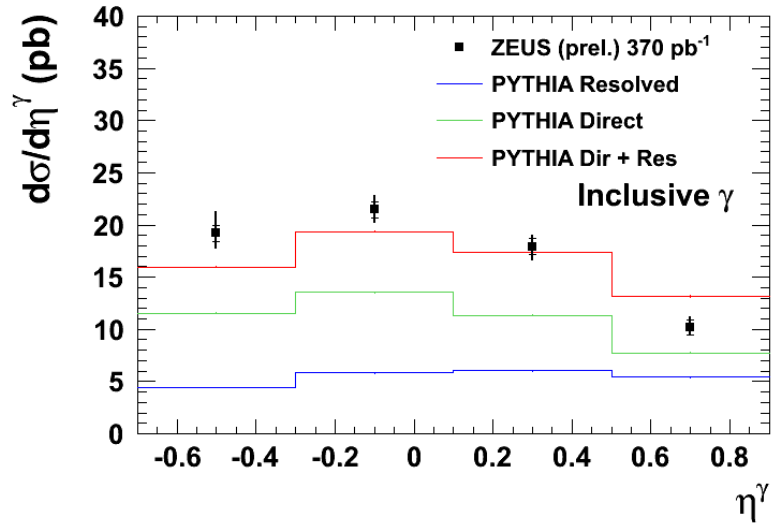
# **Post-fitting comparison with the PYTHIA model**

# Post-fitting comparison with the PYTHIA model

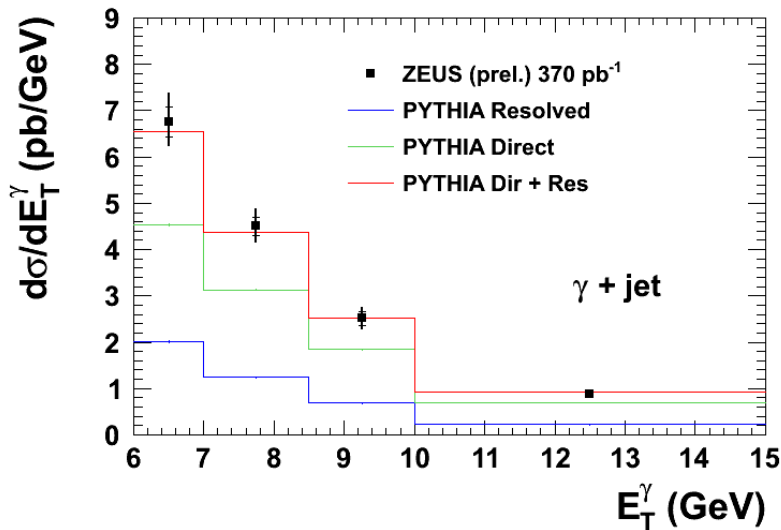
## ZEUS



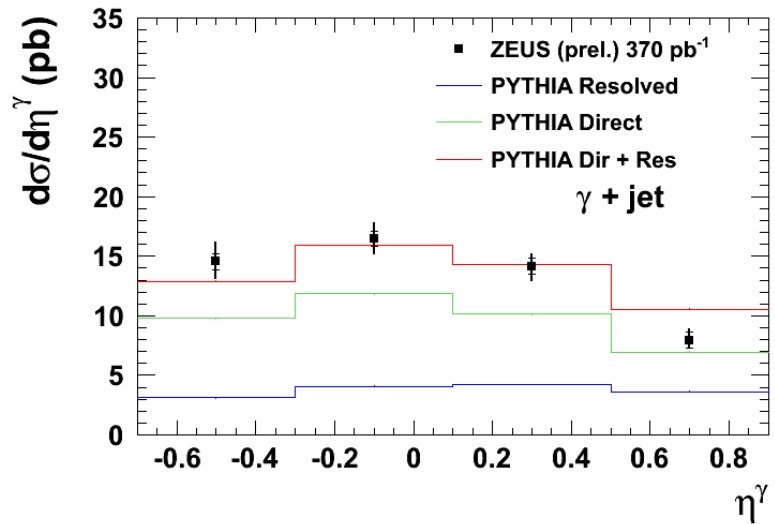
## ZEUS



## ZEUS



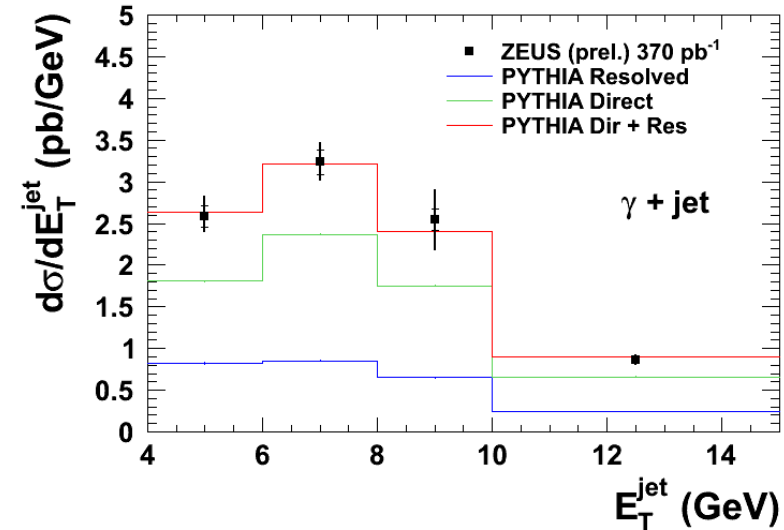
## ZEUS



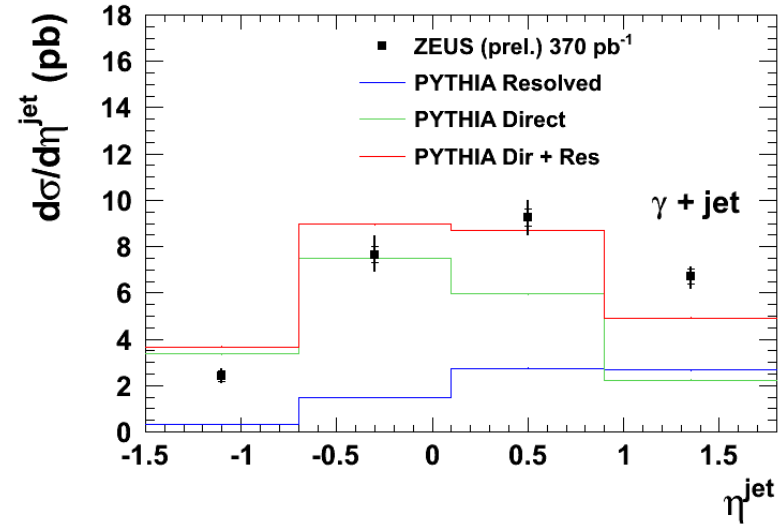
Reasonable description of data by MC model

# Post-fitting comparison with the PYTHIA model

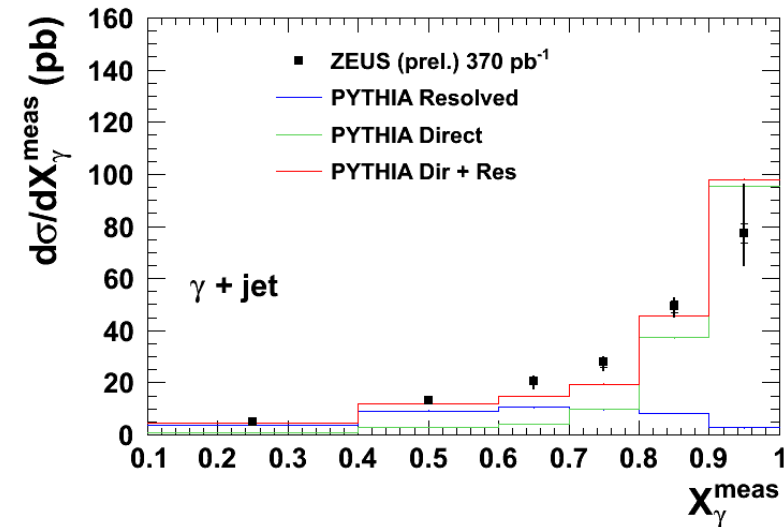
## ZEUS



## ZEUS



## ZEUS

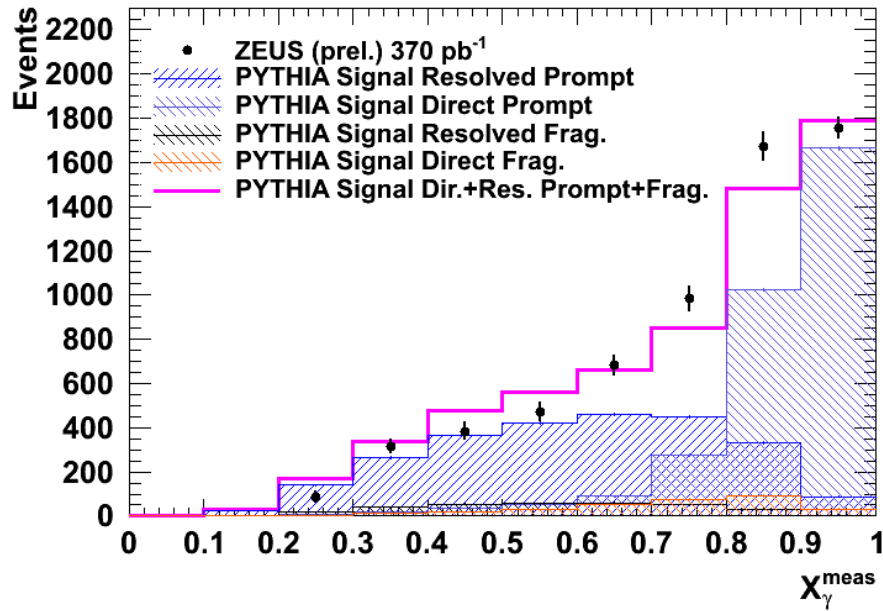


Reasonable description of data by MC model

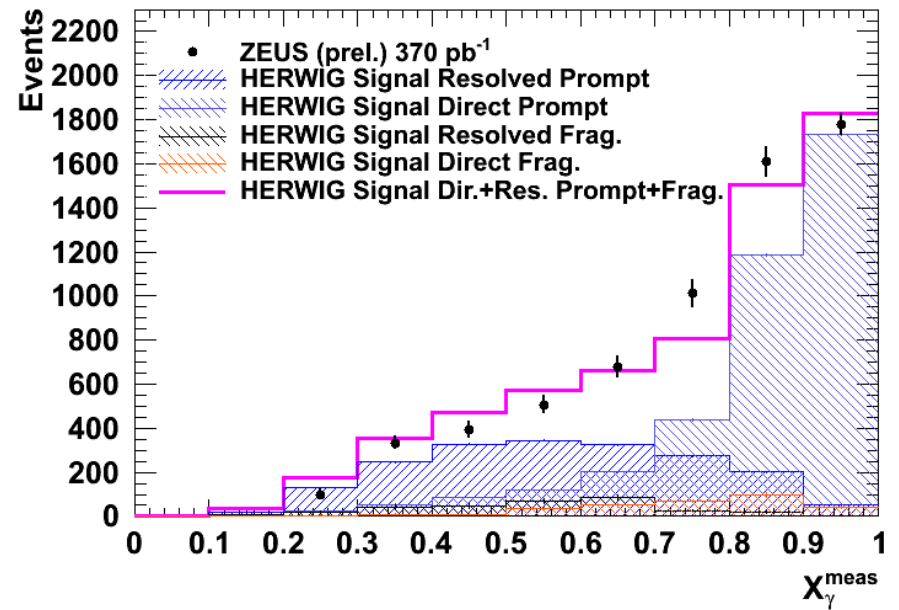
# **Additional plots**

# PYTHIA and HERWIG x-gamma

ZEUS

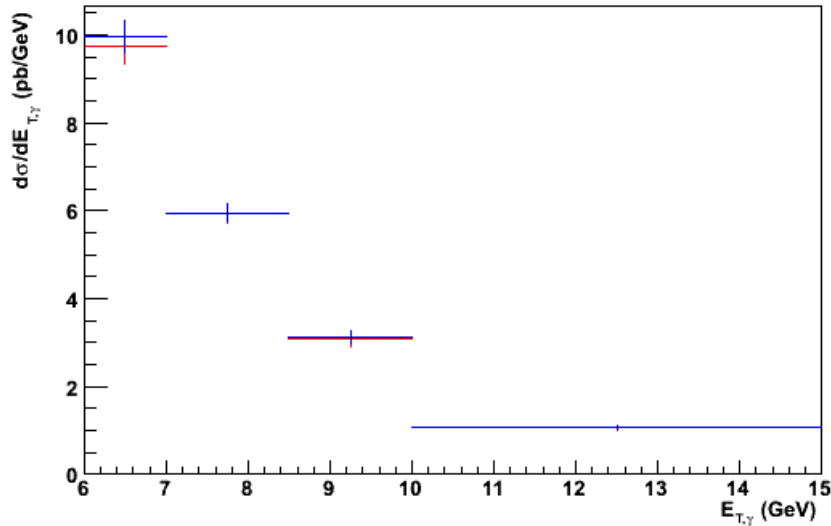


ZEUS

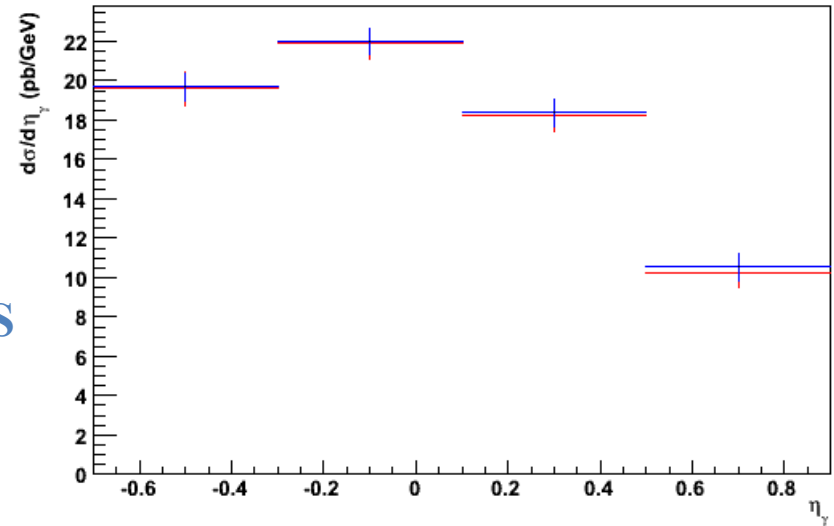


# Comparison between analyses. Inclusive photon

Inclusive  $E_{T,\gamma}$  cross-section



Inclusive  $\eta_\gamma$  cross-section

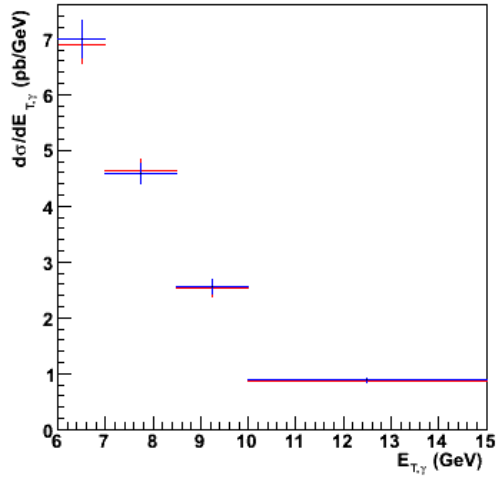


AI  
IOS

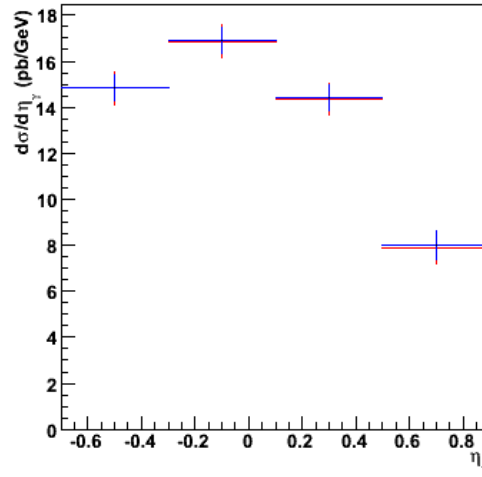
Shown are cross-section numbers with statistical uncertainties only.  
Plots provided by Volodymyr. Comparison with his results is ongoing.

# Comparison between analyses. Inclusive photon

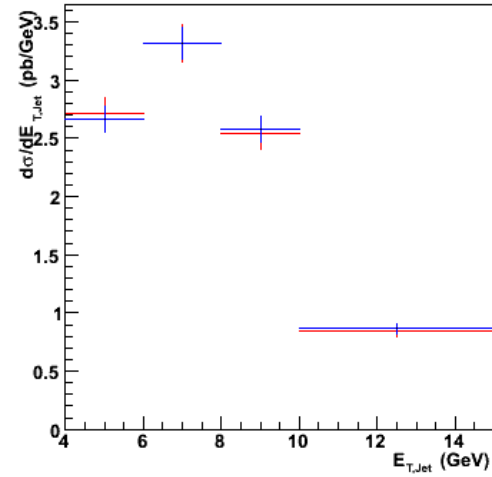
Exclusive  $E_{T,\gamma}$  cross-section



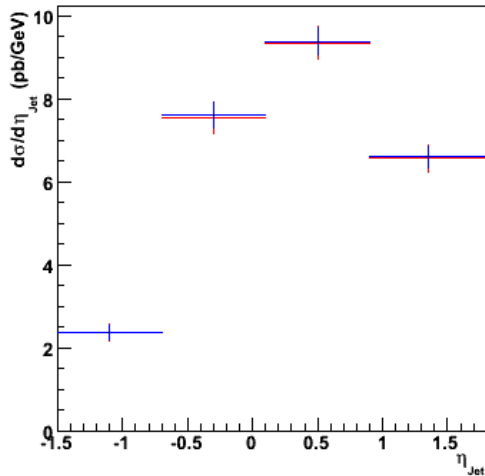
Exclusive  $\eta_\gamma$  cross-section



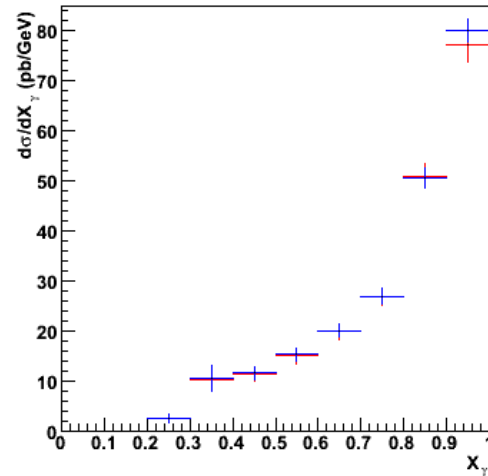
Exclusive  $E_{T,Jet}$  cross-section



Exclusive  $\eta_{Jet}$  cross-section



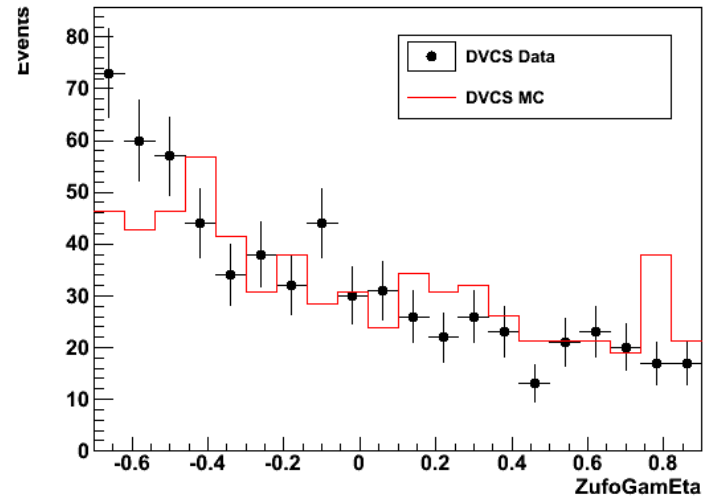
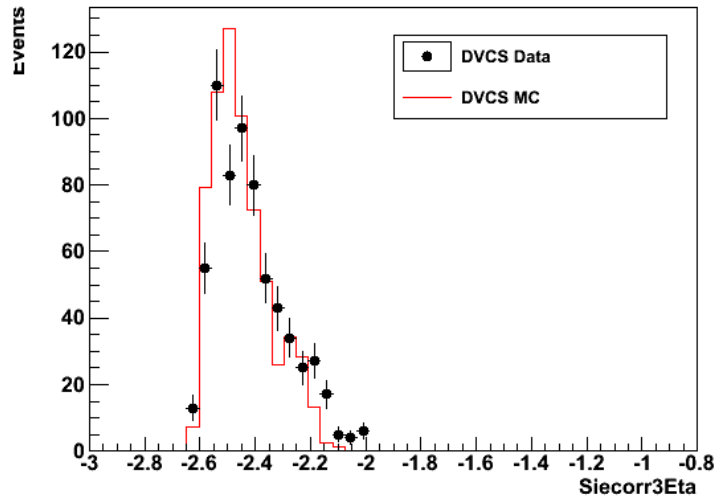
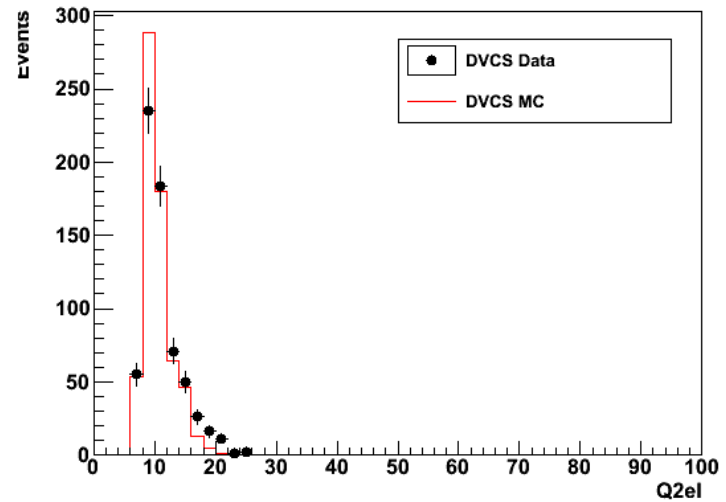
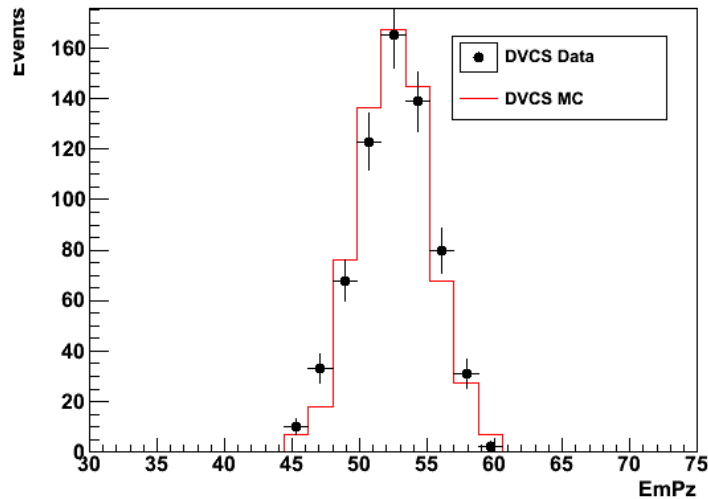
$X_\gamma$  cross-section



Excellent agreement between analyses considering that IOS is based on a 50-50 mix and no change in cone factor.

Last X-gamma bin might be caused by different implementation of selection on true level, hence different acceptance correction factor.

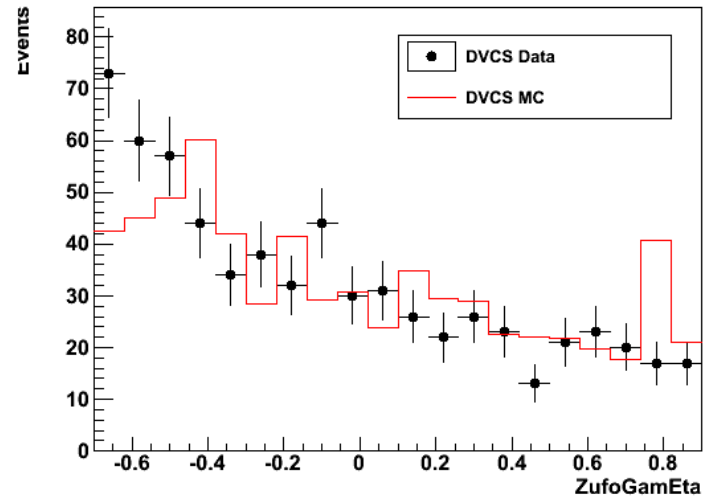
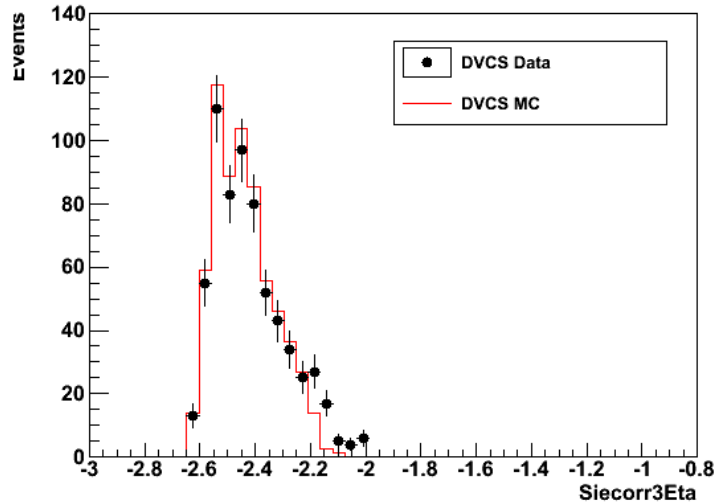
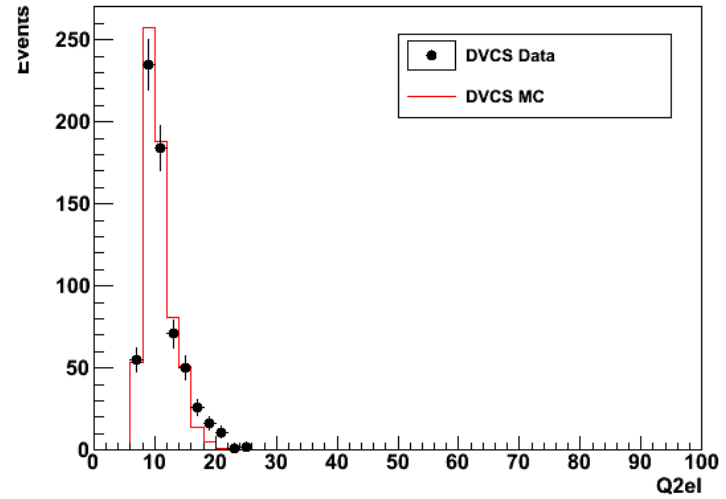
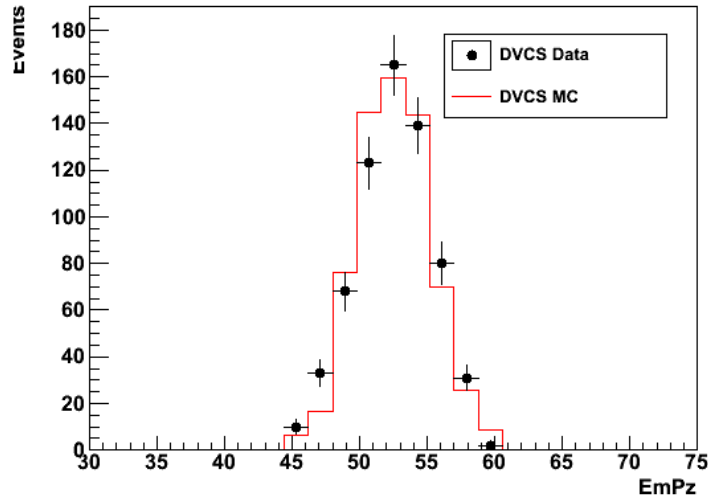
# DVCS . Control plots without reweighting



Not so good description of data by MC .

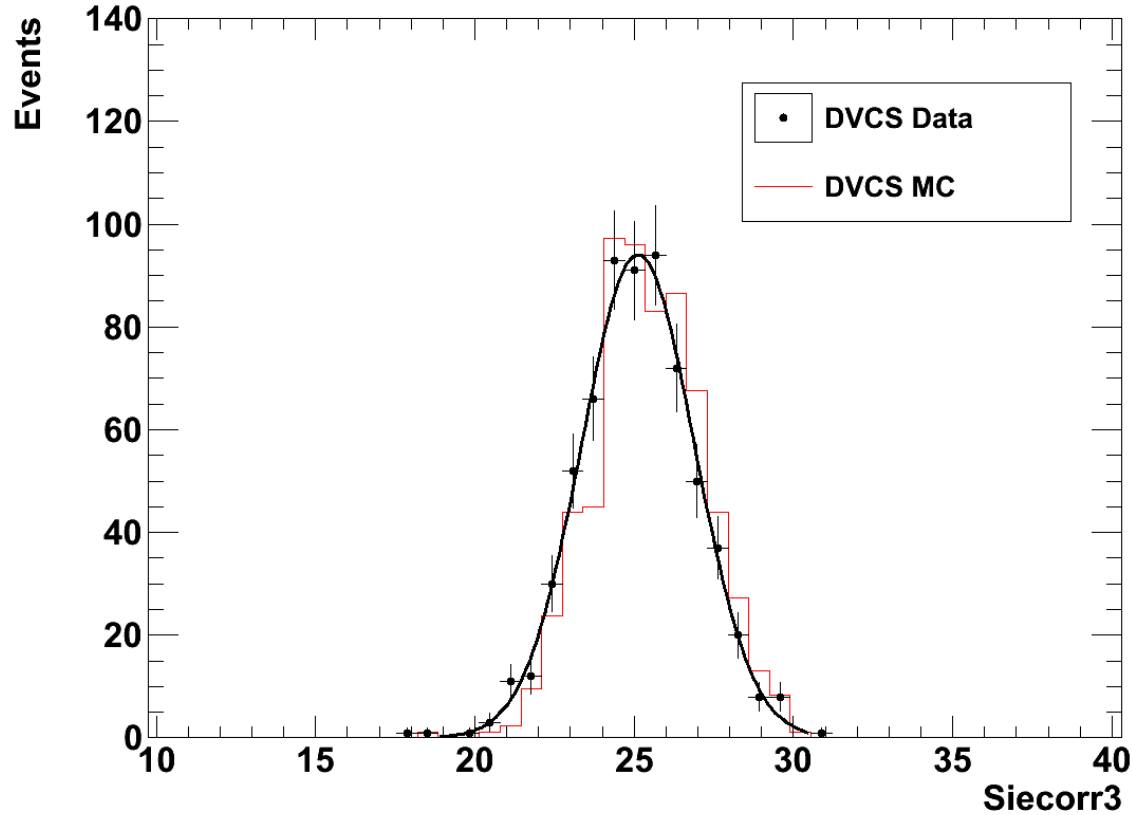


# DVCS . Control plots after reweighting



Small improvement in  $Q_2$  description can be noticed

# Fit of S iecorr3



S iecorr3 Data 18.25-31.15

Fit 2.51213e+01

Agreement within 2%